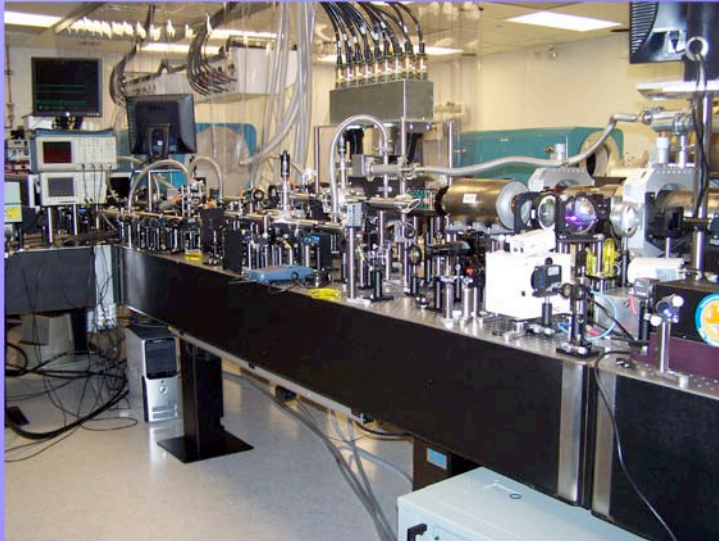


# Petawatt to Exawatt Lasers

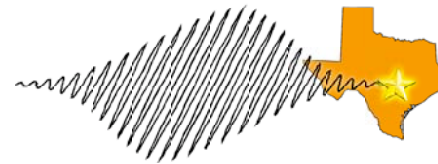
*The technology and applications of the most powerful lasers ever built*



Presented by:

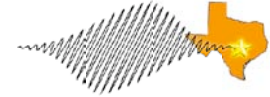
***Todd Ditmire***

**Center for High Energy Density Science  
Department of Physics  
University of Texas at Austin**



*Or... Ultrafast, ultrapowerful lasers create  
ultraintense light to study the ultrahot*

# Most important: The people who do the work



## Center Faculty:

Roger Bengtson  
Todd Ditmire  
Mike Downer  
John Keto  
Michael Marder (CNLD)  
Eric Taleff (Mech Eng)

## Center Senior Scientists:

Aaron Bernstein  
Erhard Gaul  
Gilliss Dyer  
Hernan Quevedo  
Alan Wootton

## Center Sabbatical Visitor:

Jack Glassman

## Closely Collaborating Scientists:

Jens Osterholz (Dusseldorf)  
Aaron Edens (Sandia)  
Alex Arefiev (IFS)  
Stephan Bless (IAT)  
Boris Breizman (IFS)  
Charles Chiu  
Richard Fitzpatrick (IFS)  
Wendell Horton (IFS)  
Gennady Shvets (IFS)

## Graduate students:

Woosuk Bang  
Joel Blakeney  
Sam Feldman  
Ahmed Ehlal  
Nirmala Kandadai  
Intai Kim  
Donghoon Kuk  
Sean Lewis  
Matt McCormick  
Nathan Riley  
Kristina Serratto  
Nicholas Pedrazas  
Craig Wagner  
Matt Wisher

## Post Docs:

Kay Hoffman  
Rashida Jafer  
Robert Morgan  
Heiko Thomas

## Undergraduates:

Wilson Bui  
Jennifer Chae  
Lindsay Fuller  
Blake Griffith  
Andrew Lafferty  
Trevor Latson  
Emily Leos  
David Nicholson  
James Pruet  
Giovanni Rossi

## U. Texas Petawatt and Exawatt Team

Franki Aymond  
Woosuk Bang  
Aaron Bernstein  
Joel Blakeney  
Ted Borger  
Mike Donovan  
Gilliss Dyer  
Ramiro Escamilla  
Erhard Gaul  
Jack Glassman  
Doug Hammond  
Donghoon Kuk  
Srdjan Marijanovic  
Mikael Martinez  
Marty Ringuette  
Craig Wagner  
Dave Nicholson  
Lindsay Fuller  
David Bui  
Giovanni Rossi  
Andrew Lafferty  
Blake Griffith

## LLNL

Rick Cross  
Chris Ebbers  
Al Erlandson  
John Caird  
Brent Stuart  
Bill Molander  
Tim Weiland  
Ralph Page  
Jerry Britten

## Schott Glass N. America

Matt Roth  
Eric Urruti  
Nathan Carlie

## Logos Technologies

Mike Campbell  
Jason Zweiback  
Dave Eimerl (Eimex)  
Bill Krupke

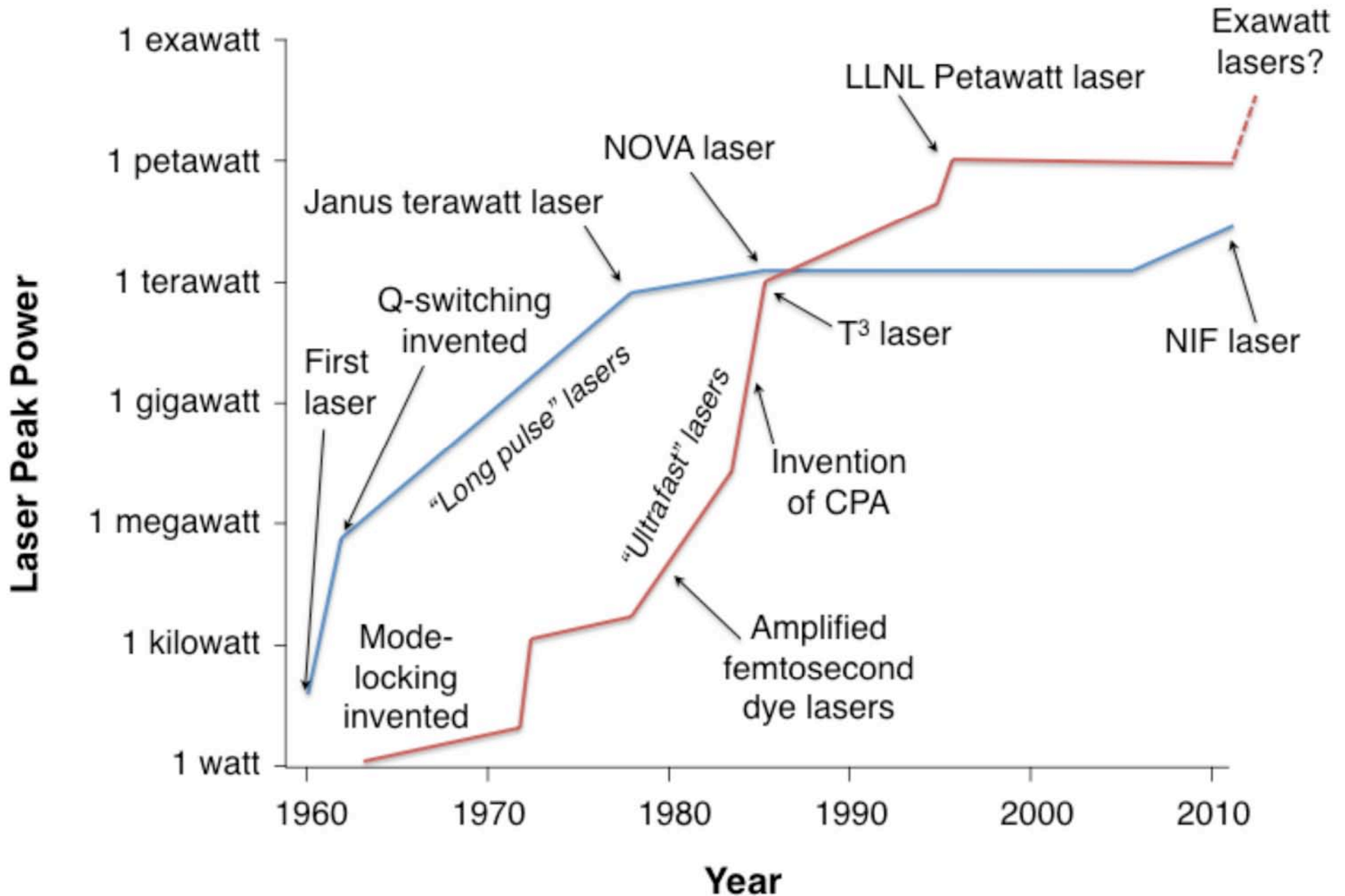
## Continuum Lasers

Curt Frederickson  
Larry Cramer

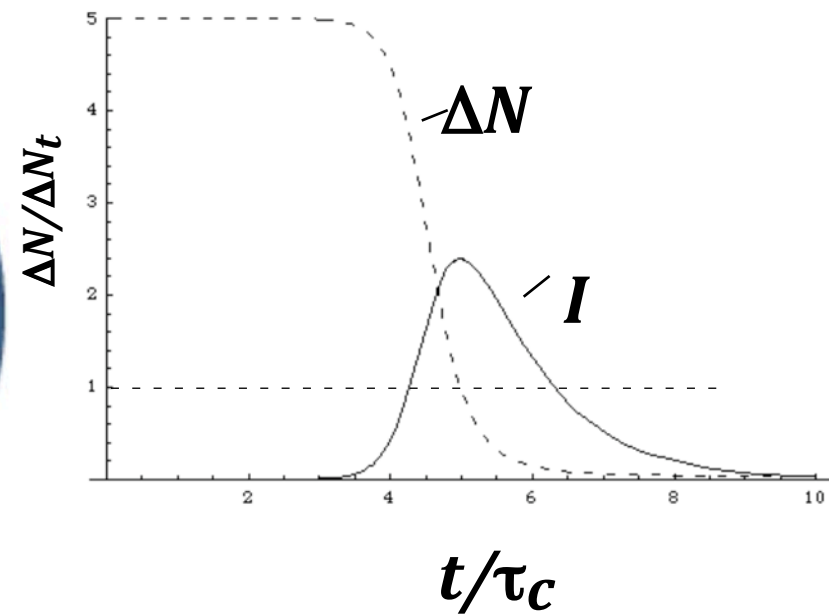
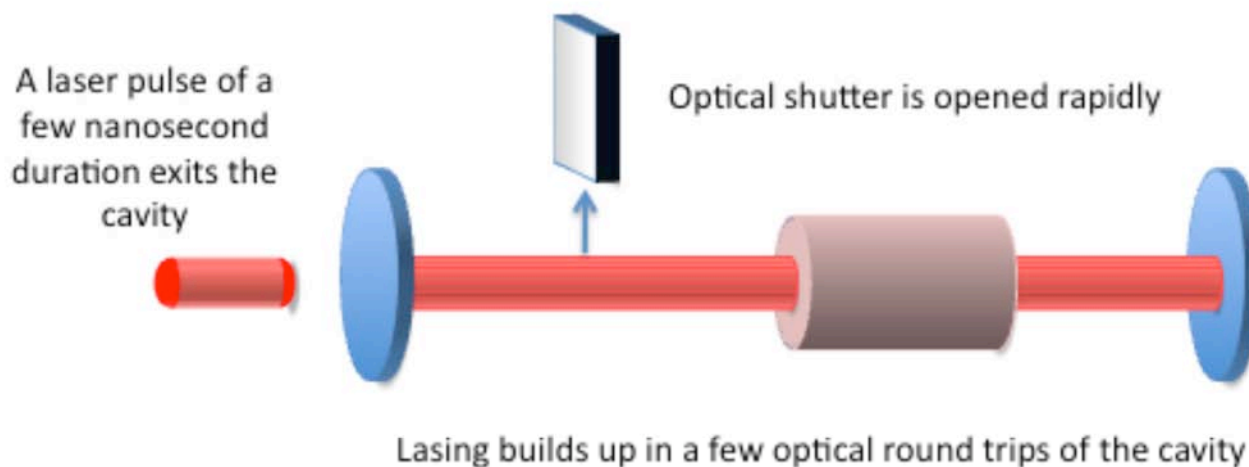
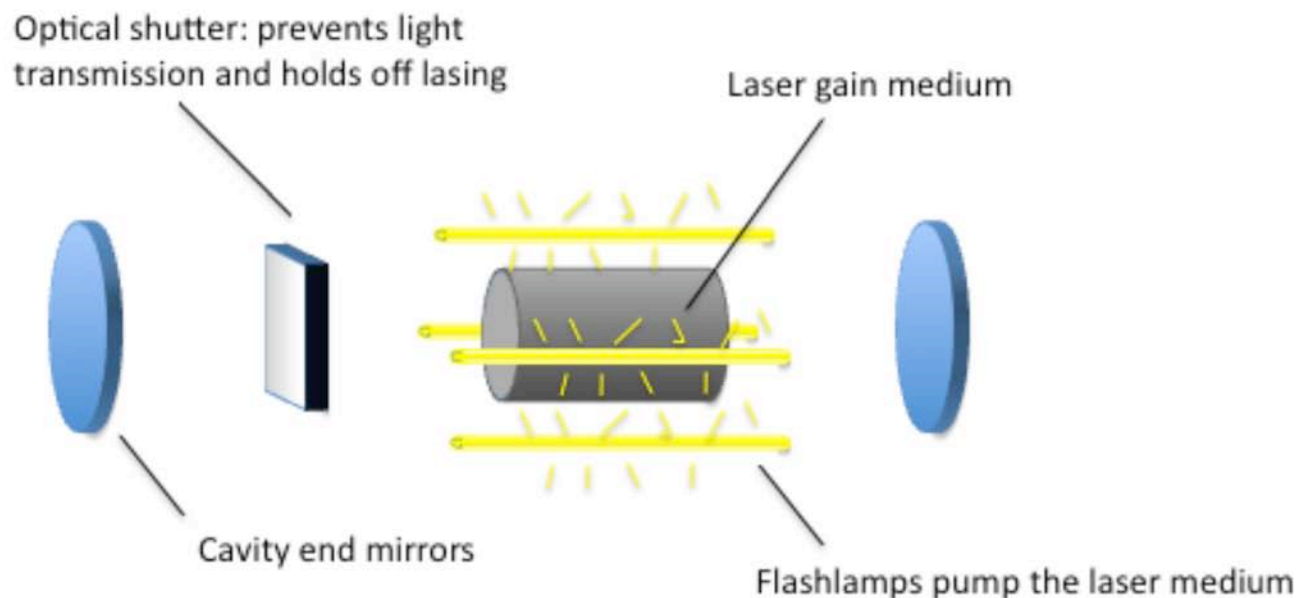
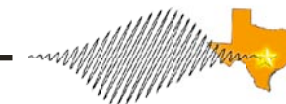
## LLE Rochester

Jon Zuegel

# Leaps in laser power have been made over the past 40 years because of technical break-throughs

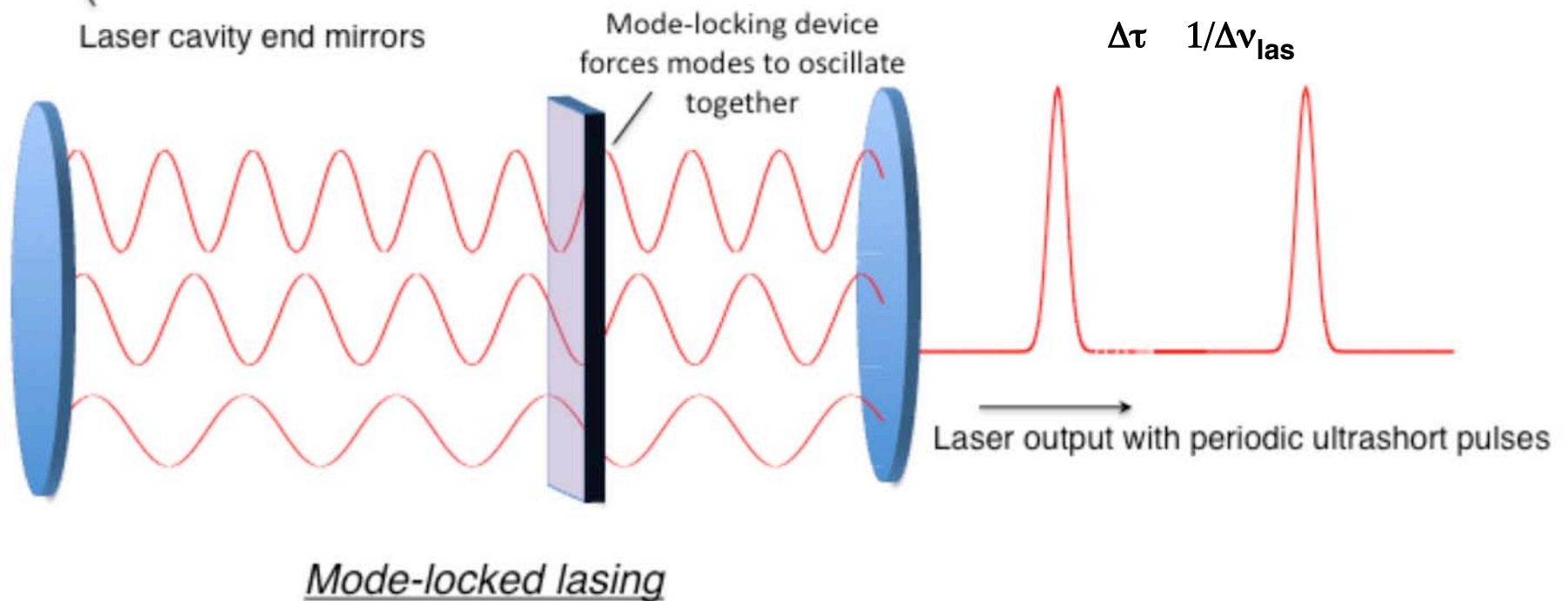
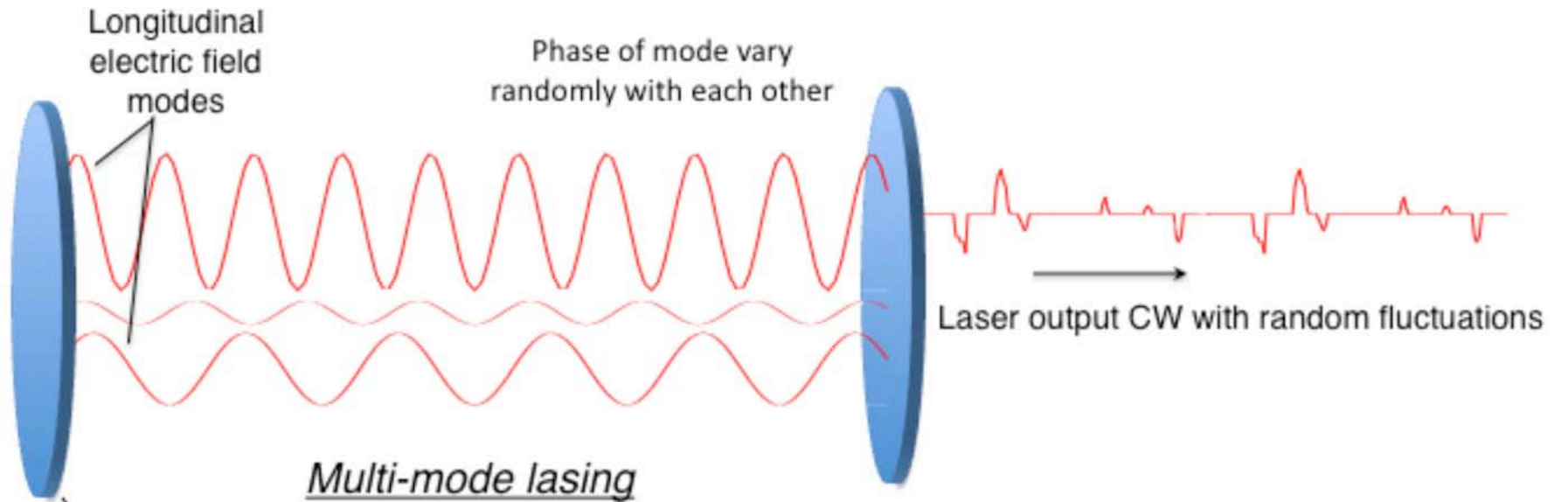
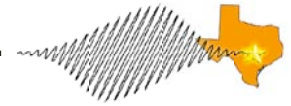


# In 1961 Q-switching was demonstrated enabling production of MW peak power, ns pulses

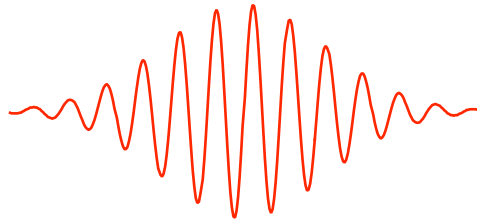




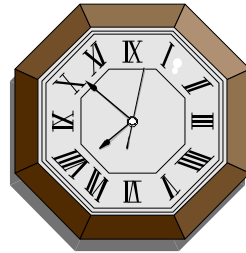
# In 1964 mode locking was demonstrated, enabling ultrafast pulse production



# A 30 femtosecond laser pulse is the same fraction of one minute that one minute is of the age of the universe



Ultrafast optical pulse:  
 $\sim 3 \times 10^{-14}$  sec



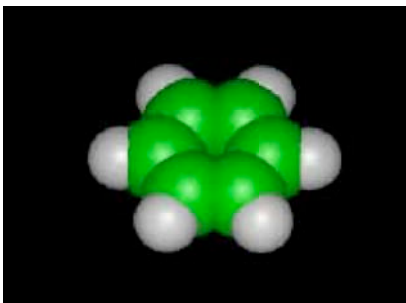
One minute:  
 $6 \times 10^1$  sec



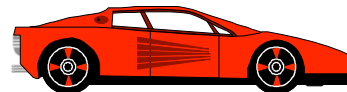
Age of the universe:  
 $\sim 2 \times 10^{17}$  sec



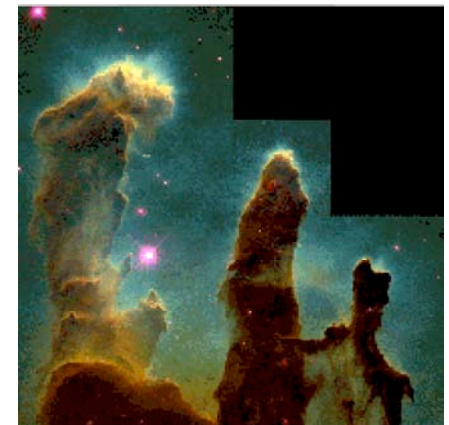
Time scale of  
atomic motions



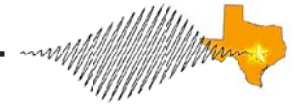
Time scale of  
human motions



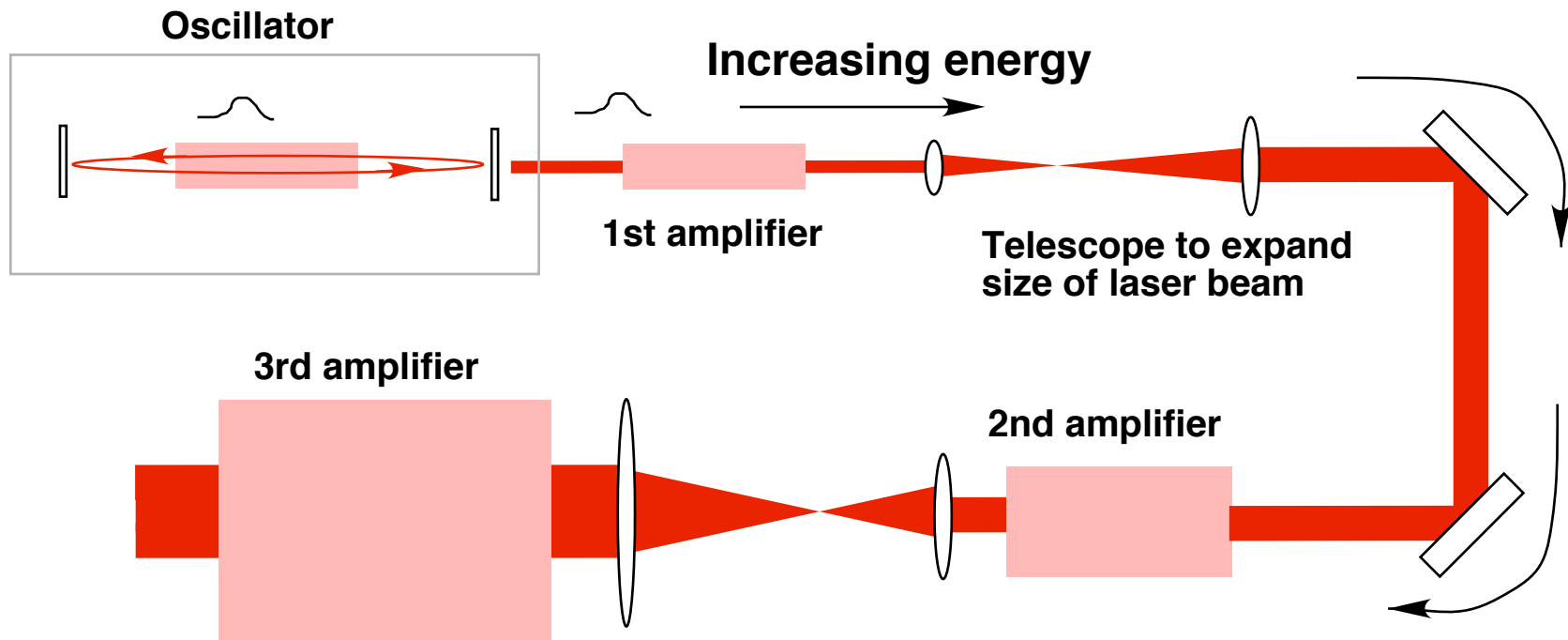
Time scale of  
interstellar  
motions



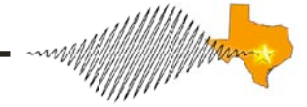
# Modern high power lasers are based on the “master oscillator, power amplifier” architecture



## MOPA laser chain



# Livermore has pushed MOPA architecture through a number of generations to higher and higher energy



**Cyclops laser (1974)**  
1 beam  
100 J



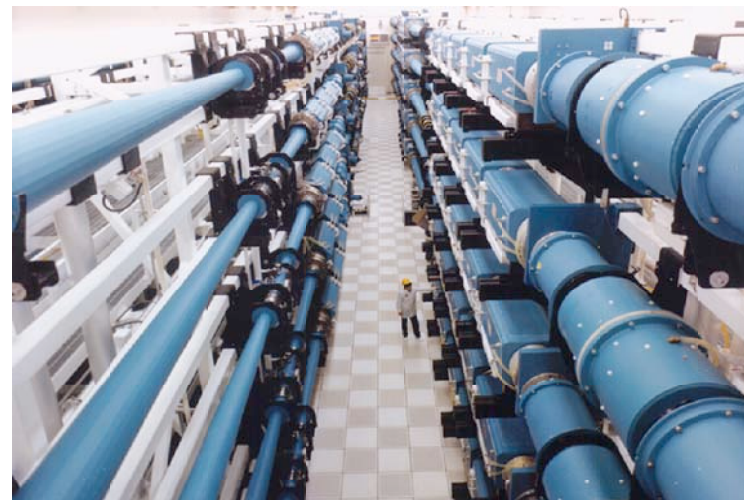
**Argus laser (1976)**  
2 beams  
1000 Joules of energy



**Shiva laser (1978)**  
20 beams  
10,000 Joules



**Nova laser (1985)**  
10 beams  
100,000 Joules





# The National Ignition Facility, recently completed at LLNL is the highest energy laser in the world

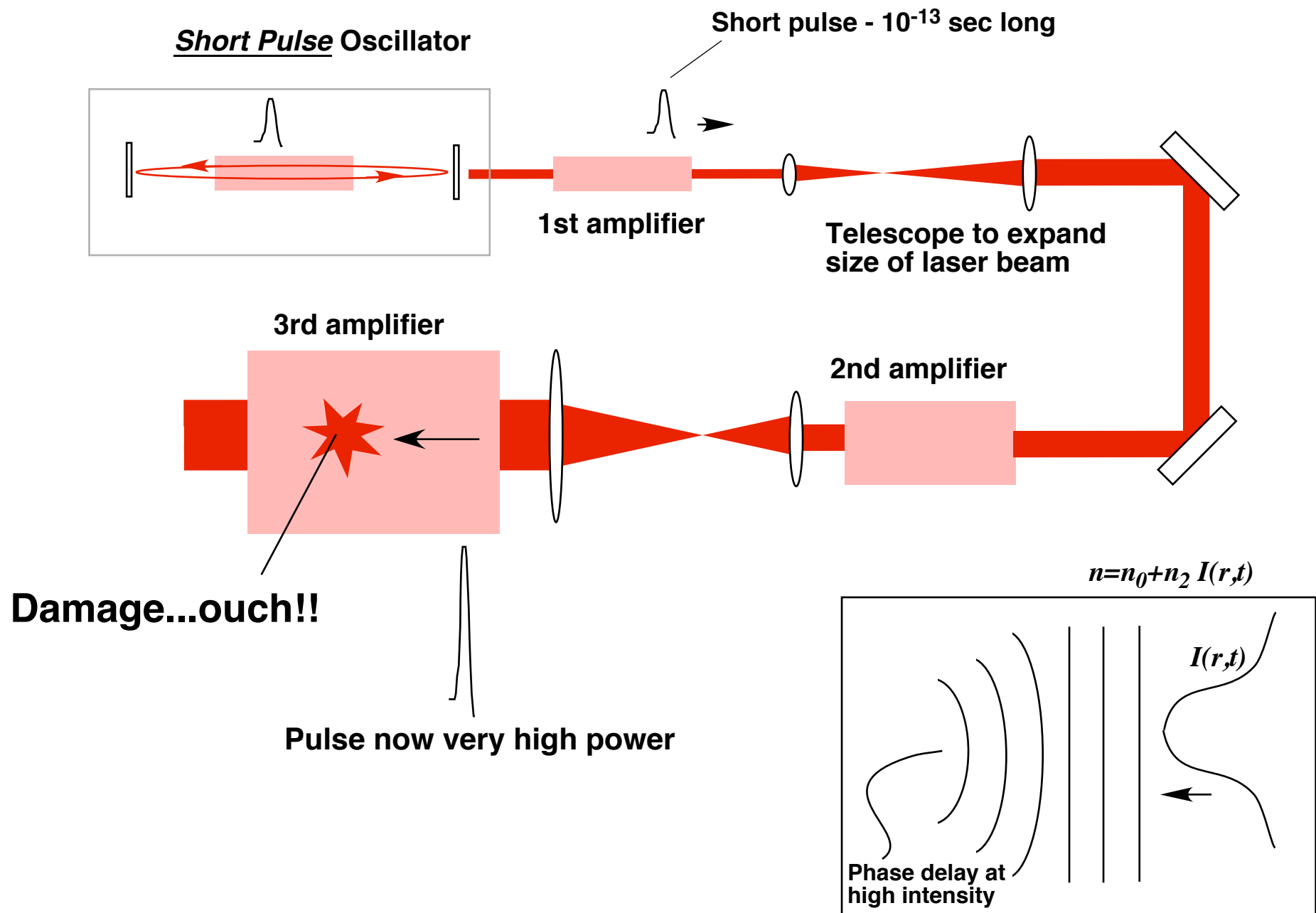
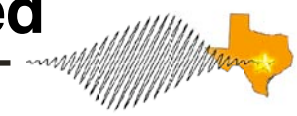
---



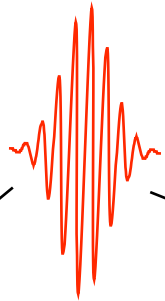
**National Ignition Facility - NIF (2009)**  
**192 beams**  
**1,800,000 Joules @ 351 nm**



# In the mid 1980's, the question of whether a short laser pulse could be amplified in a MOPA chain was broached



# The trick to amplify such short pulses is to stretch them out in time temporarily



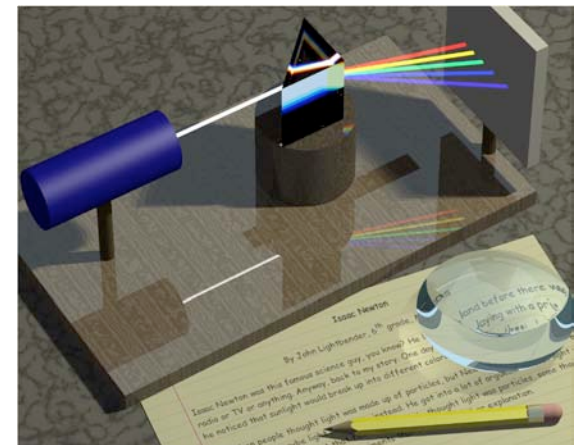
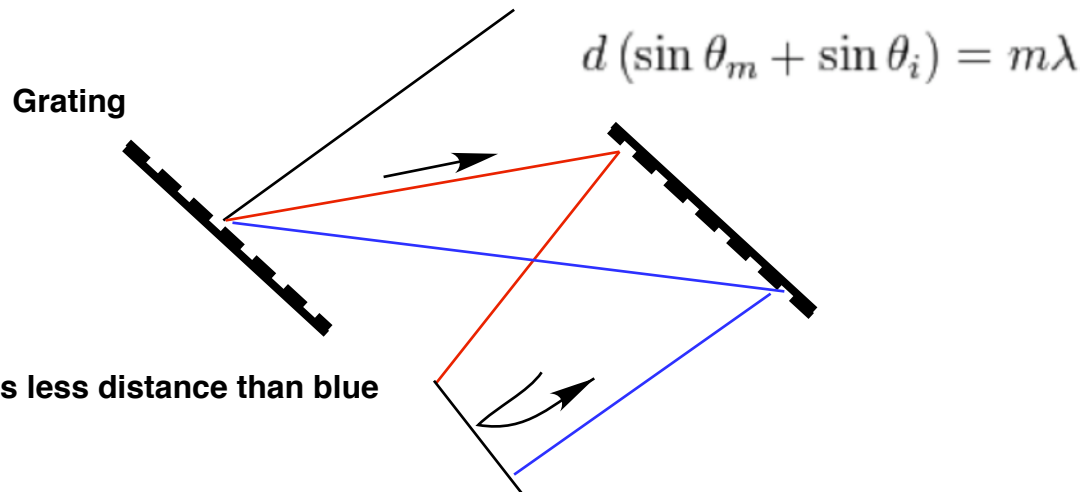
A very short laser pulse actually has a broad spectral bandwidth



*To disperse colors out in time, we spread them out in space*

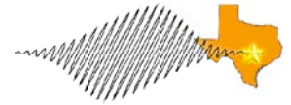


So...stretch out the spectrum

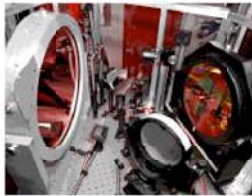




# Using this stretching trick, we can amplify very short pulses...and then recompress them after amplification



2) Using gratings, the pulse is stretched by 10,000 x



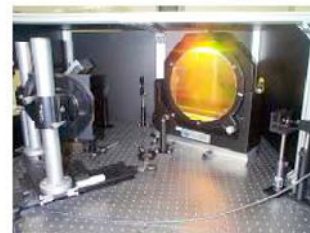
1) Oscillator produces short pulses



3) Long pulses are now safe to amplify in laser amplifiers



4) After amplification, the pulses are recompressed with gratings



5) Resulting pulse is short (20 - 1000 fs) and high energy (~ 1 J)

• Strickland and Mourou *Opt. Comm.* 56, 219 (1985)

This technique has come to be called “Chirped Pulse Amplification” (CPA)



# The current state-of-the-art ultrafast, ultraintense lasers tends to fall into two categories



## *Ti:sapphire based CPA lasers:*

Pulse energy  $\sim .001 - 30 \text{ J}$ ,  
Pulse duration  $< 30 - 100 \text{ fs}$ ,  
Peak Power  $< 100 \text{ TW}; 1 \text{ PW}$   
Repetition Rate  $\sim 1 \text{ kHz} - 1 \text{ Hz}$

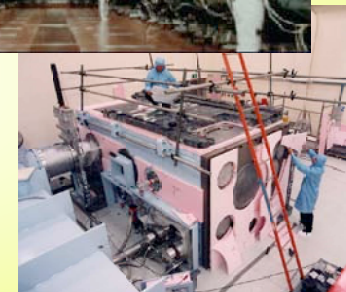
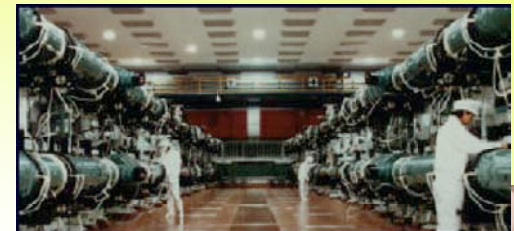
Shortest pulse systems and most  
“table-top” CPA lasers



## *Nd:glass based CPA lasers*

Pulse energy  $10 - 1000 \text{ J}$   
Pulse duration  $> 100 \text{ fs}$   
Peak power  $10 - 1000 \text{ TW}$   
Repetition rate  $\sim 1 \text{ shot/min} - 1 \text{ shot/hr}$

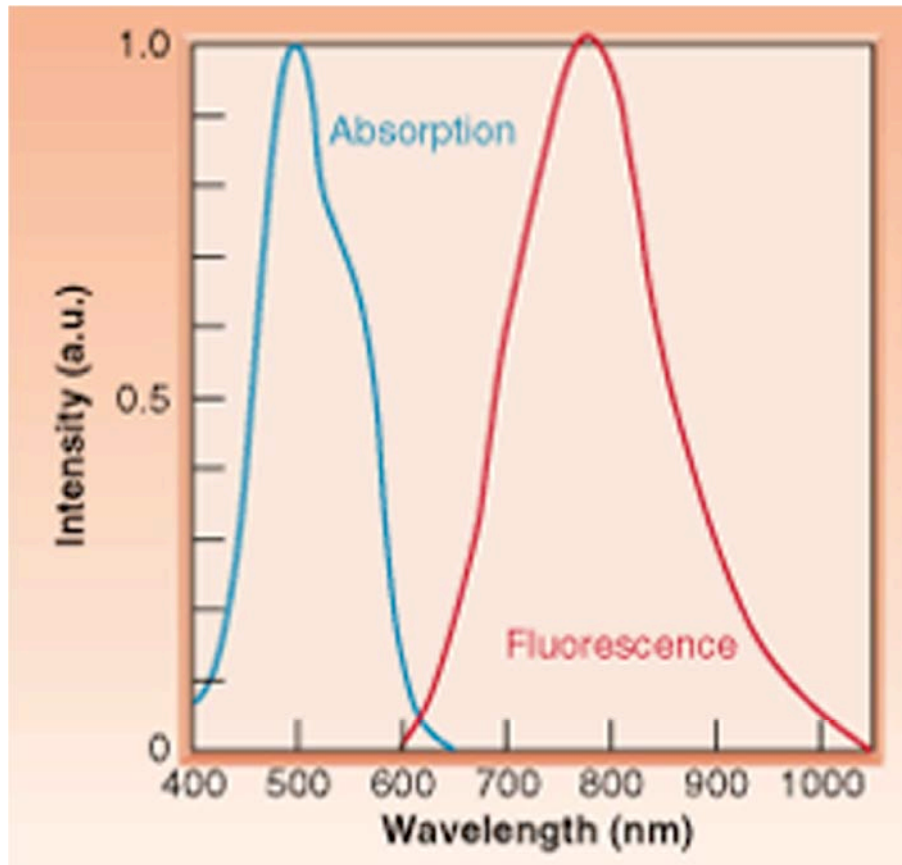
Highest energy systems, many of “facility” scale



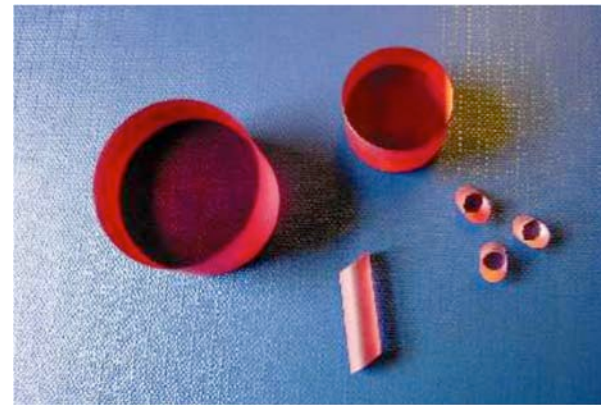
# Ti:sapphire has advantages and disadvantages in high power CPA lasers



## Absorption and emission spectrum of Ti:sapphire



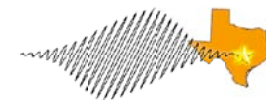
## Large scale Ti:sapphire crystals



Gain bandwidth in Ti:sapphire is very large → amplification of pulses as short as 20 fs

High quality Ti:sapphire can only be produced with aperture up to ~10 cm

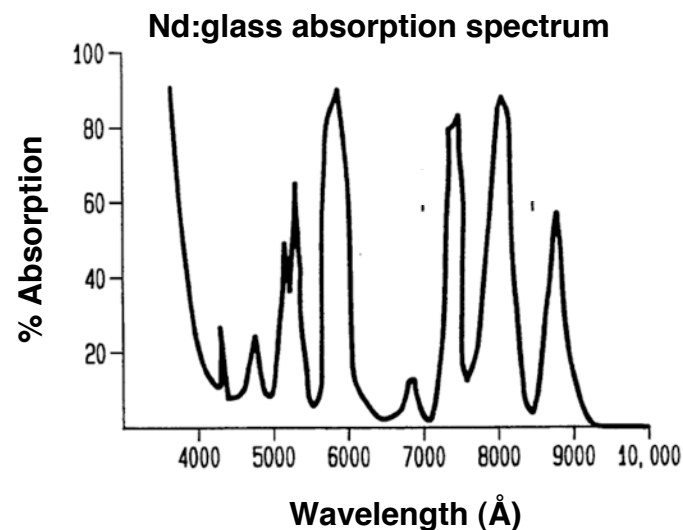
# Nd:glass is very attractive for high power lasers because it can be fabricated with large aperture



*Large aperture amplifiers can be fabricated*

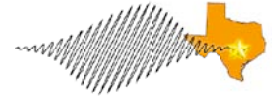


*Nd:glass can be efficiently pumped by flashlamps*

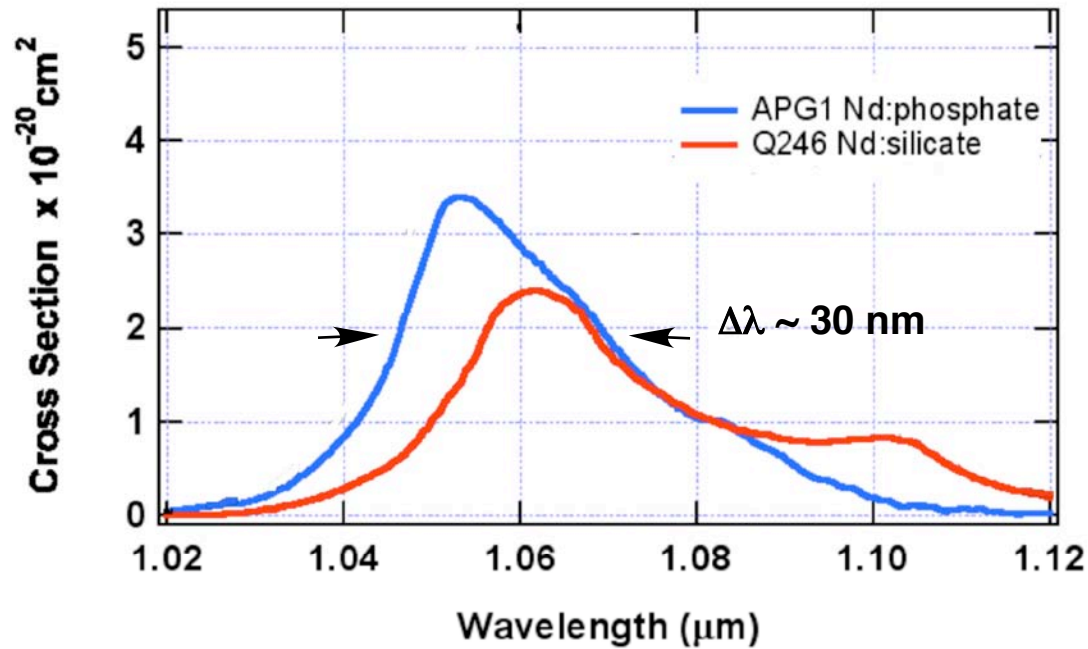




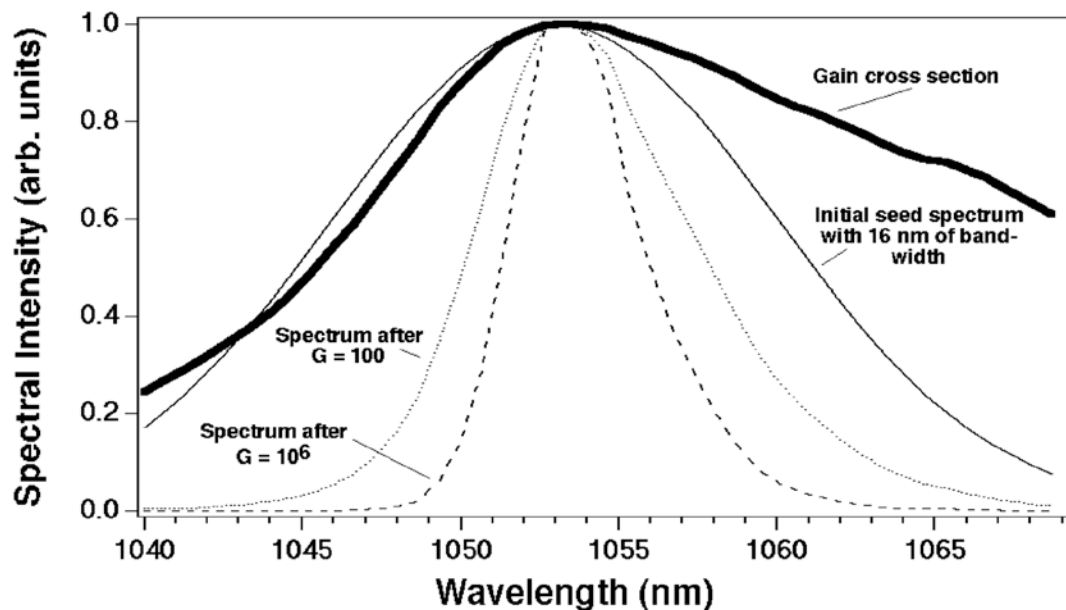
# The principal limitation to the use of Nd:glass in CPA lasers is that it exhibits limited gain bandwidth



*Gain spectrum of two kinds of laser glass*



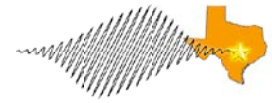
*Calculation of the effects of gain narrowing in Nd:glass*



**Gain narrowing of the ultrafast pulse spectrum tends to limit Nd:glass CPA lasers to pulse duration of 500 fs**

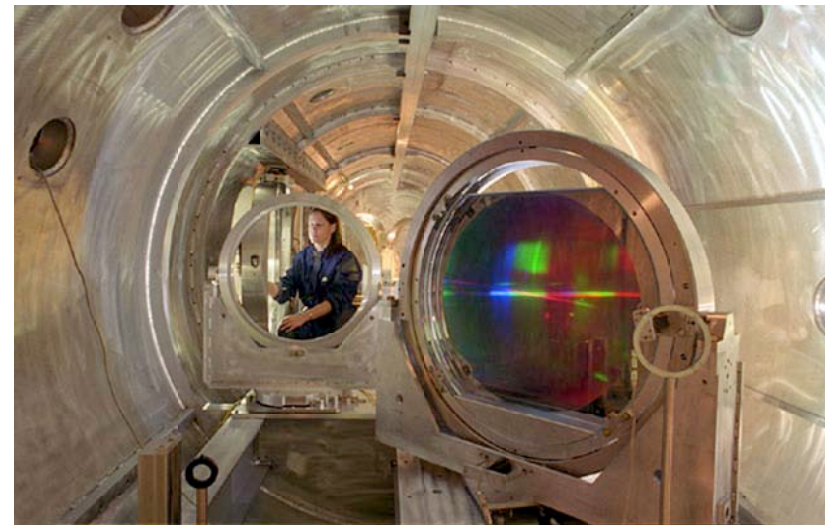


# The first Petawatt laser was demonstrated at Lawrence Livermore by implementing CPA on the NOVA laser



## The Petawatt at LLNL

Nova laser



90 cm gratings to compress Nova pulses

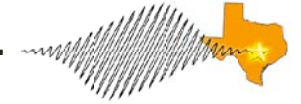
Petawatt specs:

500 J energy

500 fs pulse duration

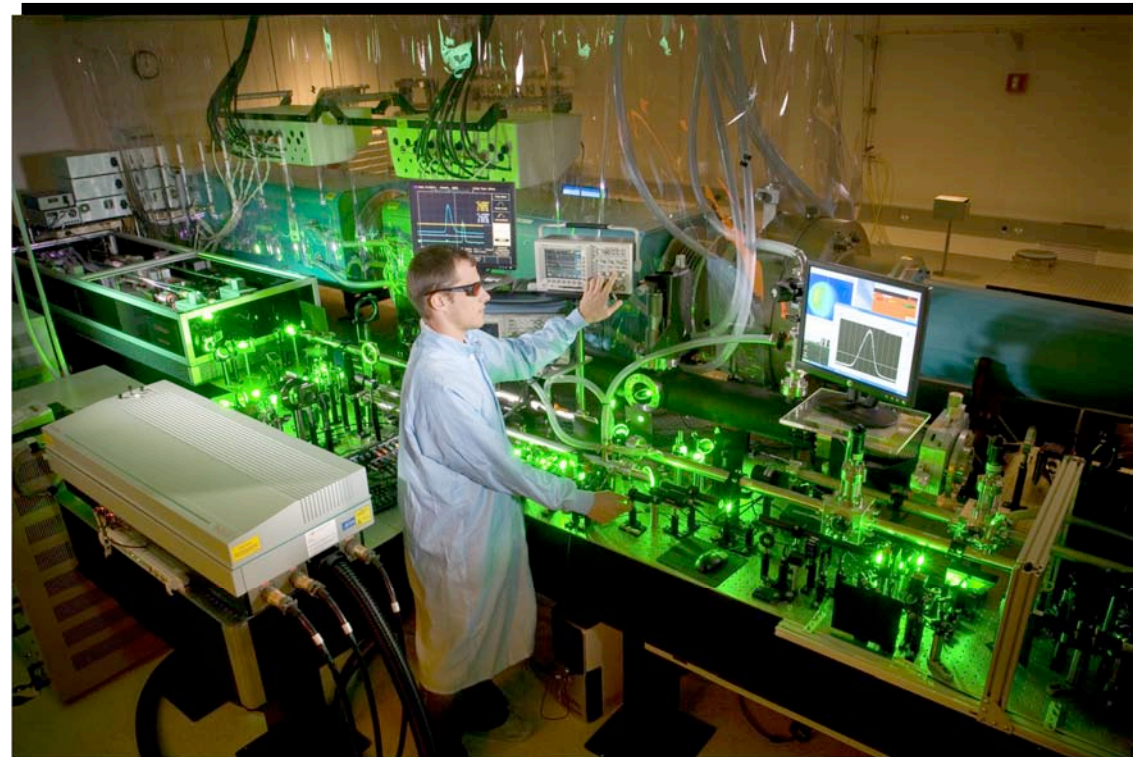
Peak intensity  $> 10^{20}$  W/cm<sup>2</sup>

# A “Petawatt” is many times more power than all the power delivered by all the power plants in the US



Power output of U.S. electrical grid:  
~ 80,000 GWh/week = 0.5 TW

Petawatt laser

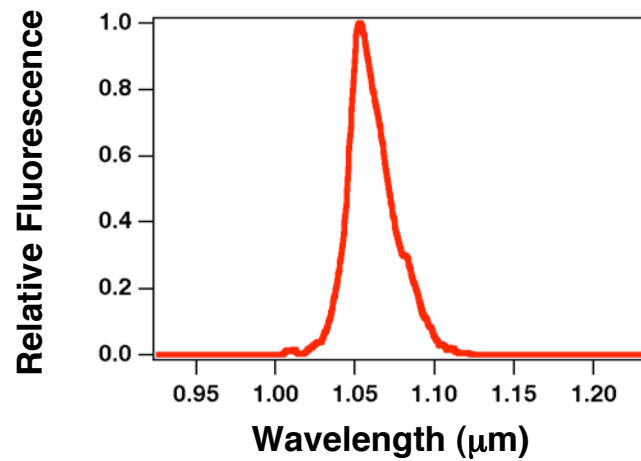


***A state-of-the-art petawatt laser has 2000 times the power output of all power plants in the US***

# There are different kinds of laser glass which each have slightly different center wavelengths



**Q-98 Phosphate glass**



**Peak Wavelength: 1054 nm**

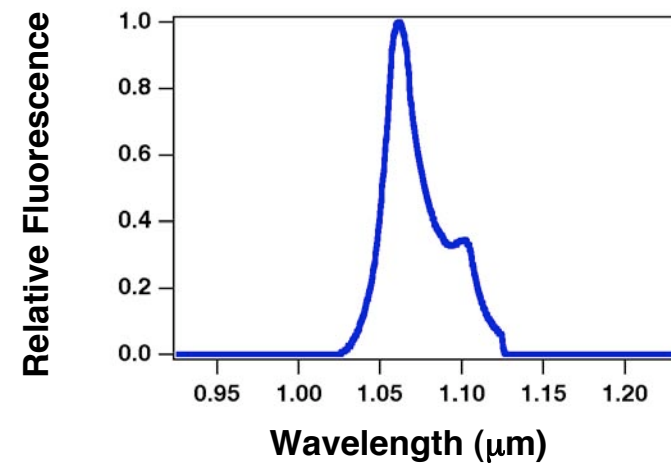
**Peak cross section:  $4.3 \times 10^{-20} \text{ cm}^2$**

**Linewidth (FWHM): 21.1 nm**

**$\text{Nd}_2\text{O}_3$  ~3%**

**$\text{P}_2\text{O}_5$  ~97%**

**LG-680 Silicate glass**



**Peak Wavelength: 1061 nm**

**Peak cross section:  $2.9 \times 10^{-20} \text{ cm}^2$**

**Linewidth (FWHM): 28.2 nm**

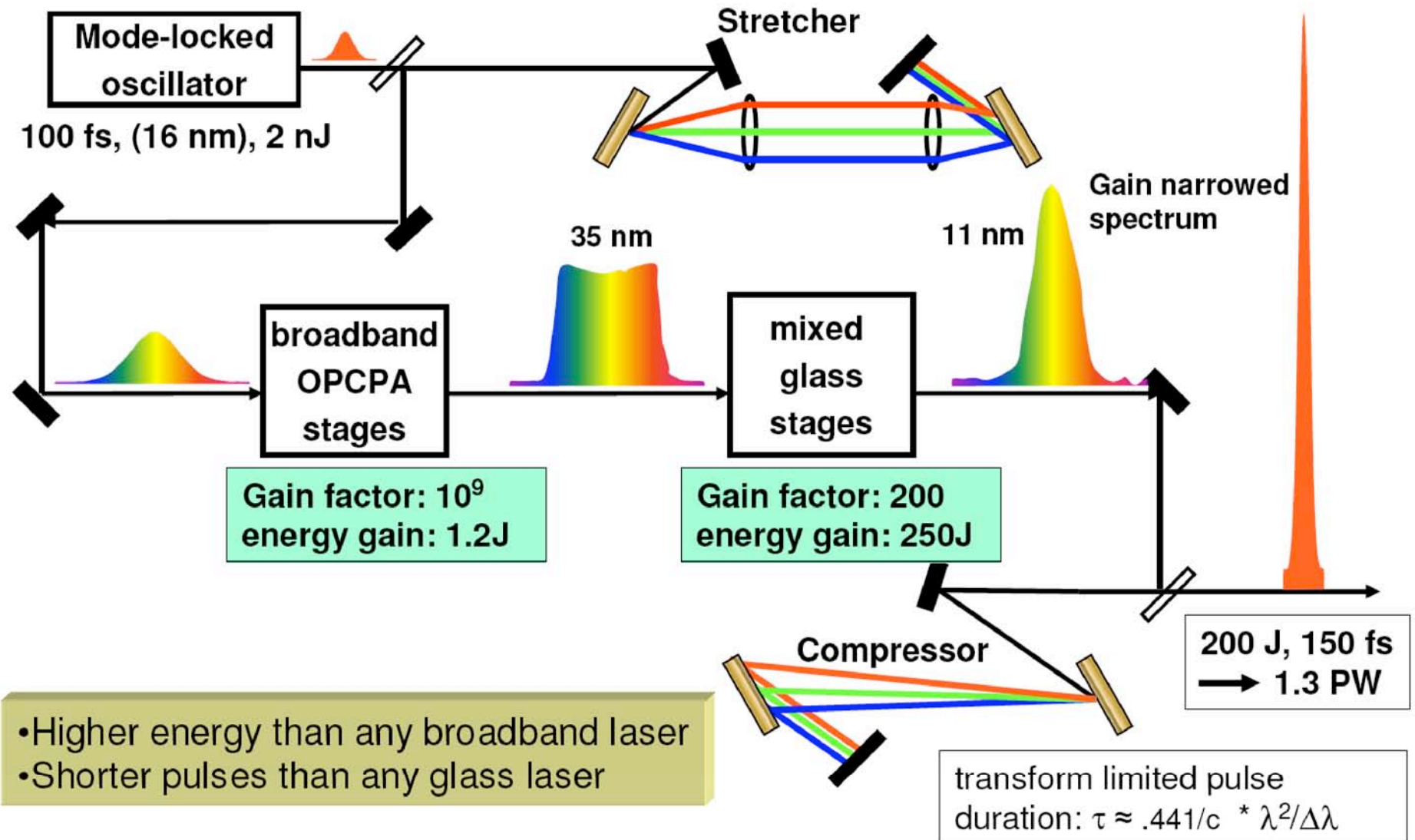
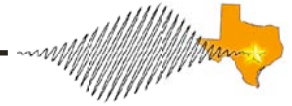
**$\text{Nd}_2\text{O}_3$  ~3%**

**$\text{SiO}_2$  ~97%**



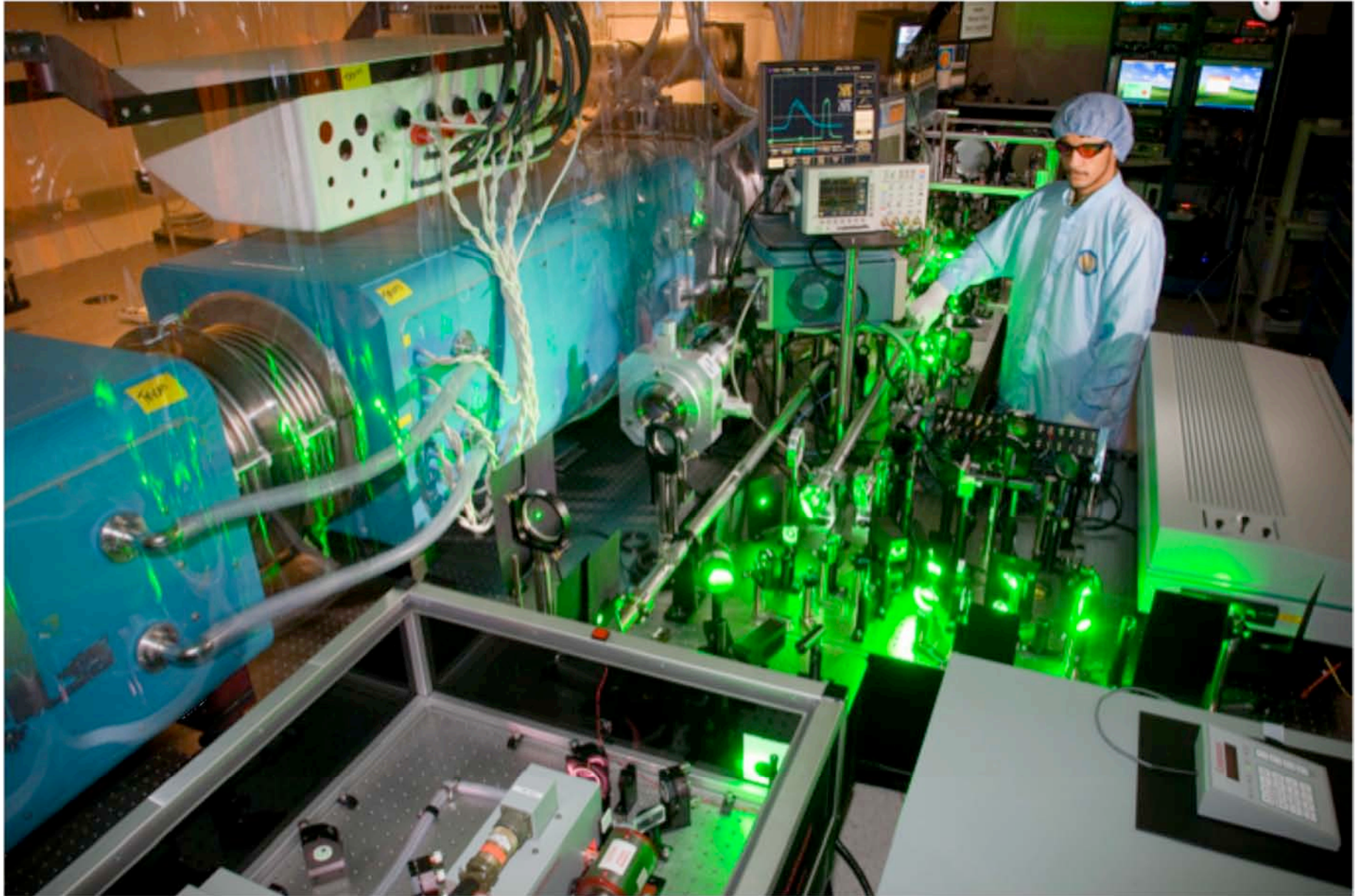
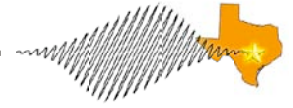


# The Texas Petawatt is designed to maintain broad bandwidth through the entire chain





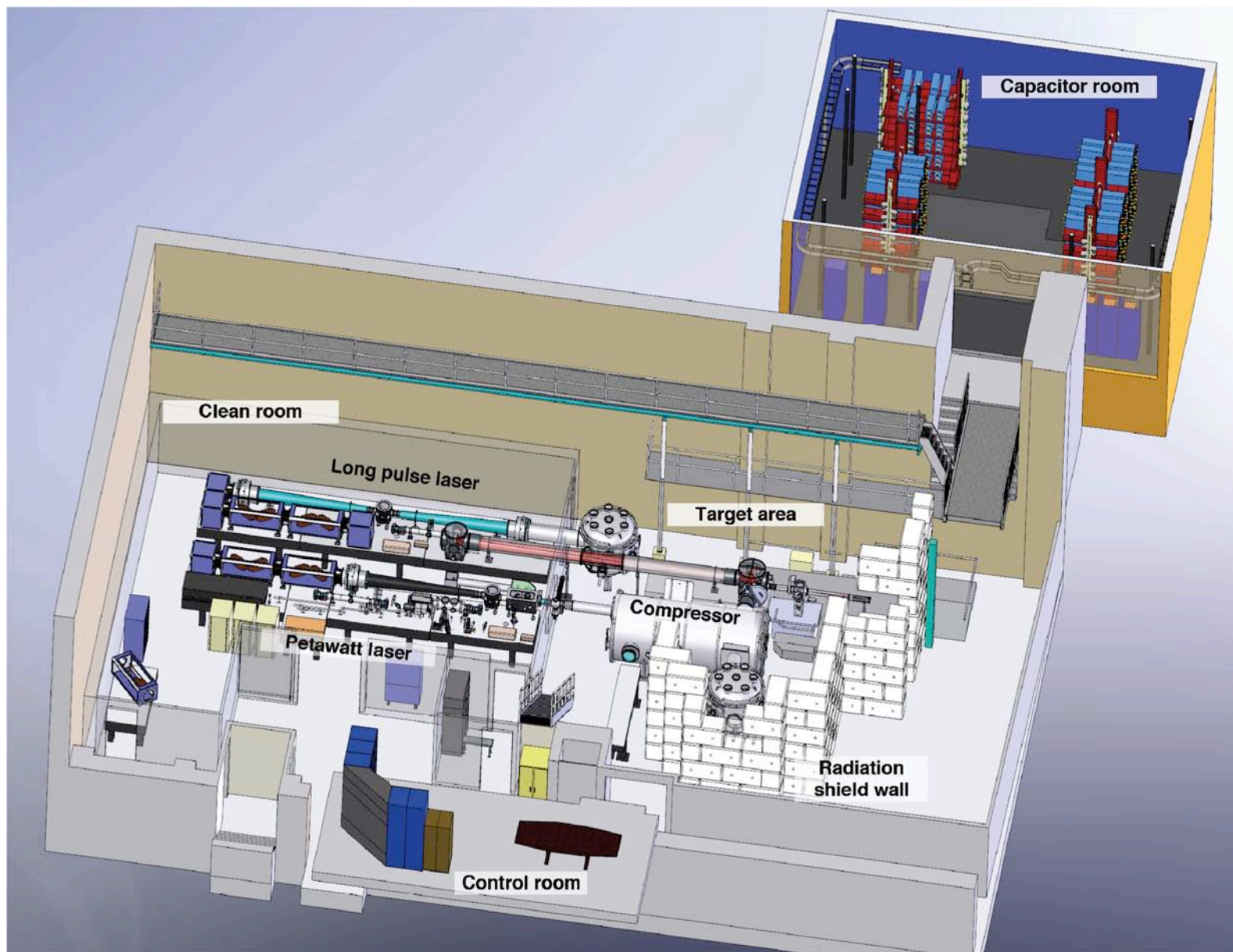
# The Texas Petawatt uses four-passes in NOVA 31 cm disk amplifiers



# The Texas Petawatt Laser is housed in the RLM High Bay

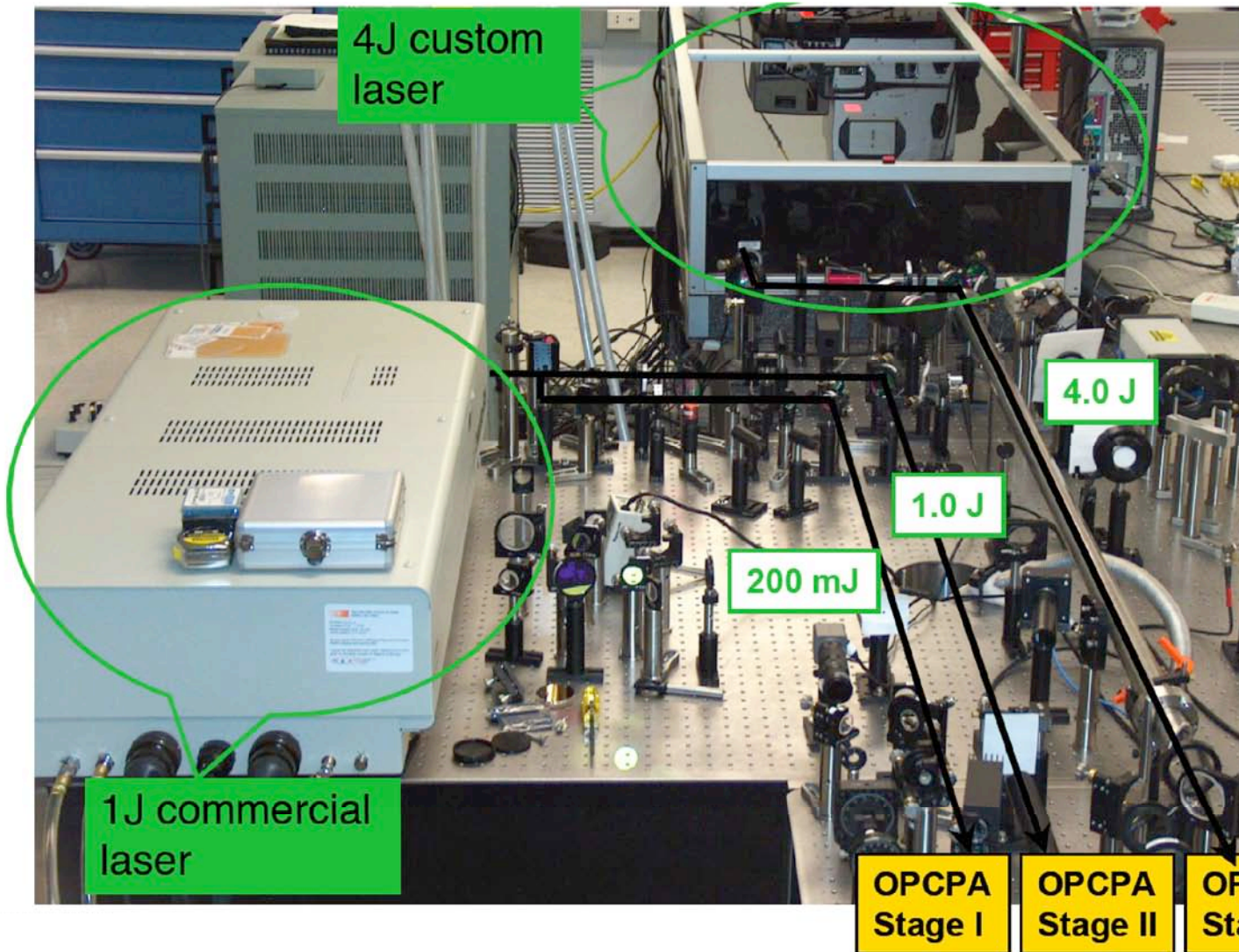
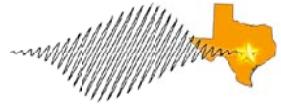


## Layout of the Texas Petawatt Facility in the Robert Lee Moore Basement

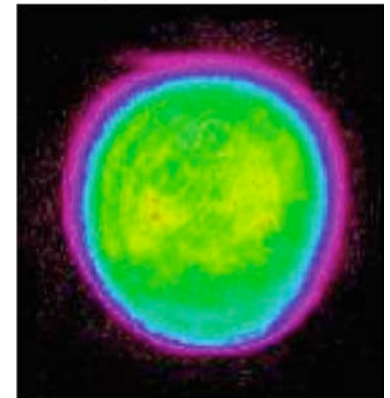




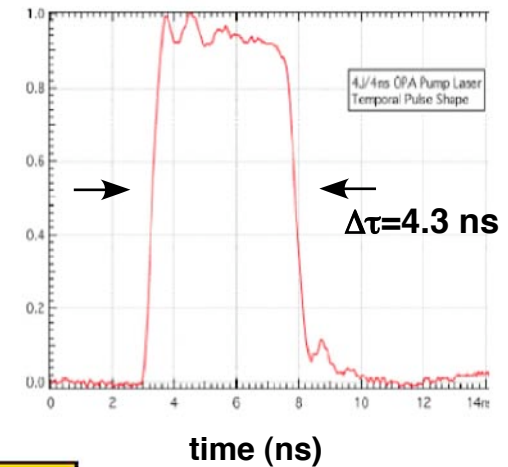
# We are operating an OPCPA front end whose principal component is a custom 4J pump laser



Pump beam spatial profile @ 4J

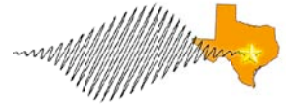


Pump beam temporal profile @ 4J





# The Nd:silicate rod amplifier has been installed and tested

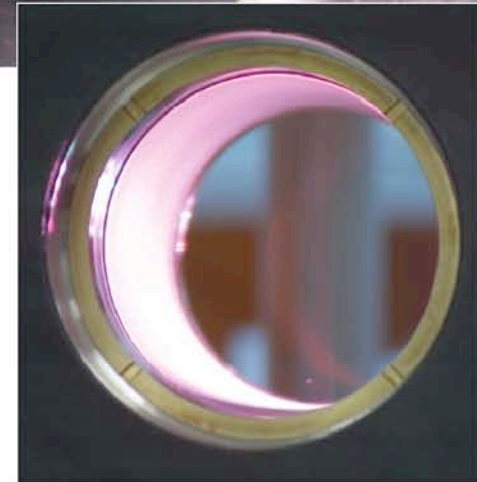


The capacitor bank is above the cleanroom

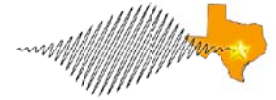


64 mm rod amplifier with N0306 silicate glass

70 mm static Faraday rotator



# Two 315 mm disk amplifiers from NOVA are used in each beamline as the final laser amplifiers



The NOVA Laser at LLNL

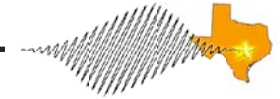


31 cm Disk amps from NOVA



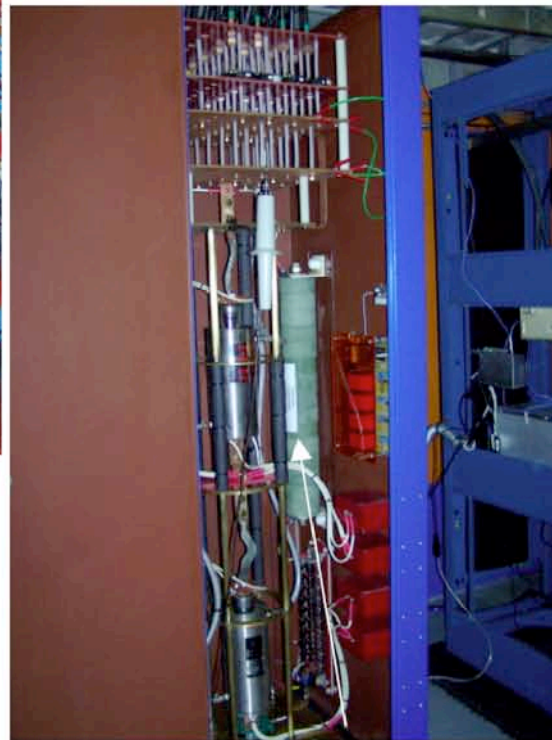
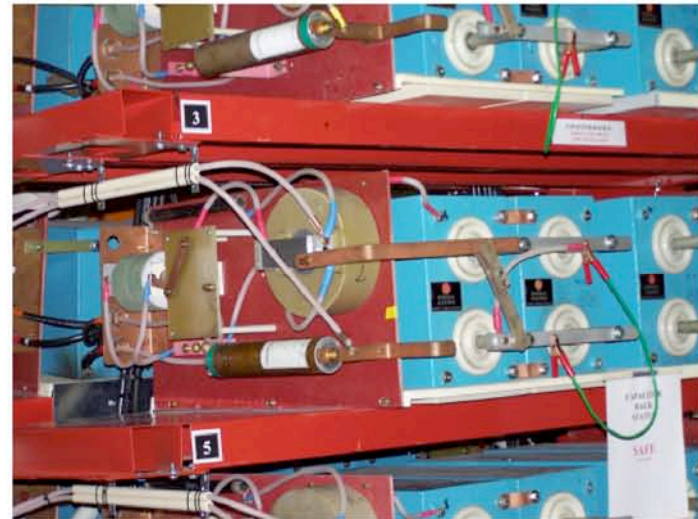


# The pulsed power for the 31 cm disk amplifiers is assembled and delivers 500 kJ of electrical energy

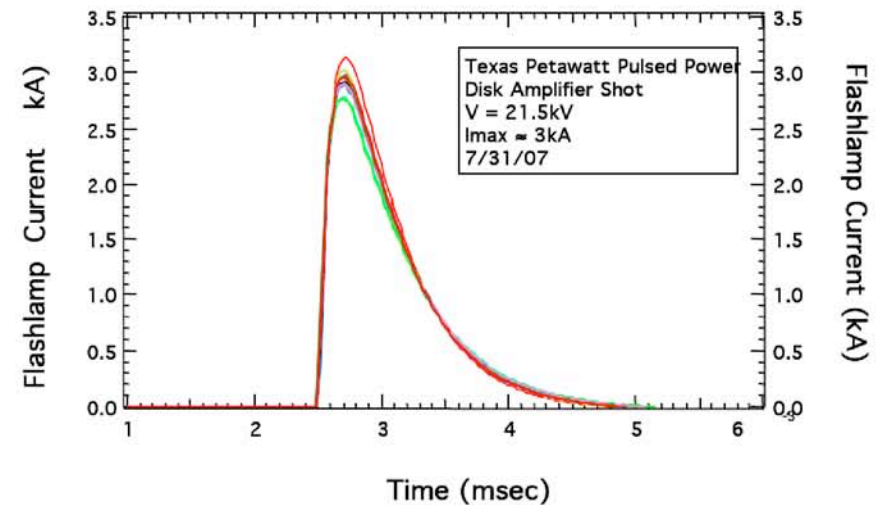


One side of bank drives  
One of two 31 cm amps

Pulse forming networks:  
2 52uF Capacitors  
1 450uH Inductor  
Critically Damped @ 400ms

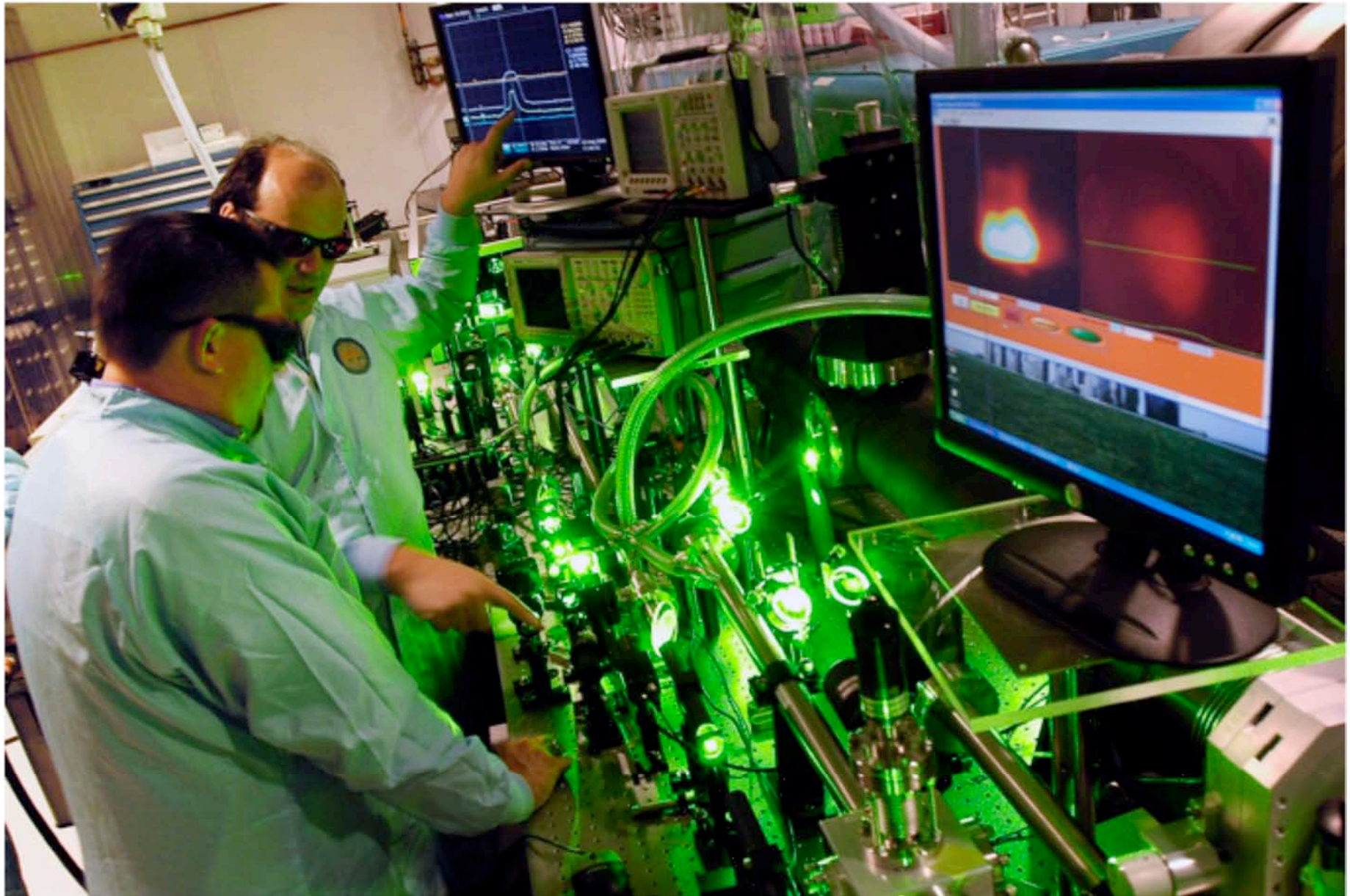
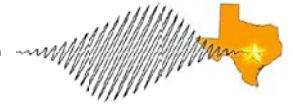


Two series ignitrons switches

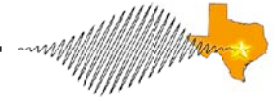




# Here Erhard Gaul explains to Todd Ditmire what a “laser” is



# Concrete shielding blocks have been installed in the target bay



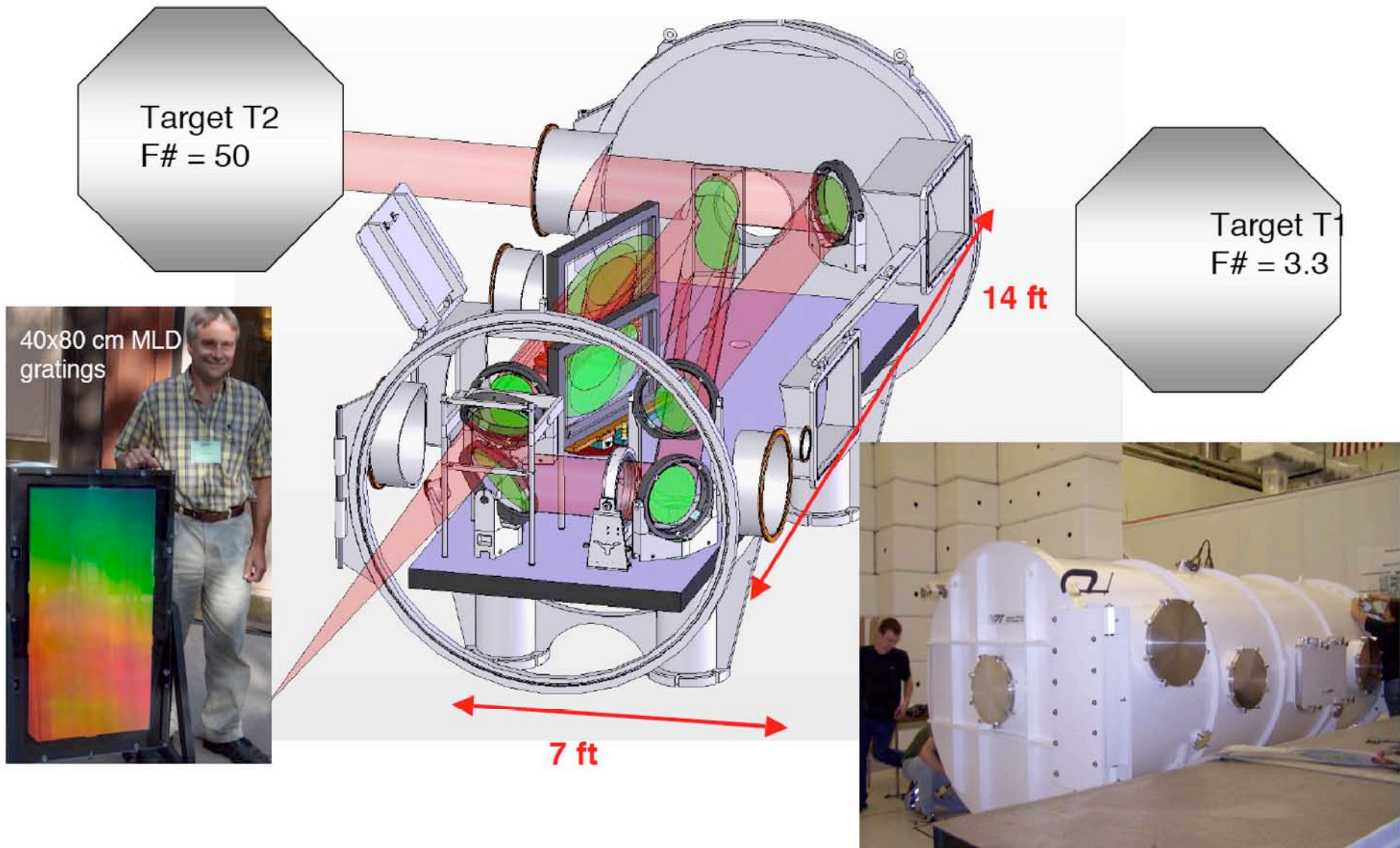
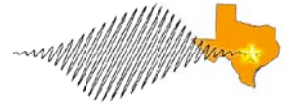
**Texas Petawatt Target Bay  
prior to shielding installation**



**Texas Petawatt Target Bay with shielding  
installed around the Petawatt beam target area**

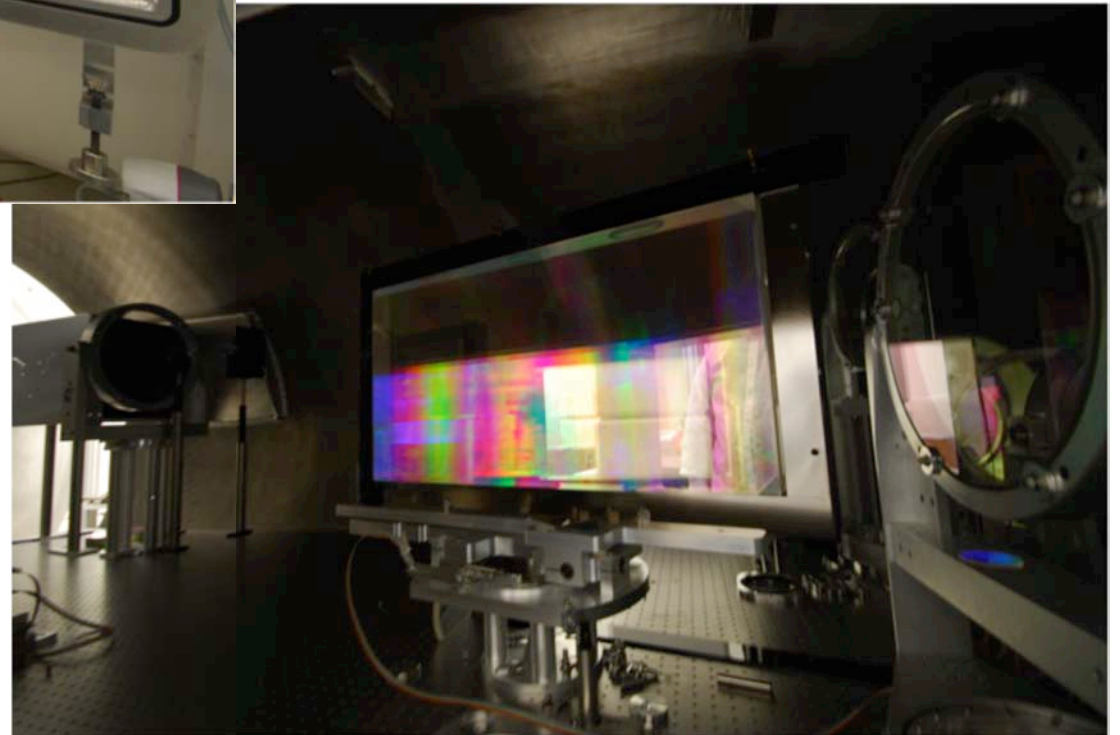
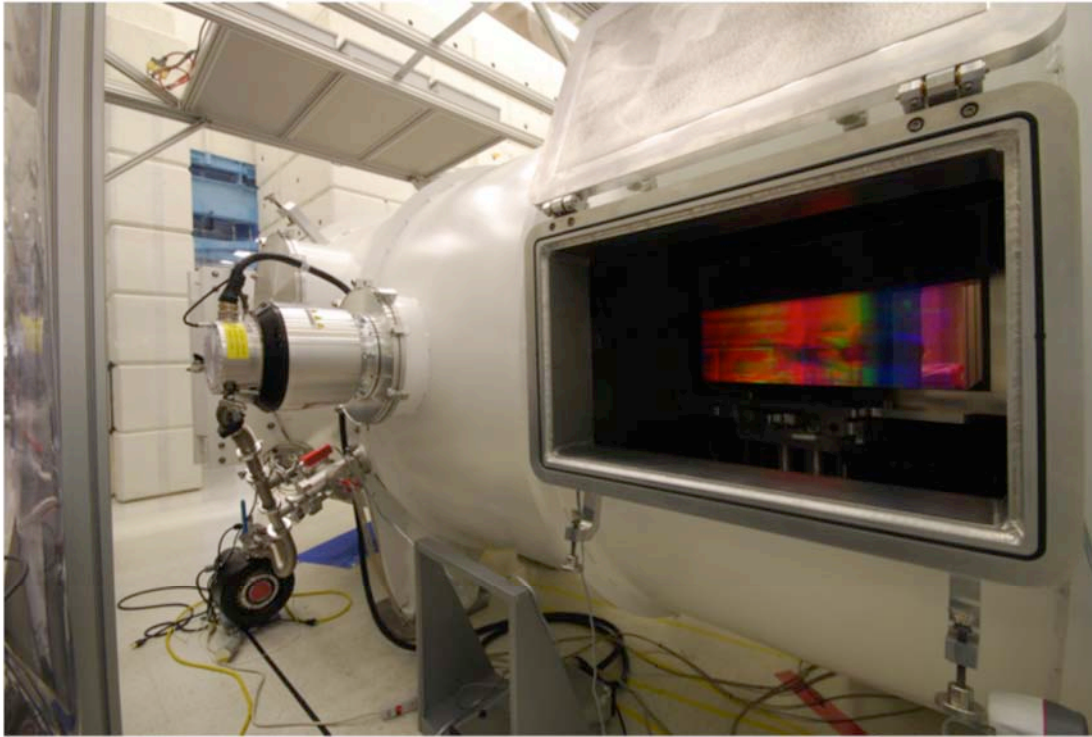
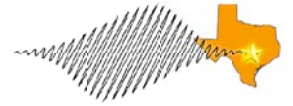


# Two MLD 80x40 cm gratings compress the TPW pulse

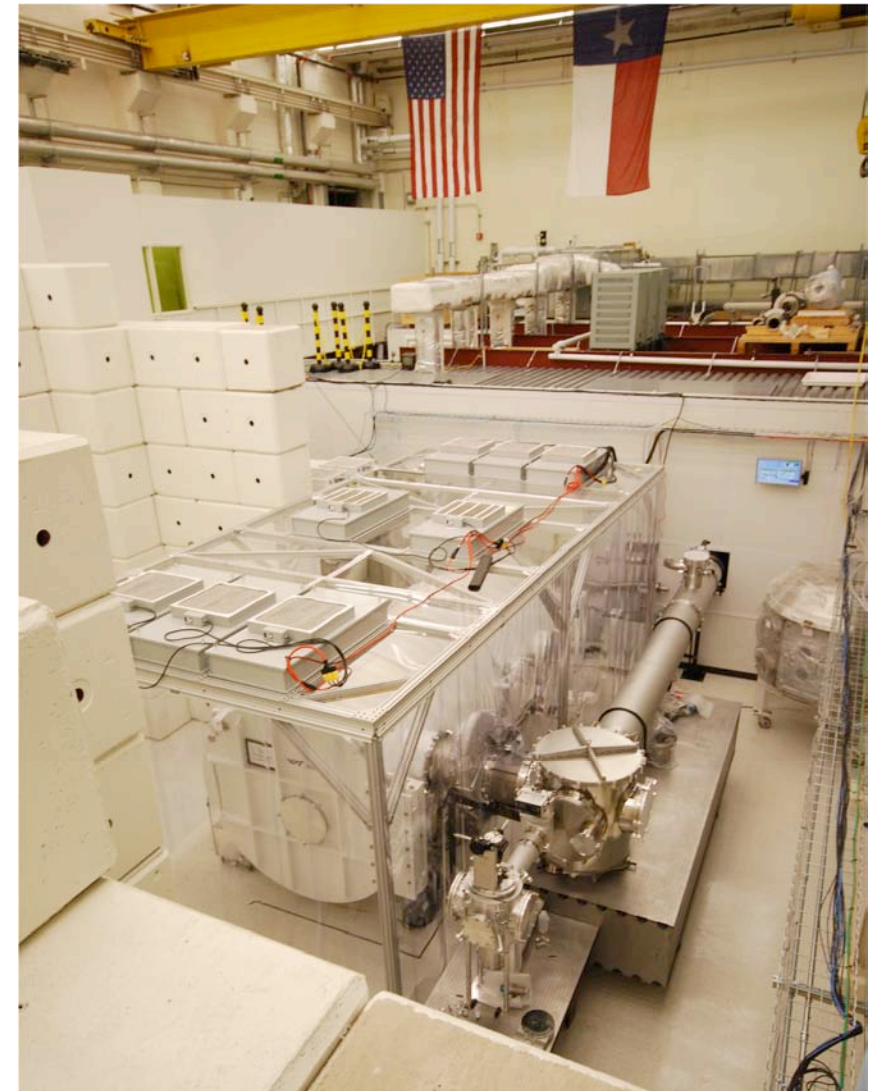
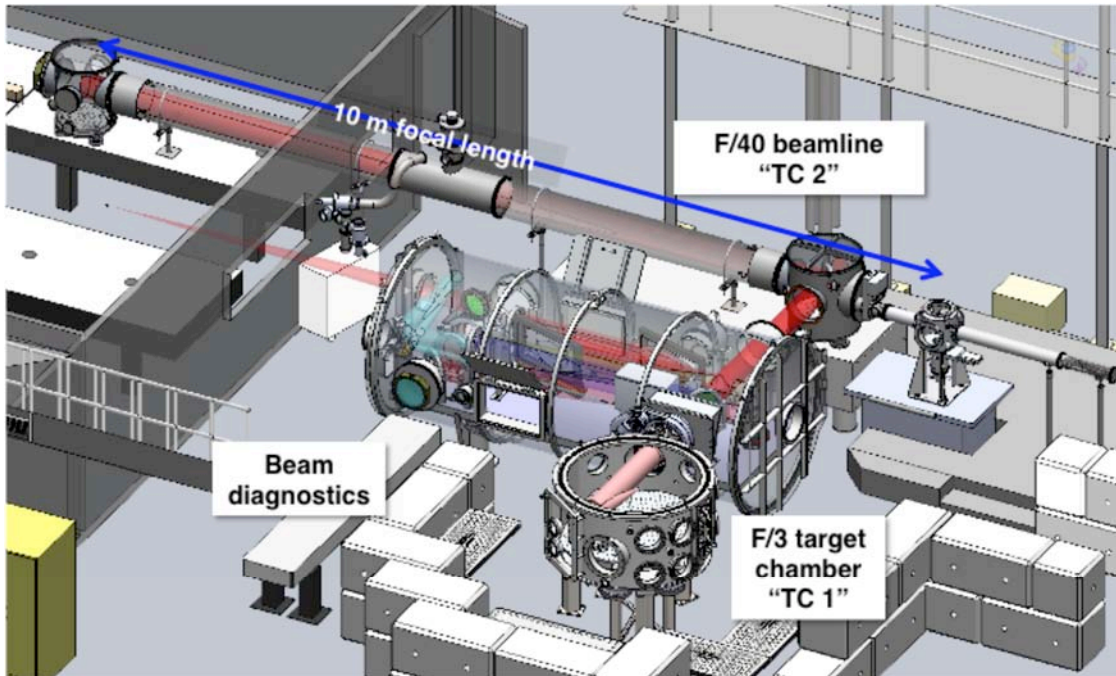
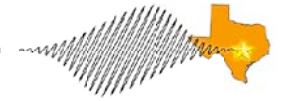




# The MLD gratings in the TPW perform well with high diffraction efficiency

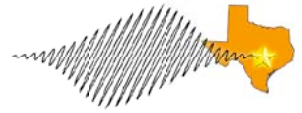


# The Texas Petawatt Laser has multiple target areas





# A modern ultra intense laser is the brightest known source of light

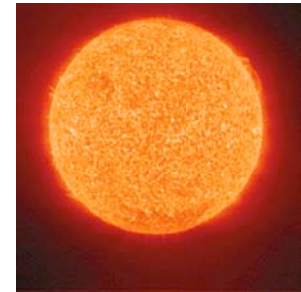


## Light intensity...

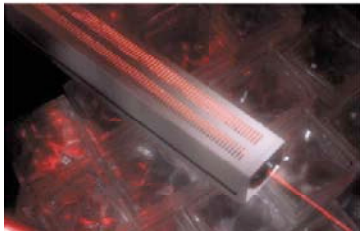


...near a light bulb  $\sim .03 \text{ W/cm}^2$

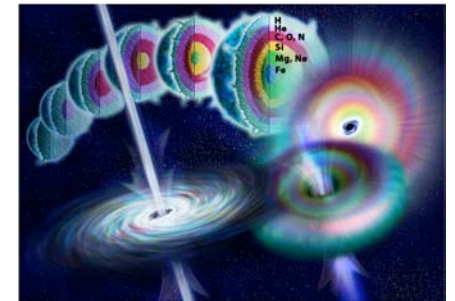
... of sun light on Earth  $\sim .14 \text{ W/cm}^2$



...of a focused HeNe laser  $\sim 500 \text{ W/cm}^2$



...at the surface of the sun  $\sim 7000 \text{ W/cm}^2$



...near a black hole during a gamma ray burst  
 $\sim 10,000,000,000,000,000,000 (10^{20}) \text{ W/cm}^2$

...of a petawatt CPA laser  
 $> 1,000,000,000,000,000,000,000$   
(a billion-trillion or  $10^{21}$ )  $\text{W/cm}^2$





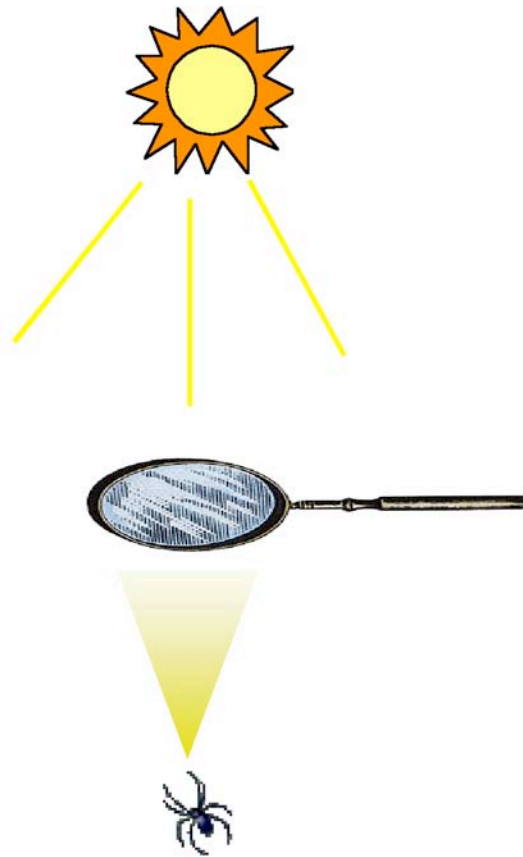
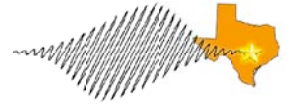
**What happens to matter irradiated at such extreme intensities?**

---



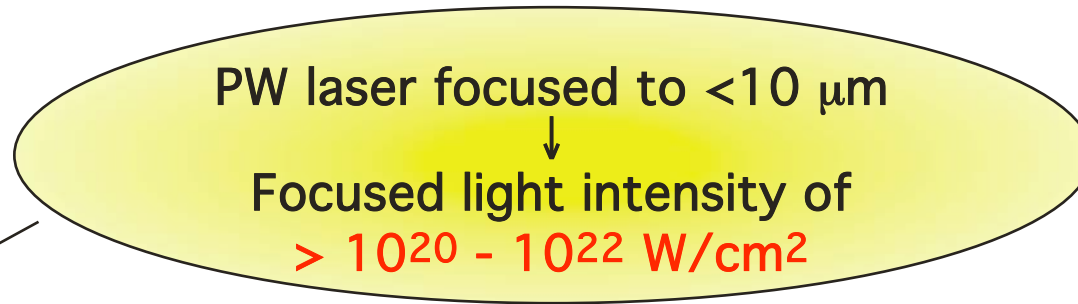
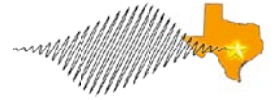
# What happens to matter irradiated at such extreme intensities?

---



**It gets hot!**

# Petawatt lasers access extreme regimes of physical parameter space



## High Field Science

### High electric fields

$$E \sim 10^{11} - 10^{12} \text{ V/cm}$$

Field strength is 100 to 1000 times that of the electric field felt by an electron in a hydrogen atom

### High electron quiver energy

$$U_{\text{osc}} = 1 \text{ MeV} - 10 \text{ MeV}$$

Electron motion can become relativistic ( $U_{\text{osc}} > m_e c^2 = 512 \text{ keV}$ )

## High Energy Density Science

### Concentrated energy

Energy density in a femtosecond pulse is  $10^{11} \text{ J/cm}^3$

Corresponds to  $\sim 1 \text{ MeV}$  per atom at solid density

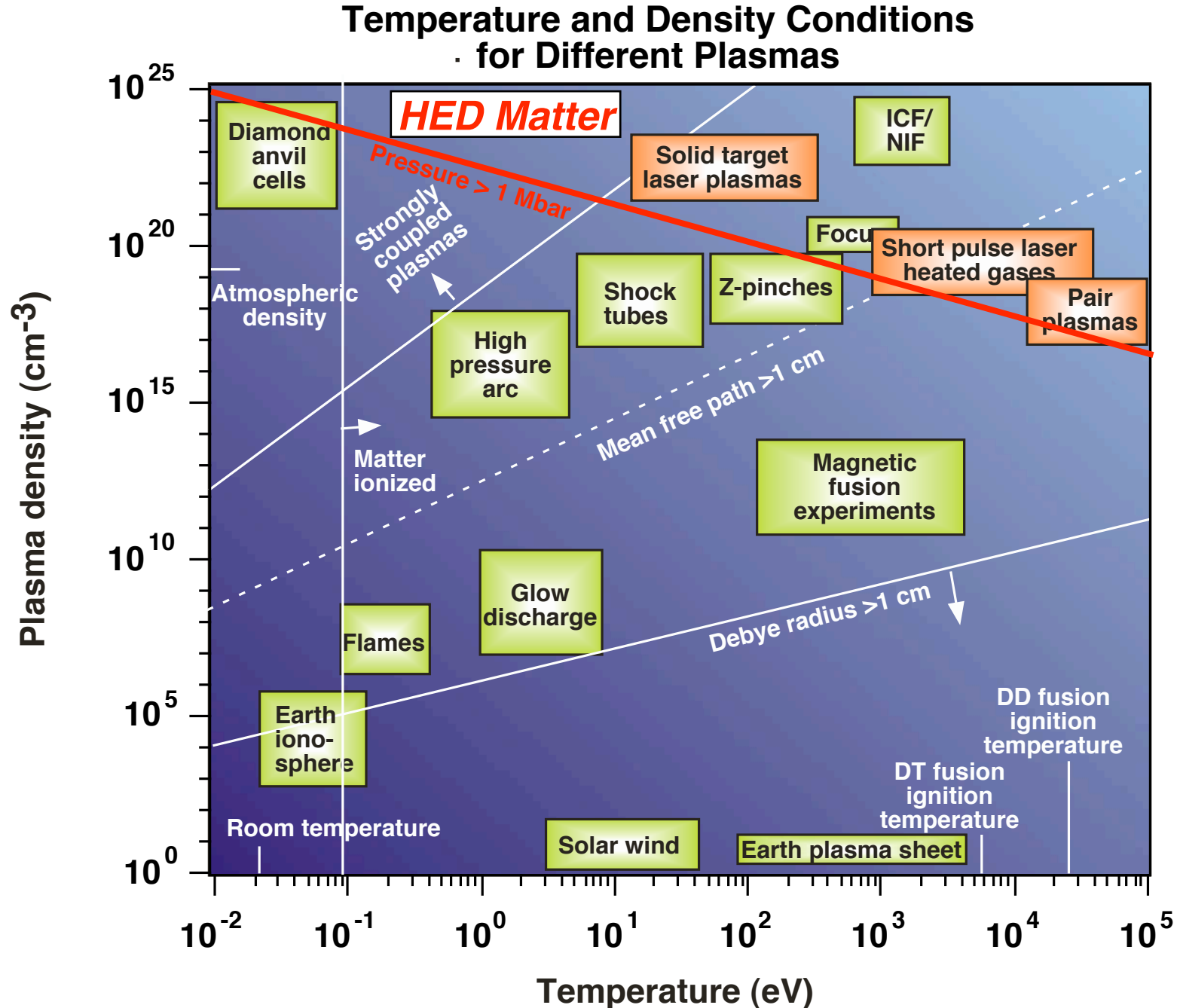
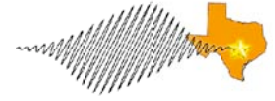
### High brightness and pressure

Radiance exceeds that of a 1 MeV black body

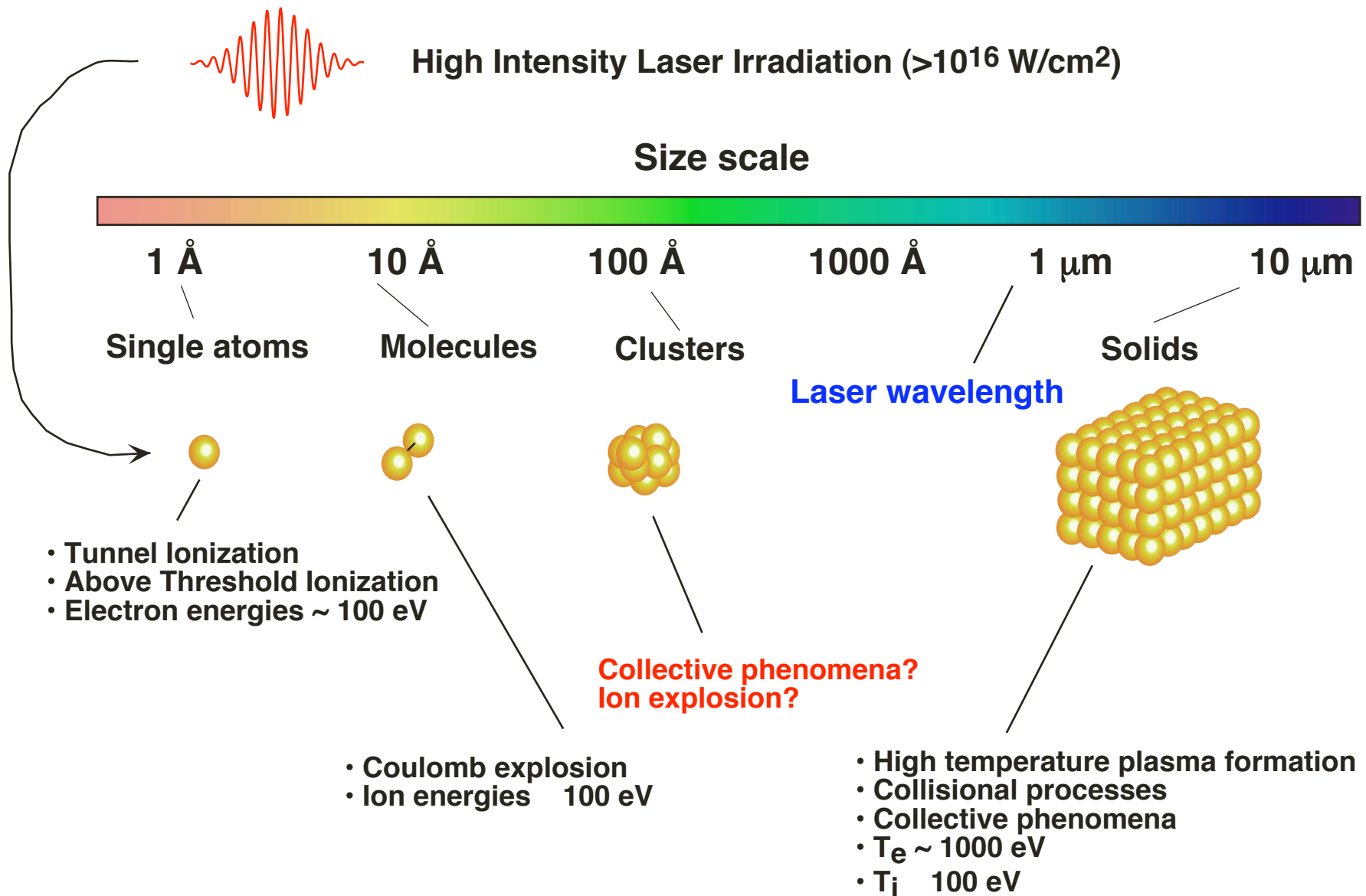
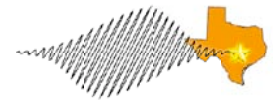
Light pressure  $P = I/c = 30 - 3000 \text{ Gbar}$



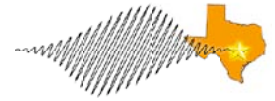
# High energy density matter is created by heating dense plasma to very high temperature



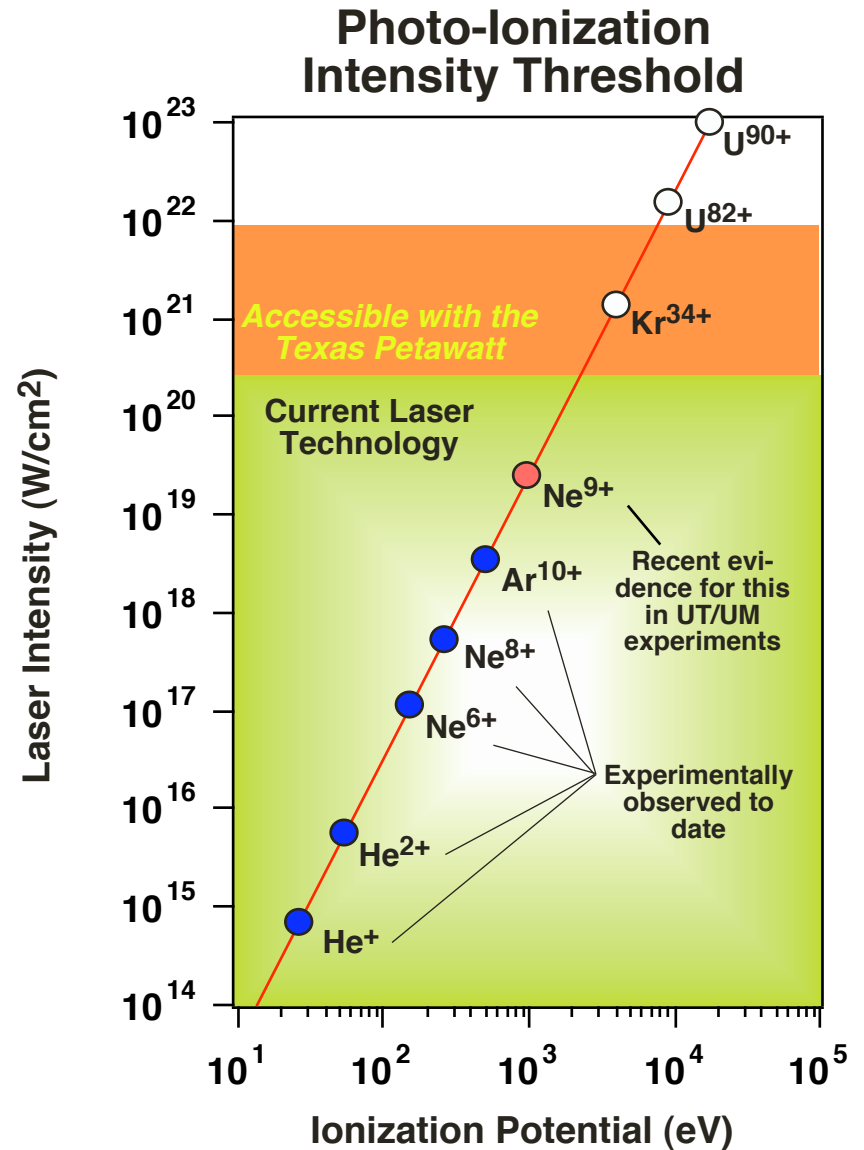
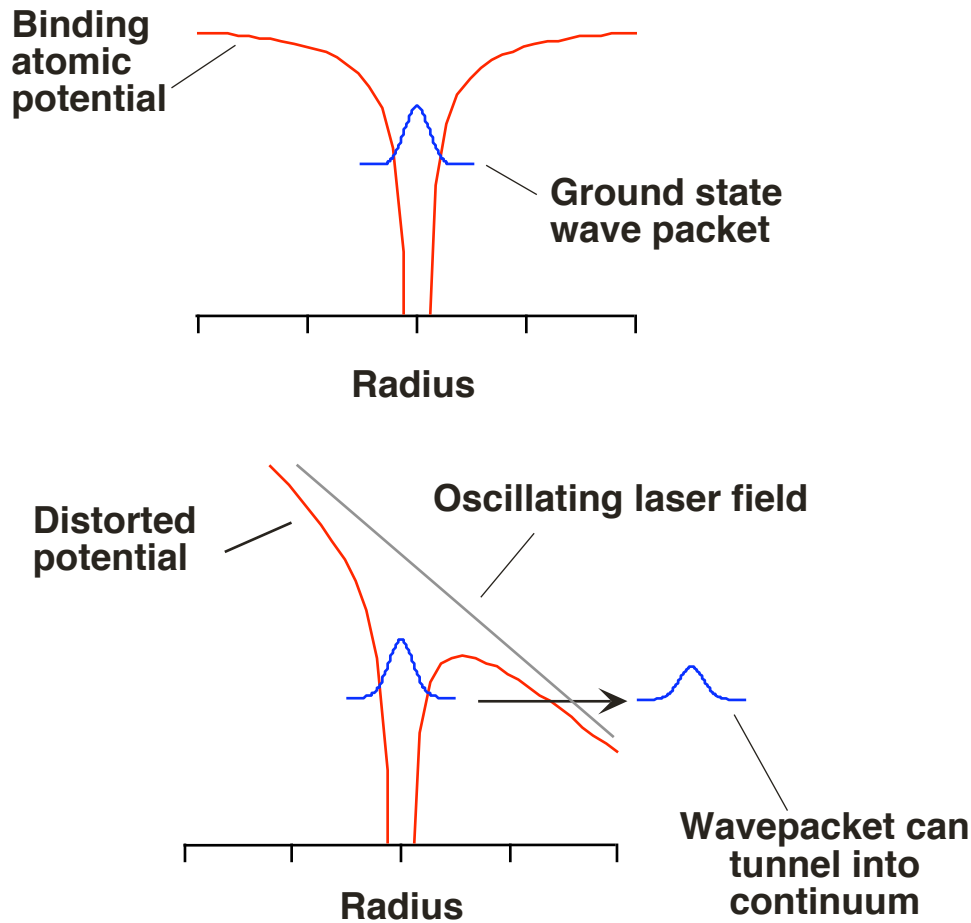
# The interaction of intense laser pulses with different targets leads to quite different physical effects



# Highly non-perturbative laser intensities push photo-ionization into the tunneling regime



A strong laser field can ionize an atom by tunneling





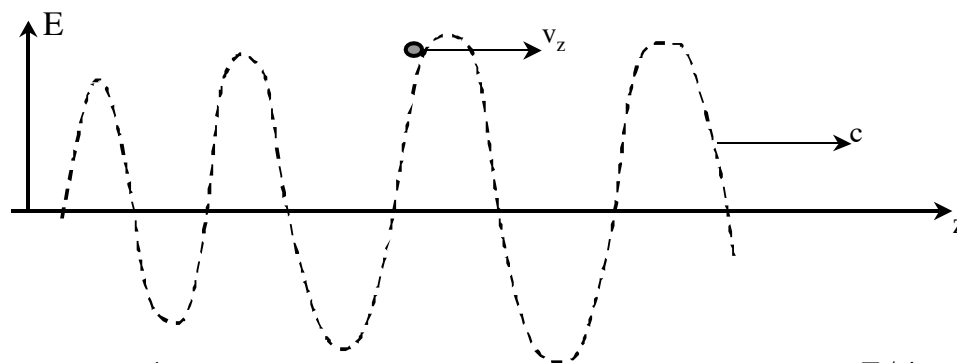
# Ionization in a strongly relativistic beam will be accompanied by strong, free-wave acceleration of the ejected electrons



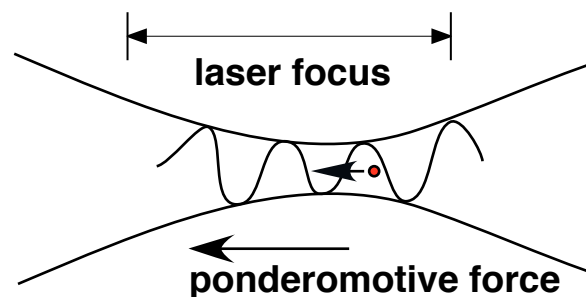
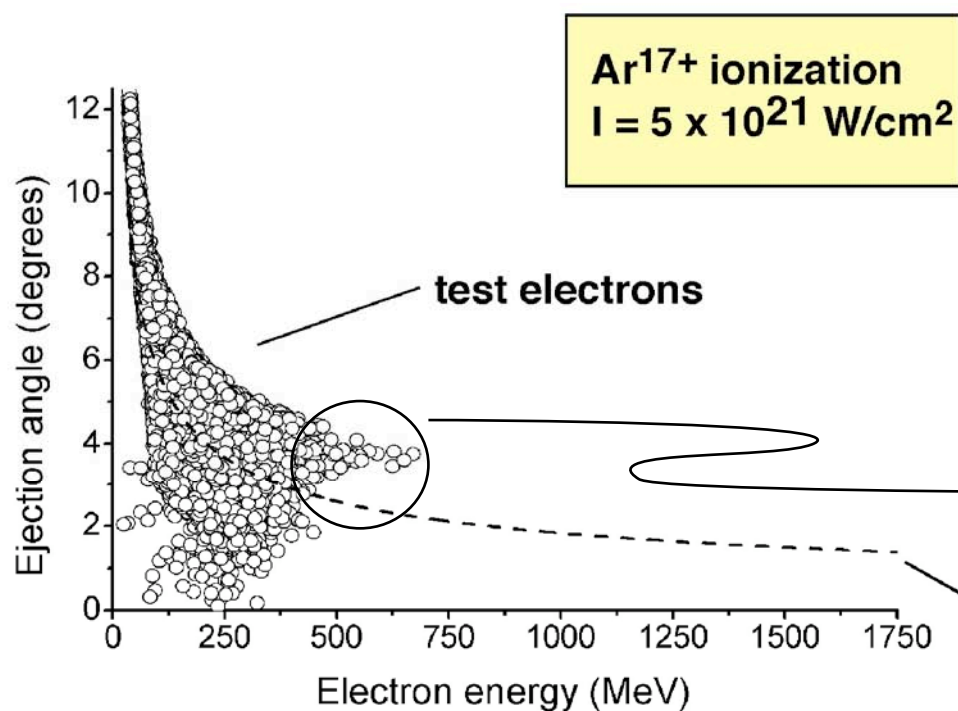
- Ionization "injects" electrons right at the peak of the field

- This yields much greater energies than if the laser interacts with free electrons

Electrons with large longitudinal momentum can "ride" the laser wave



Monte Carlo simulation using a focused PW beam



$$\gamma_{\max} \approx eE_{\max} \theta z_R / 2m_e c^2$$

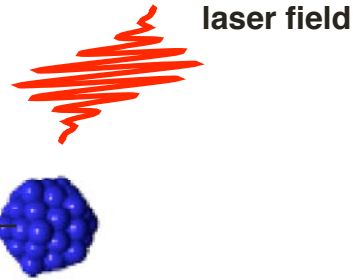
Near GeV electrons ejected within 4° of the laser axis

$$\tan^2 \theta = 2/(\gamma-1)$$

# A cluster irradiated by an intense fs laser, violently explodes, ejecting ions with high kinetic energy



55 atom argon cluster  
before irradiation

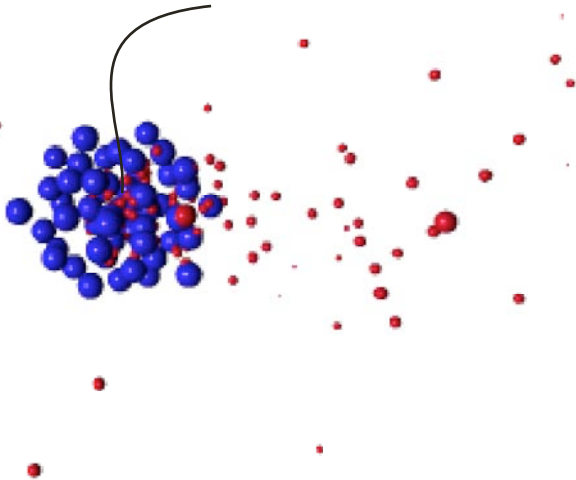


Simulation with  $I = 10^{16} \text{ W/cm}^2$   
 $\Delta\tau = 50 \text{ fs}$

1) time = - 100 fs

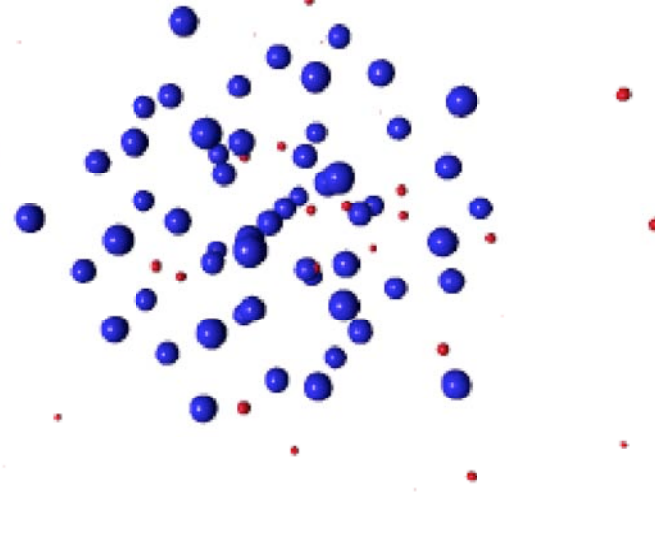
2) time = - 20 fs

electrons confined by  
space charge forces



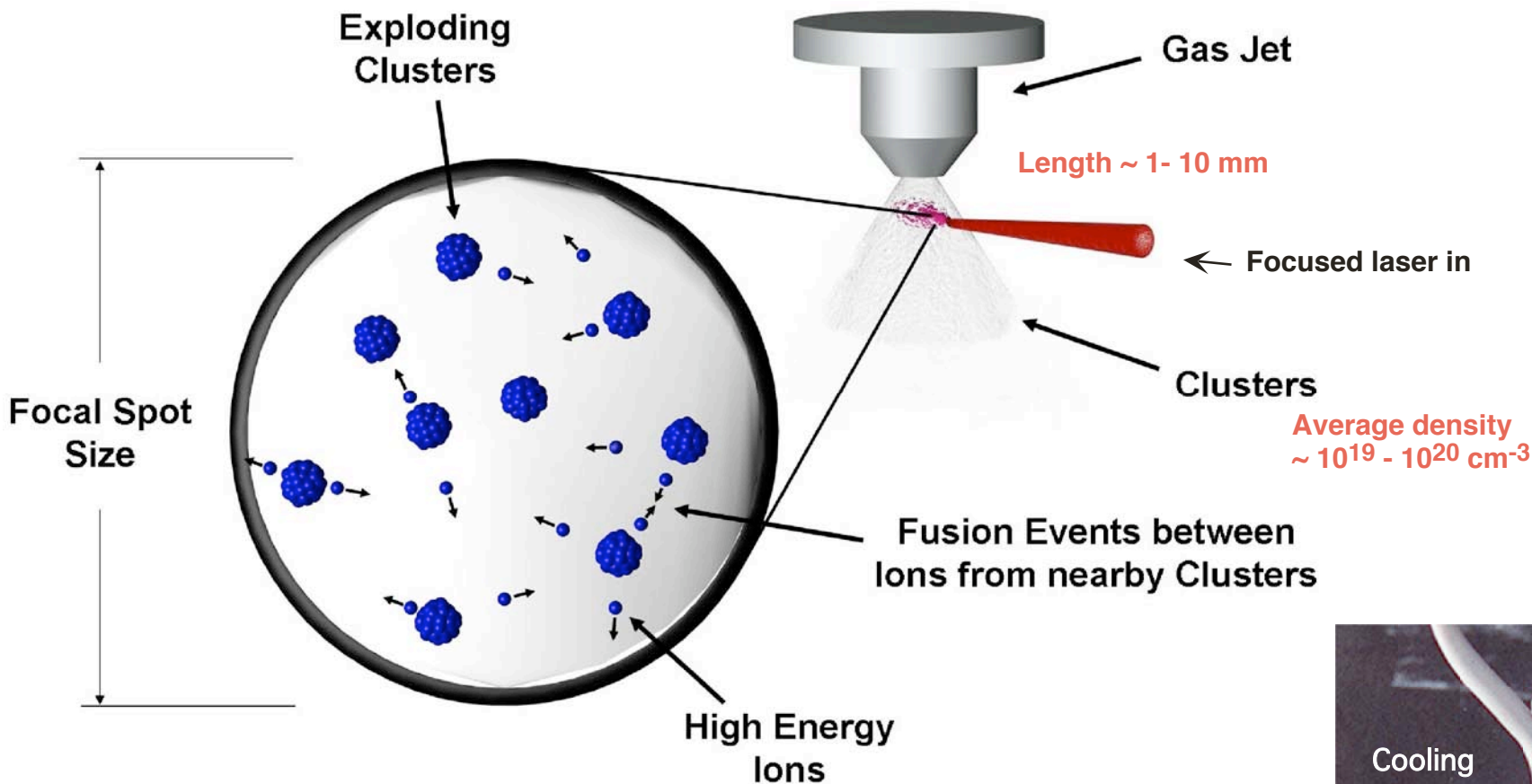
Laser field begins to heat and  
expel electrons from cluster

3) time = + 20 fs

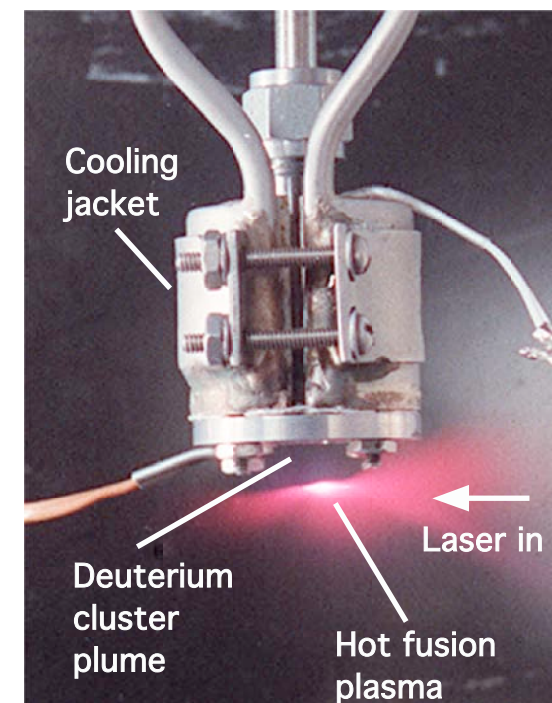


Ions explode by Coulomb forces

# A gas of exploding deuterated clusters can produce a burst of fusion

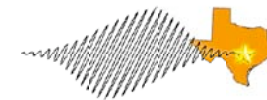


Relevant fusion reactions:

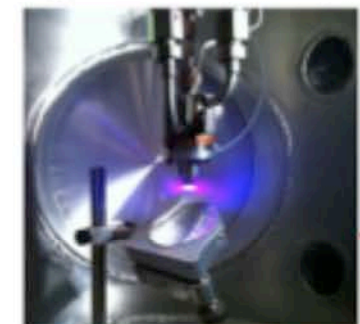




# We have produced $\sim 2 \times 10^7$ DD fusion n/shot in clusters with the Texas Petawatt Laser ( $\sim 100$ J energy)



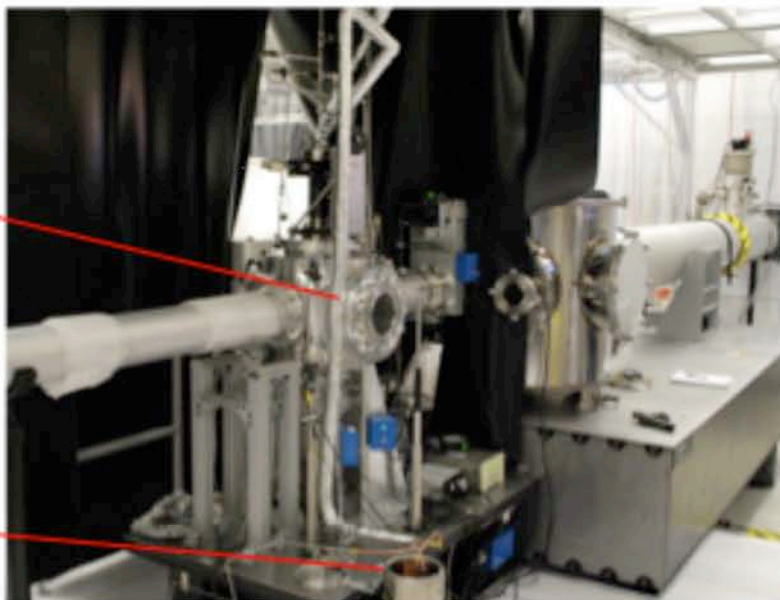
## Target area of the Texas Petawatt for the cluster fusion experiment



Deuterium plasma (OPA shot w/ DFM)

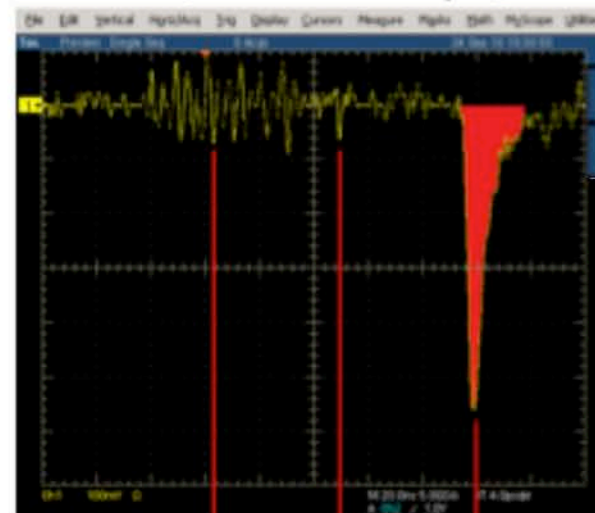


LN2 cooling line



Target chamber

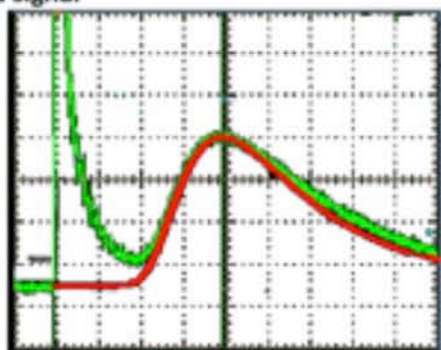
Shot 1282, deuterium 750psi,  $-186^\circ\text{C}$



EMP X-ray neutron

## Faraday cup data showed $kT=2\text{keV}$ deuterium ion peaks on 10J rod shots.

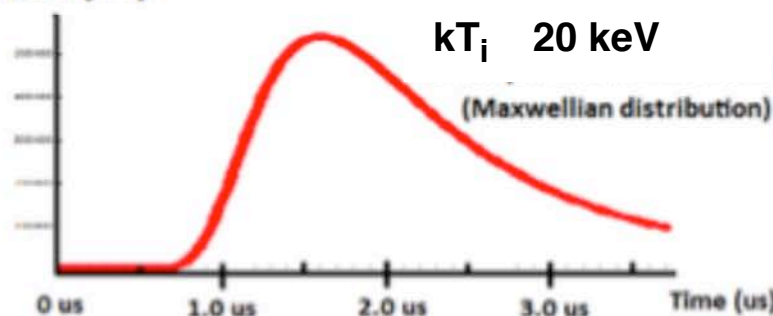
Faraday cup signal (20mV/div)



Oscilloscope trace  
Maxwellian Fit

Time (400ns/div)

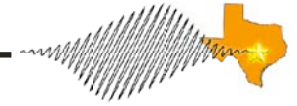
Ion Count (arb.)



$kT_i$  20 keV

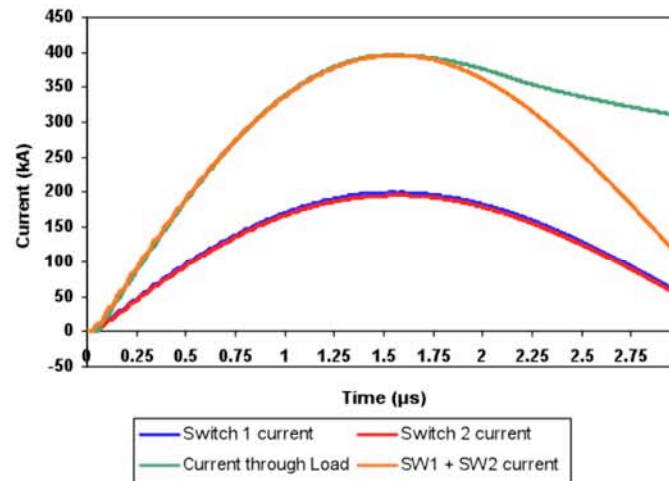
(Maxwellian distribution)

# A 1 MA pulsed power source constructed at Sandia will be installed on the TPW in February



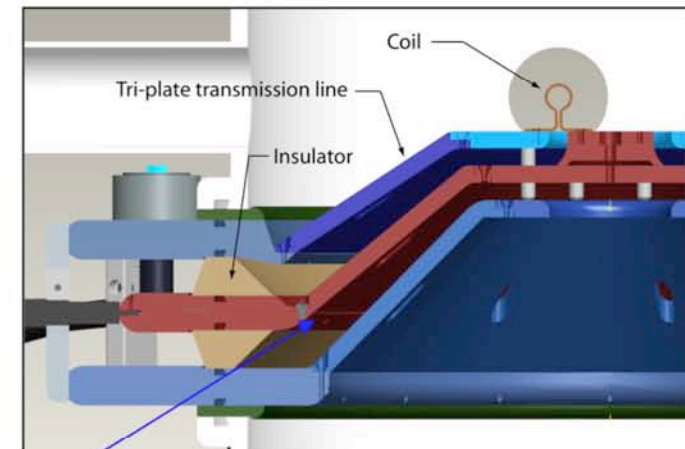
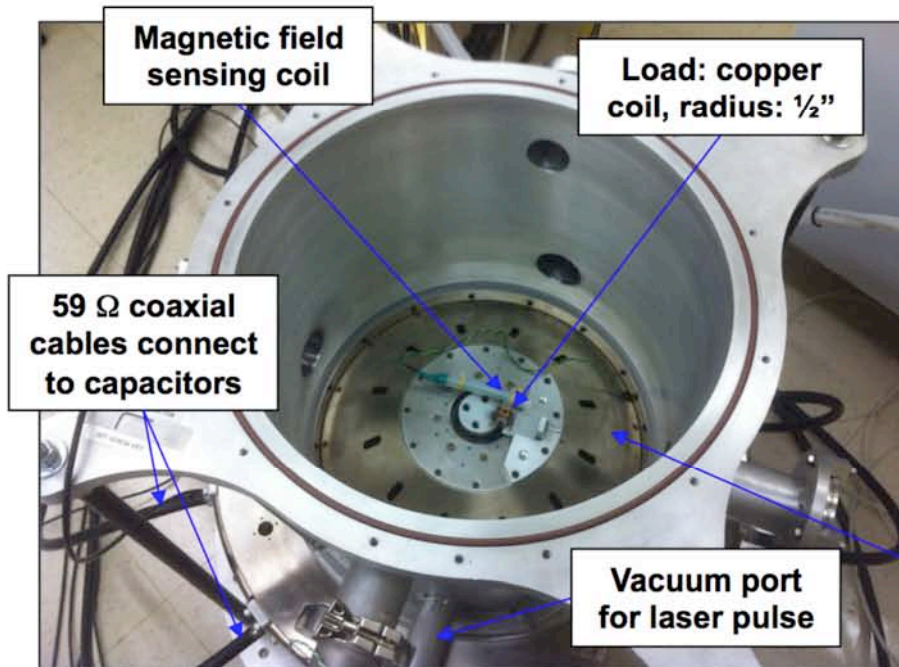
## Recent Results with 1/2 MA Test Chamber

- 400 kA peak current
- Time to peak ~ 1.5 ms
- Load: SS shorting bar
- Estimated 28 T for present coil geometry



x 4 units to be installed on Texas Petawatt

- Anticipate 150 - 200 T



Conical transmission line, similar to Z Machine



# The 100 fs TPW will allow flat-field resonant, multi-GeV laser wakefield acceleration of electrons

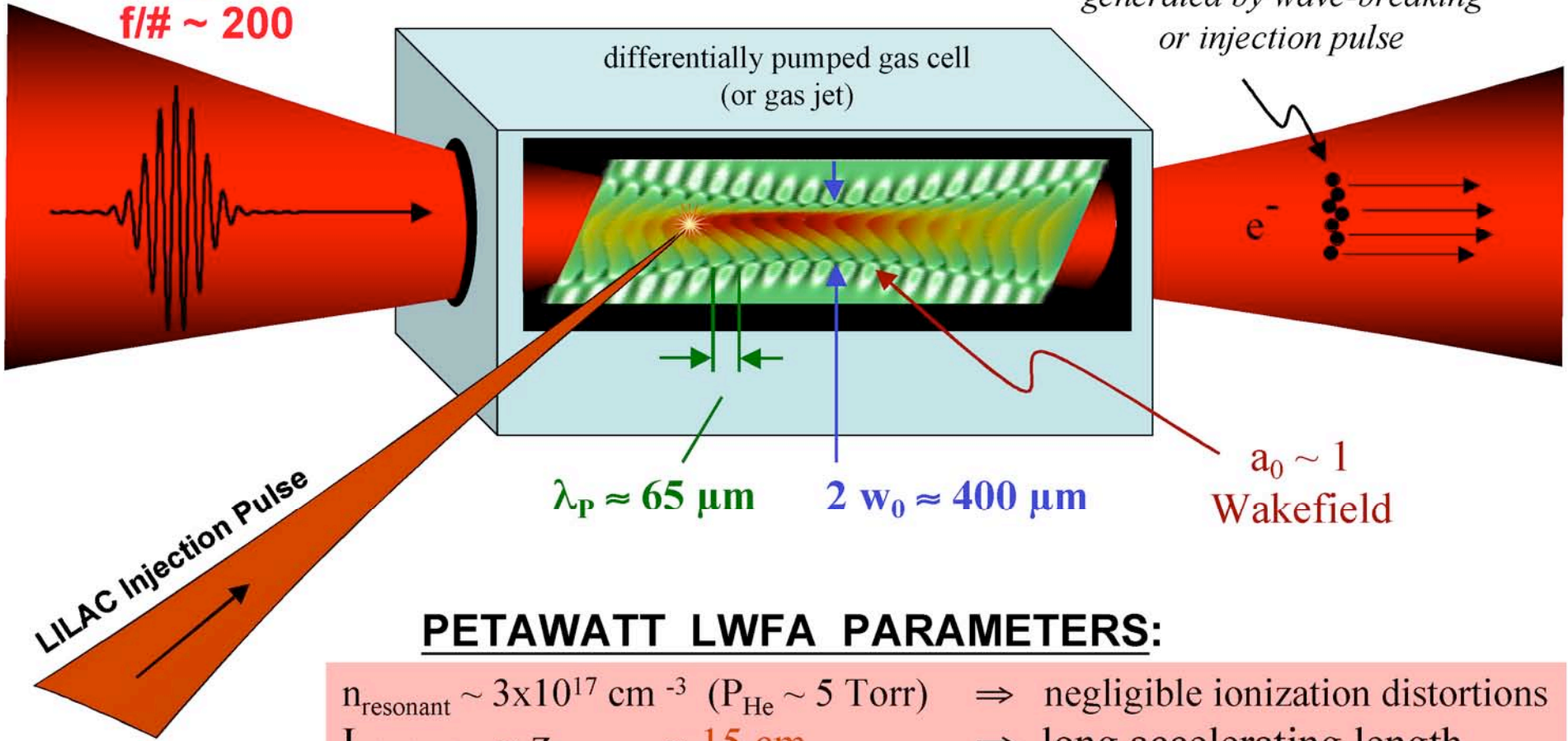


Petawatt Pulse

100 J

150 fs

$f/\# \sim 200$



Relativistic electrons  
generated by wave-breaking  
or injection pulse

differentially pumped gas cell  
(or gas jet)

$\lambda_p \approx 65 \mu\text{m}$

$2 w_0 \approx 400 \mu\text{m}$

$a_0 \sim 1$   
Wakefield

## PETAWATT LWFA PARAMETERS:

- $n_{\text{resonant}} \sim 3 \times 10^{17} \text{ cm}^{-3}$  ( $P_{\text{He}} \sim 5 \text{ Torr}$ )  $\Rightarrow$  negligible ionization distortions
- $L_{\text{dephasing}} \sim Z_{\text{Rayleigh}} \sim 15 \text{ cm}$   $\Rightarrow$  long accelerating length
- $E_z \sim 0.5 \text{ GV/cm}$   $\Rightarrow$  large accelerating field

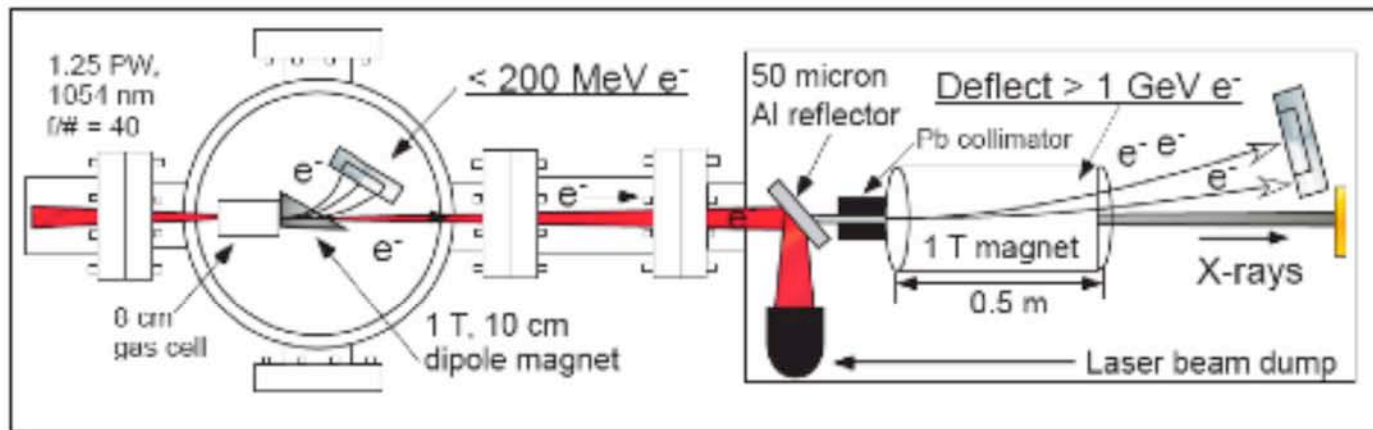
Umstadter, PRL ('96)

Linear wakefield best  
driven by fs pulses

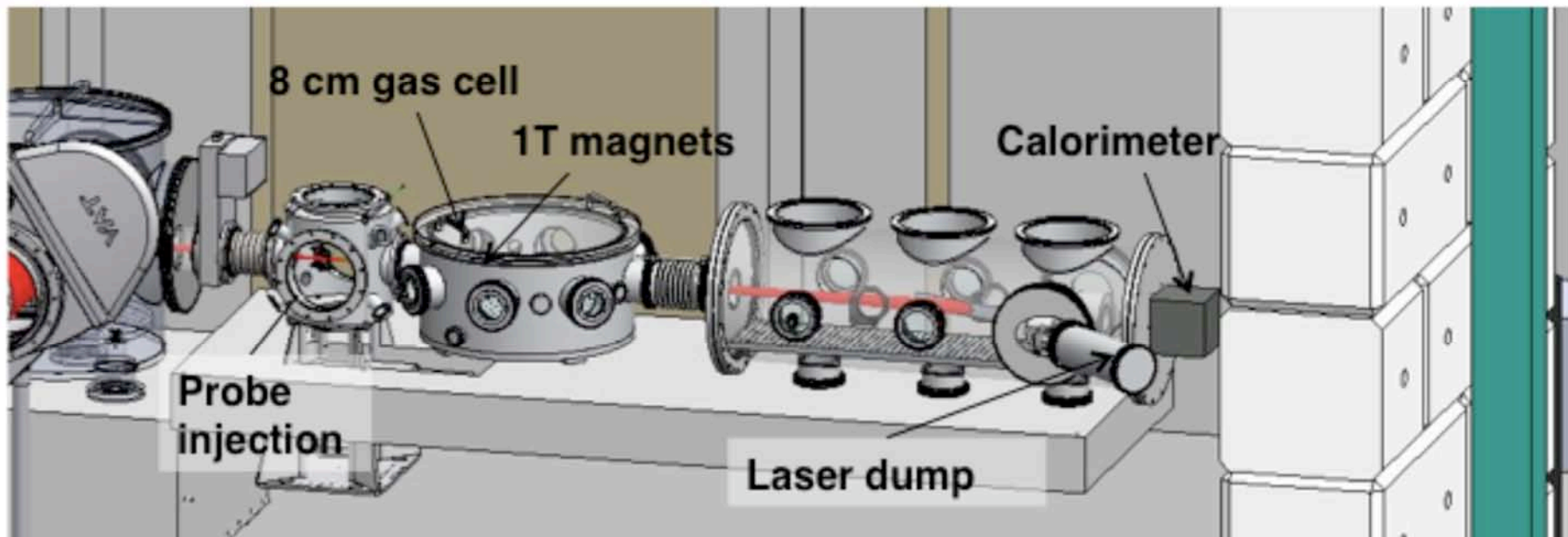
Accelerating gradient  $\sim n_e^{1/2} \longrightarrow n_e \sim (1/\Delta t_{\text{pulse}})^2$



# Electron acceleration on the TPW will be explored at electron energies up to the 10 GeV level

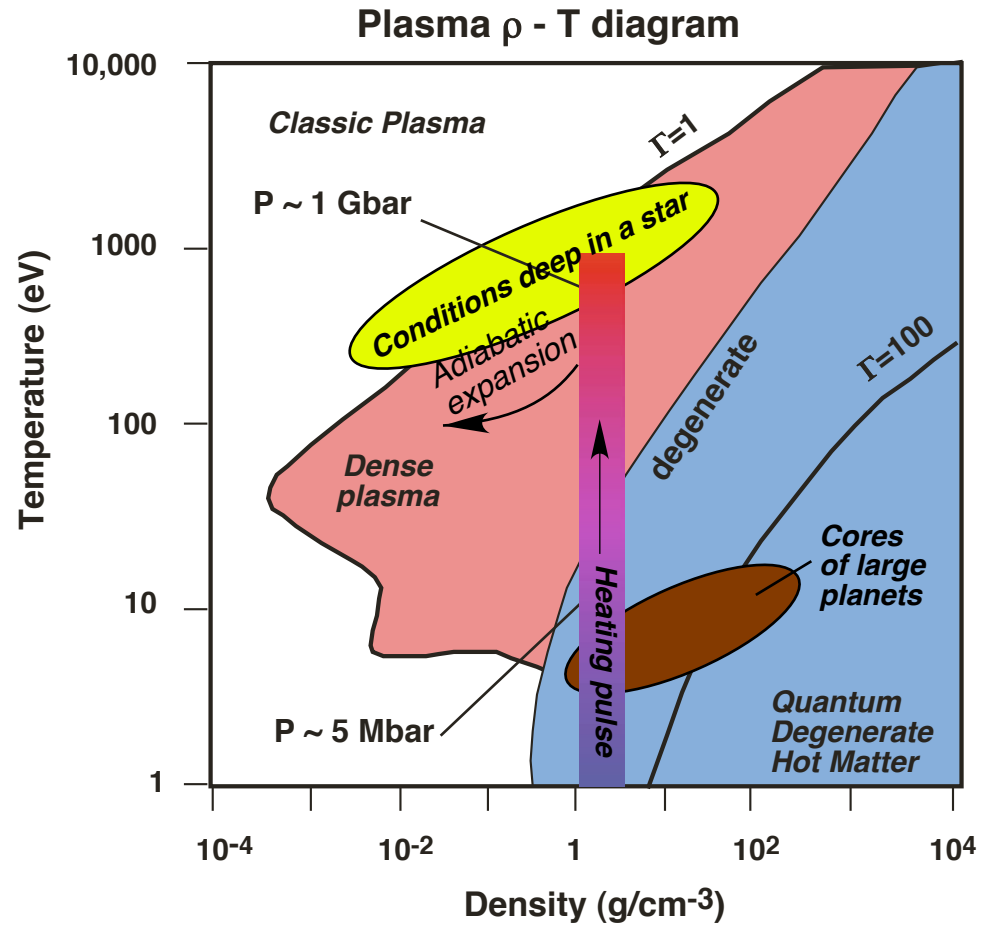
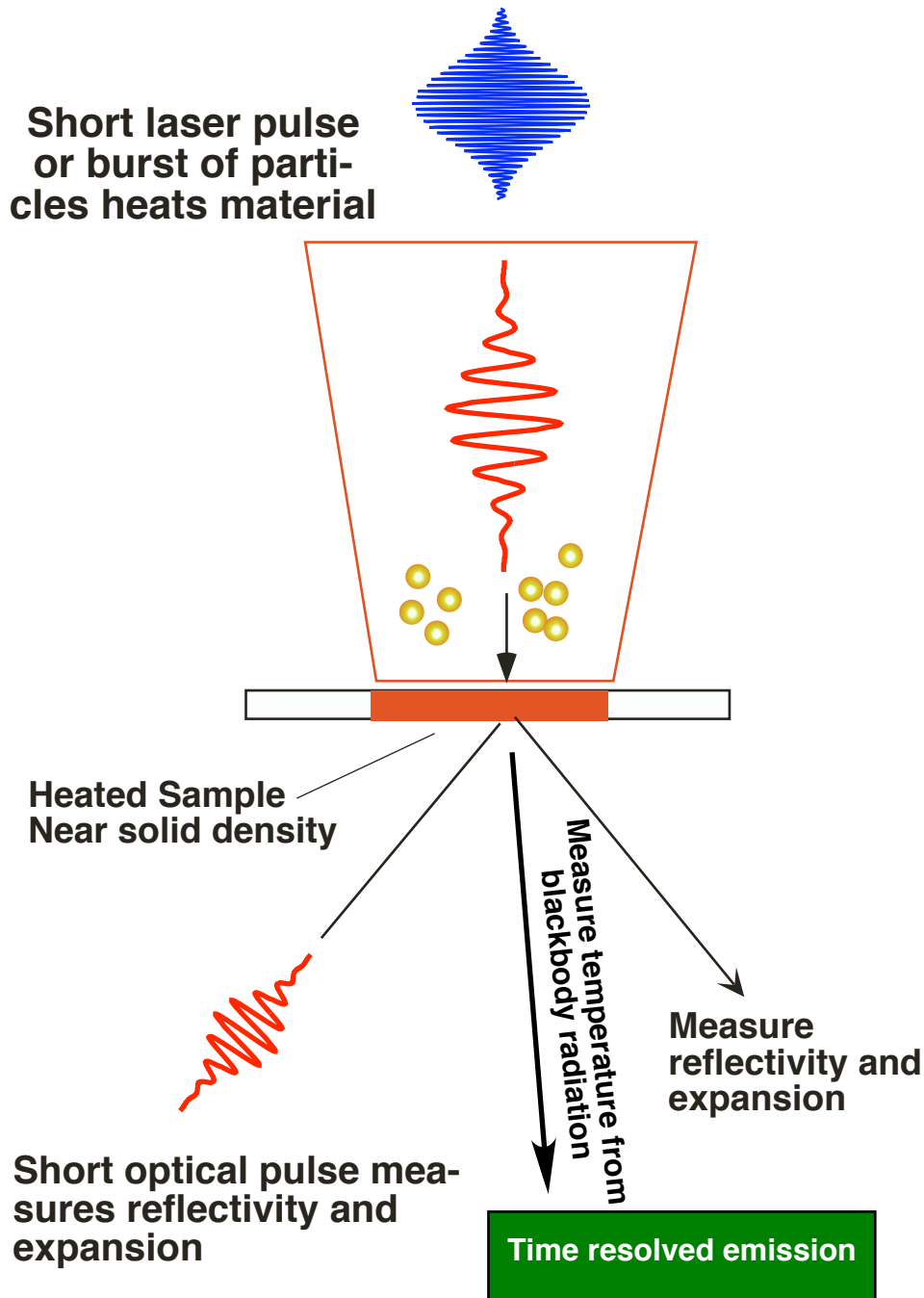


- $> 1 \text{ GeV}$  electrons predicted
- Frequency domain holography visualization
- Challenge of 1/hr shots

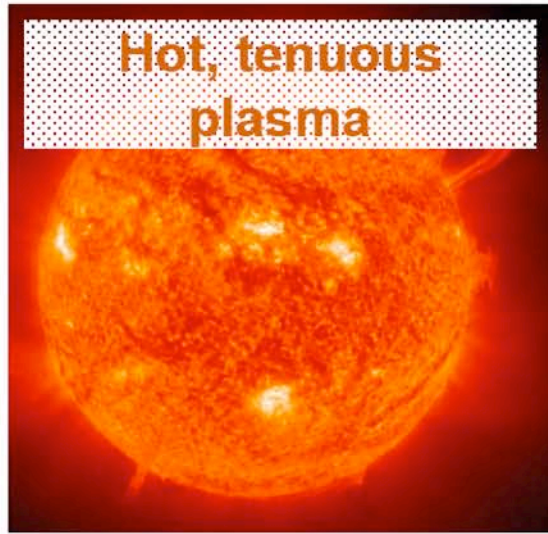
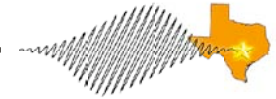


These chambers will be useful for future experiments

# Isochoric heating can be combined with optical and x-ray probes to derive information about a hot dense plasma



# A quantitative understanding of these HED plasma physics issues is of considerable practical importance

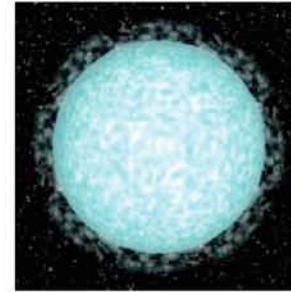


Hot, tenuous plasma

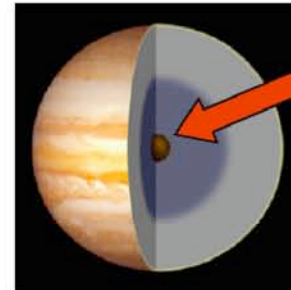
Hot dense matter



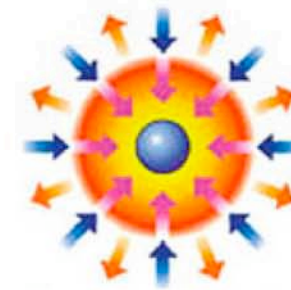
Condensed matter



White dwarfs



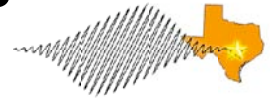
Cores of giant planets and brown dwarfs



Fusion fuel capsule compression

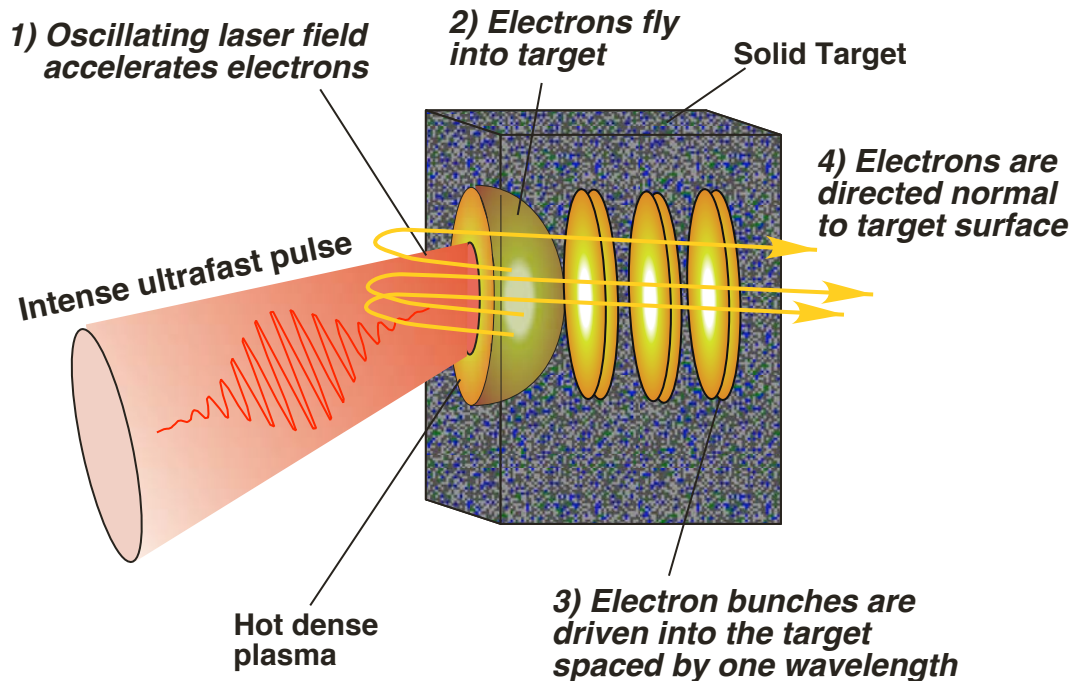
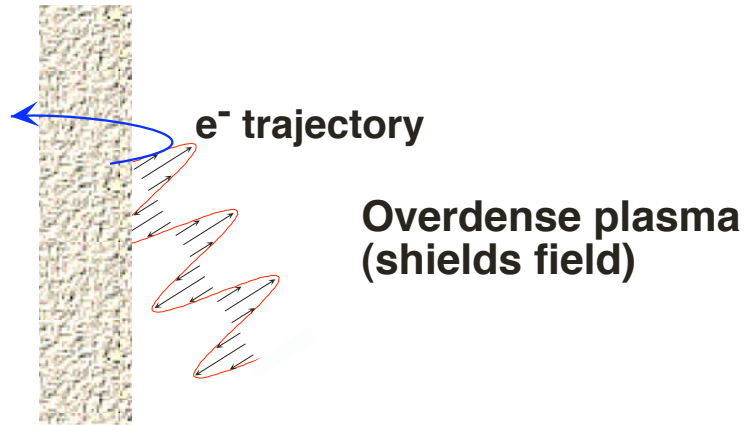


# “Rescattering” of laser driven electrons at the surface of a sharp plasma gradient can produce pulses of fast electrons

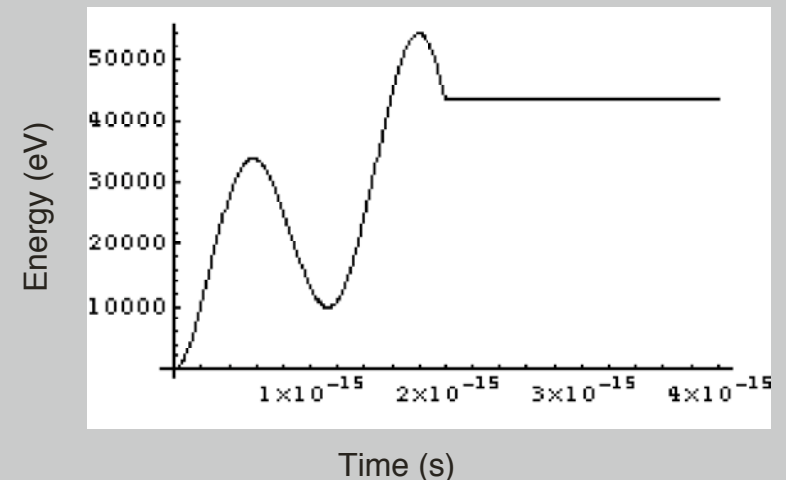
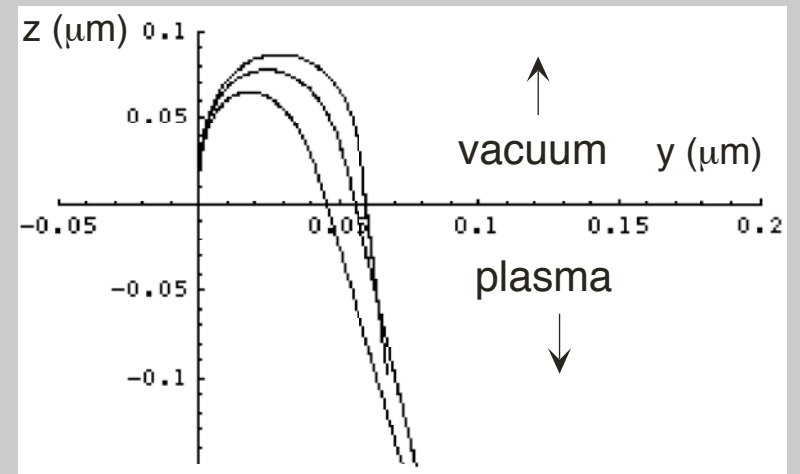


Scattering at the conducting surface breaks the adiabaticity of the laser oscillation

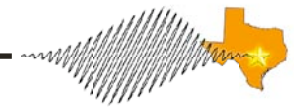
→ **"Brunel absorption"**



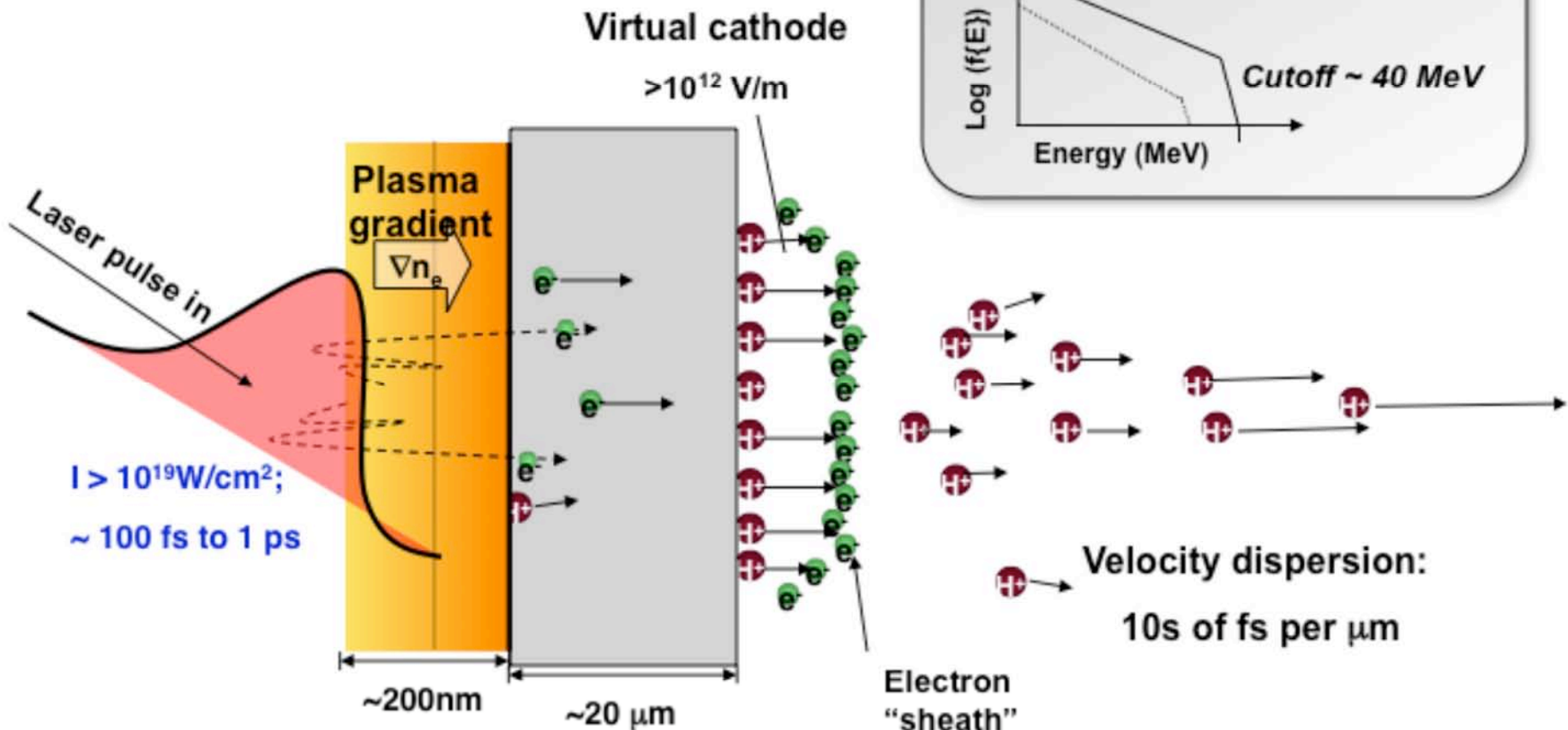
Simulated e<sup>-</sup> trajectories and energy in relativistic field at plasma surface  
 $U_p = 50 \text{ keV} (5 \times 10^{17} \text{ W/cm}^2)$



# With a petawatt laser, very intense, energetic pulses of protons can be produced



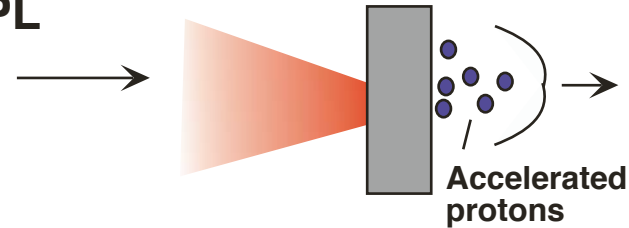
Protons are present on metal foil surfaces  
in the form of  $H_2O$  and hydrocarbons  
(moderate vacuum  $\sim 10^{-5}$  Torr)



# Target normal sheath acceleration of protons to >10 MeV was observed first at the LLNL Petawatt Laser

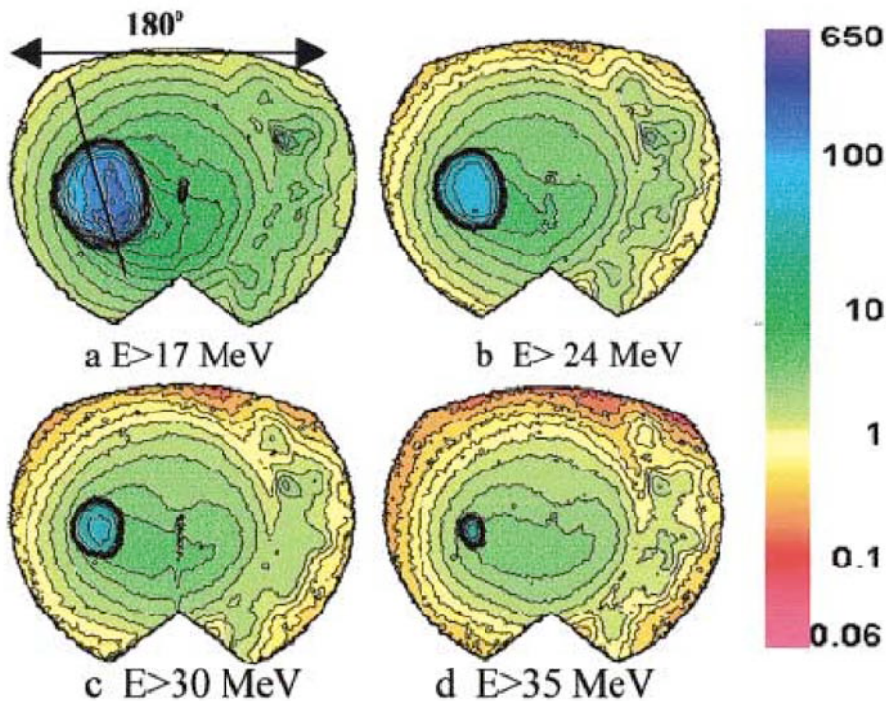


Relativistic electrons produced by a SPL can accelerate protons from a solid

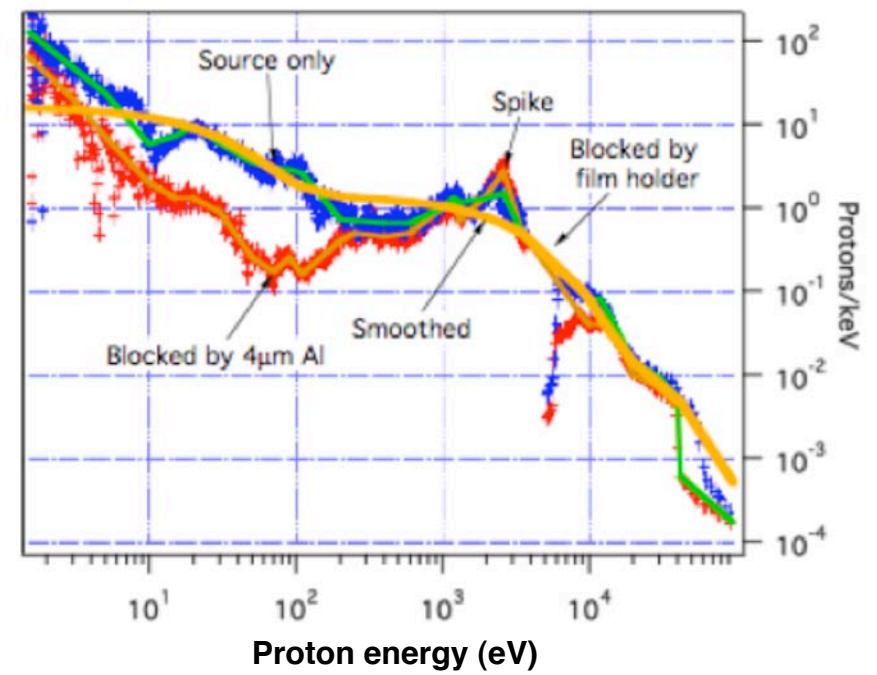


At  $3 \times 10^{20}$  W/cm<sup>2</sup> on LLNL PW laser:

Up to 48J of protons (12% of laser energy) were observed in protons with energy >10 MeV

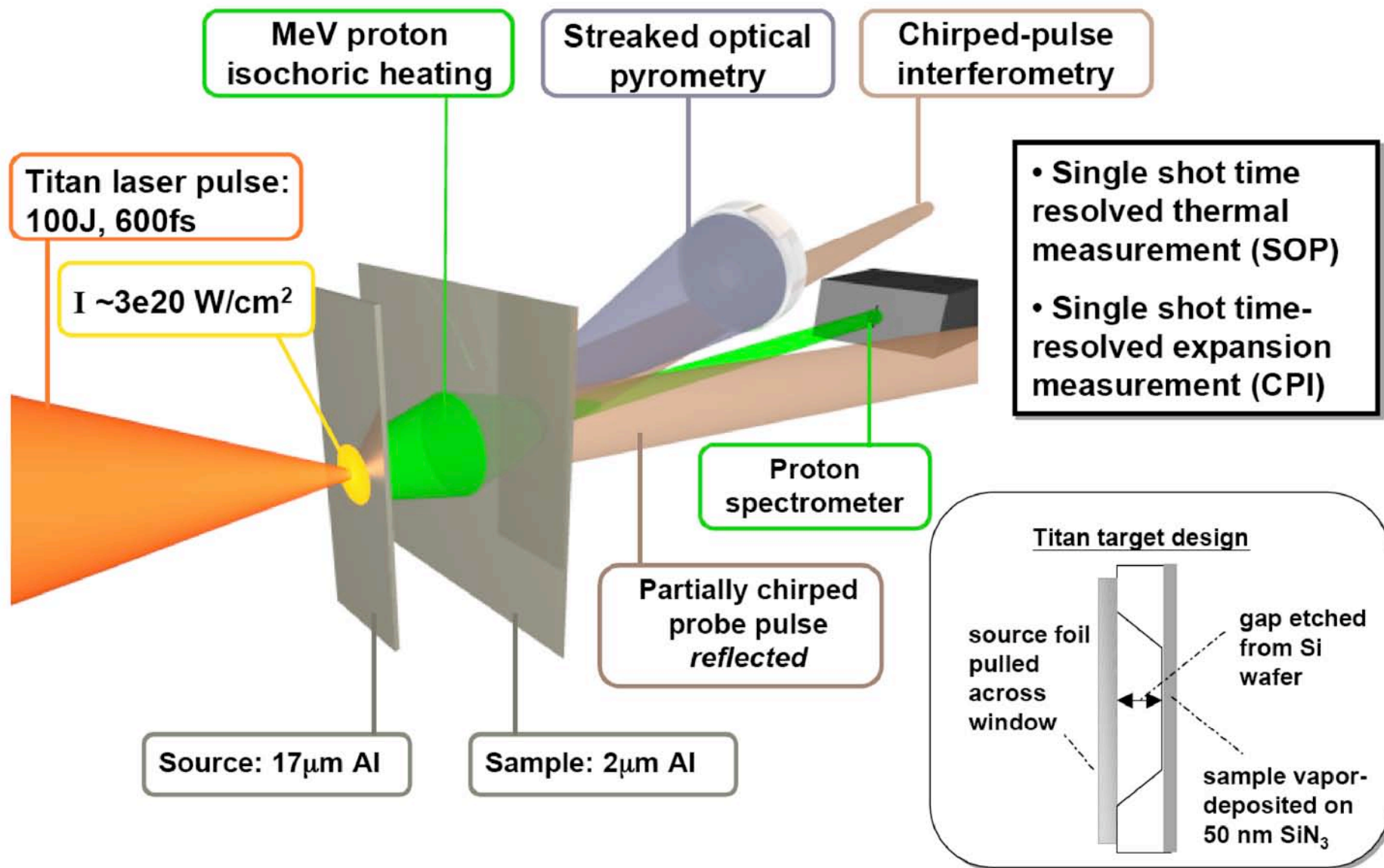


Proton spectrum taken on the Titan 250 TW laser

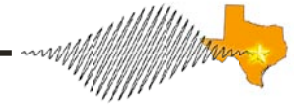




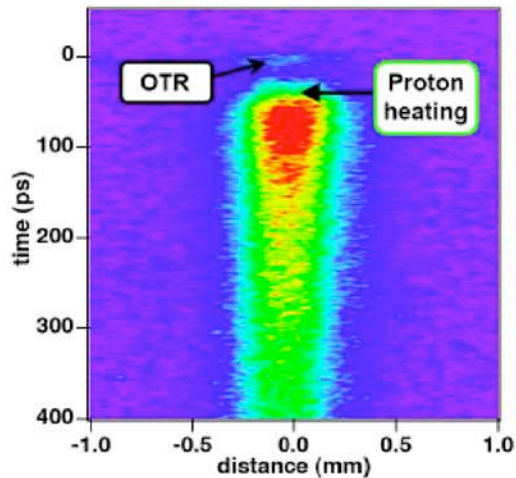
# We have demonstrated proton heating using the LLNL 200 TW Titan laser



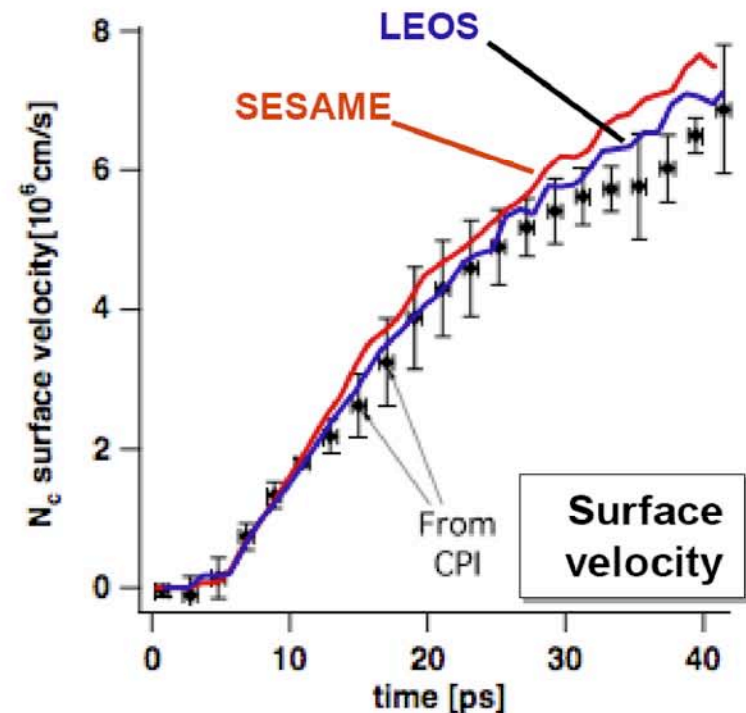
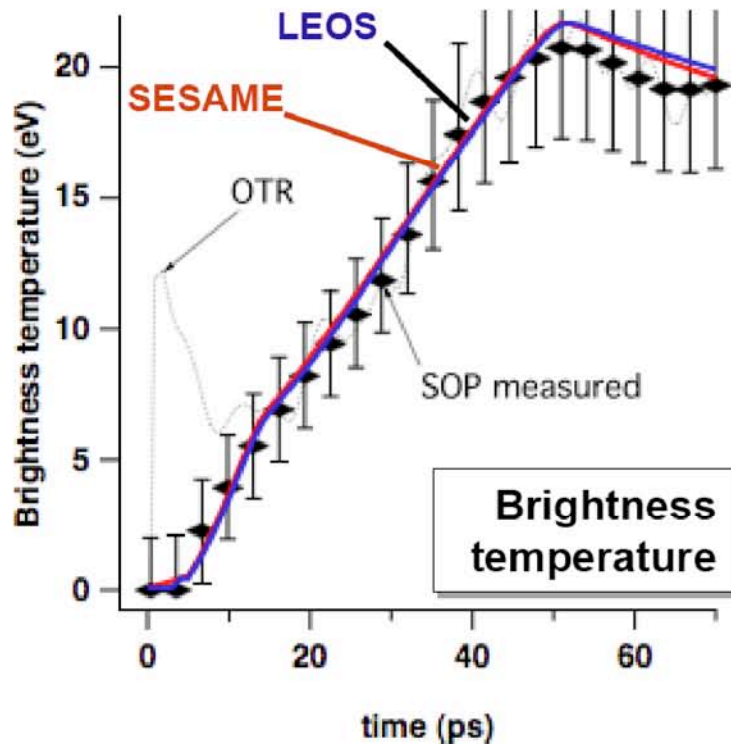
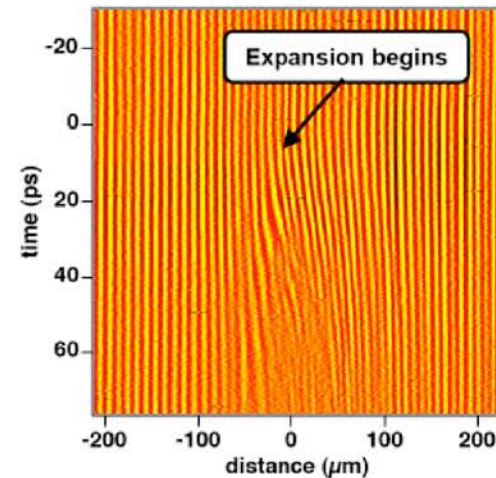
# Using hydrodynamic simulations we are able to evaluate various EOS tables for Al in the 10 - 20 eV region



SOP: Time-resolved temperature

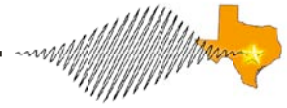


CPI: Time-resolved expansion

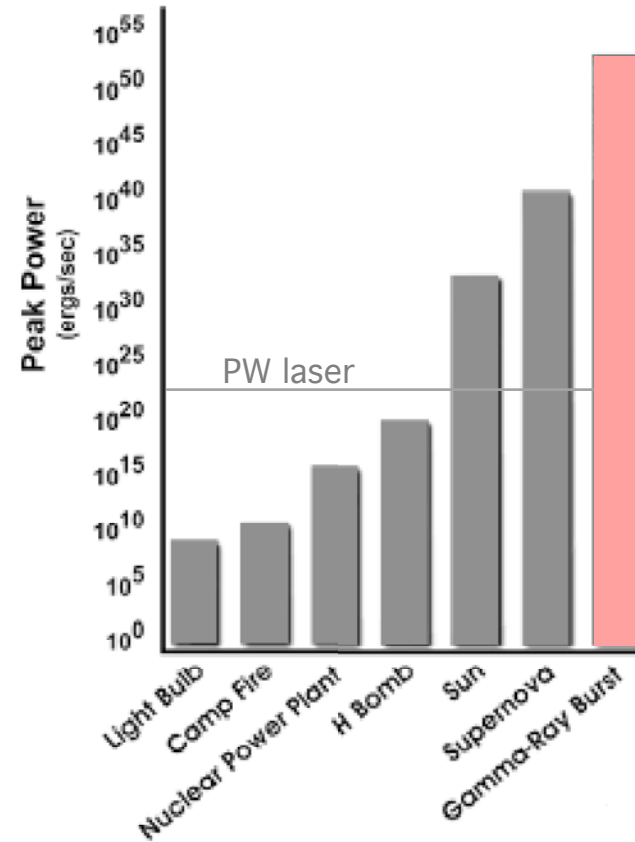
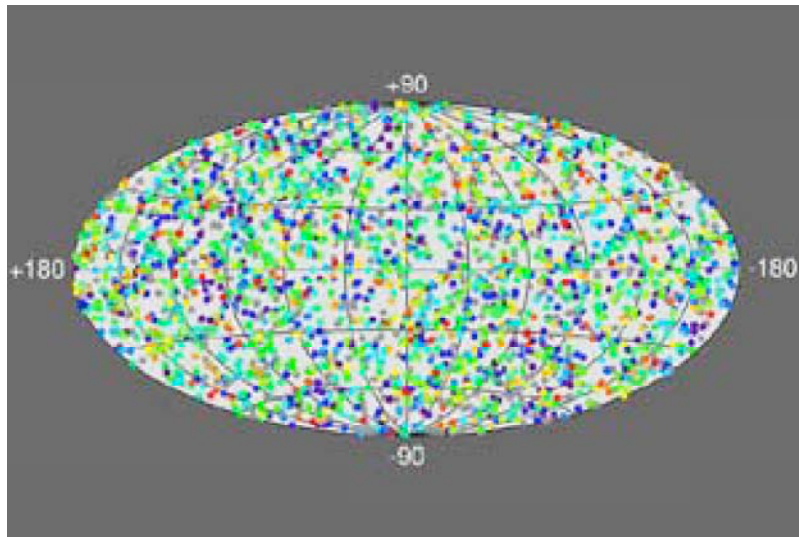


Agreement within 10%

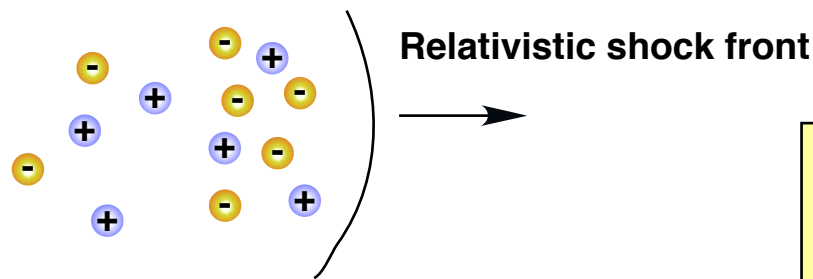
# Gamma Ray Bursts are among the most energetic and enigmatic events in the Universe



The distribution of gamma-ray burst sources on the sky



One possible mechanism is the production of internal shocks in relativistic plasma jets - collisions of electron positron plasmas



*Matter- antimatter plasma  
(electrons and positrons)*

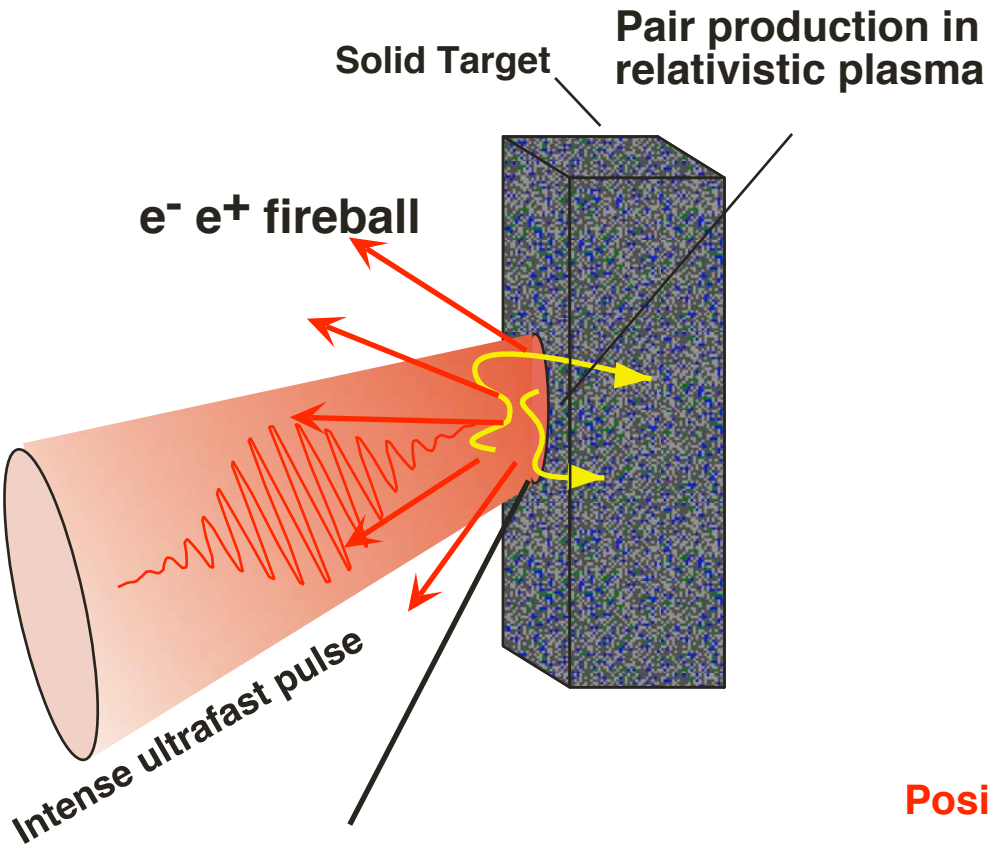
**Pair plasmas may exist near black hole candidates and may play a role in GRBs**



# At sufficiently high intensity, direct positron production by laser driven electrons is possible

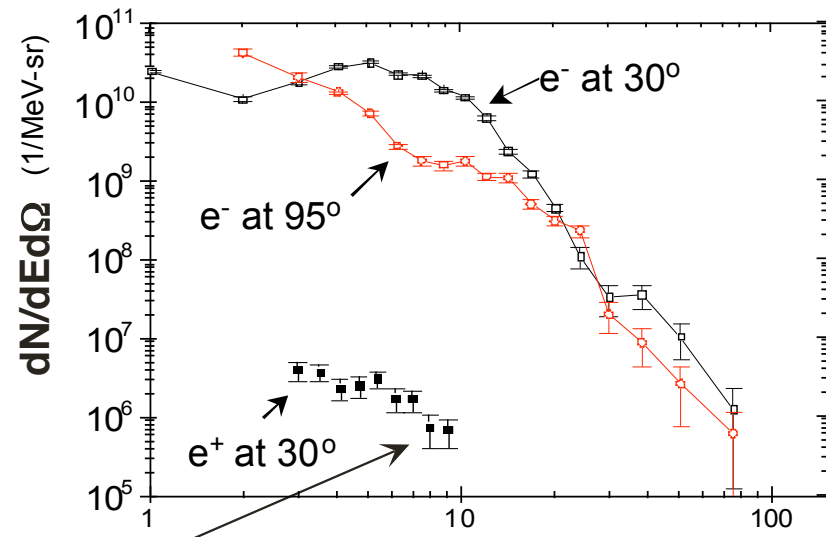


Experiment proposed by Liang et al. PRL 81, 4887 (1998)



Strong ponderomotive forces drive electrons  $\rightarrow T_{\text{hot}} \sim 1 \text{ MeV}$

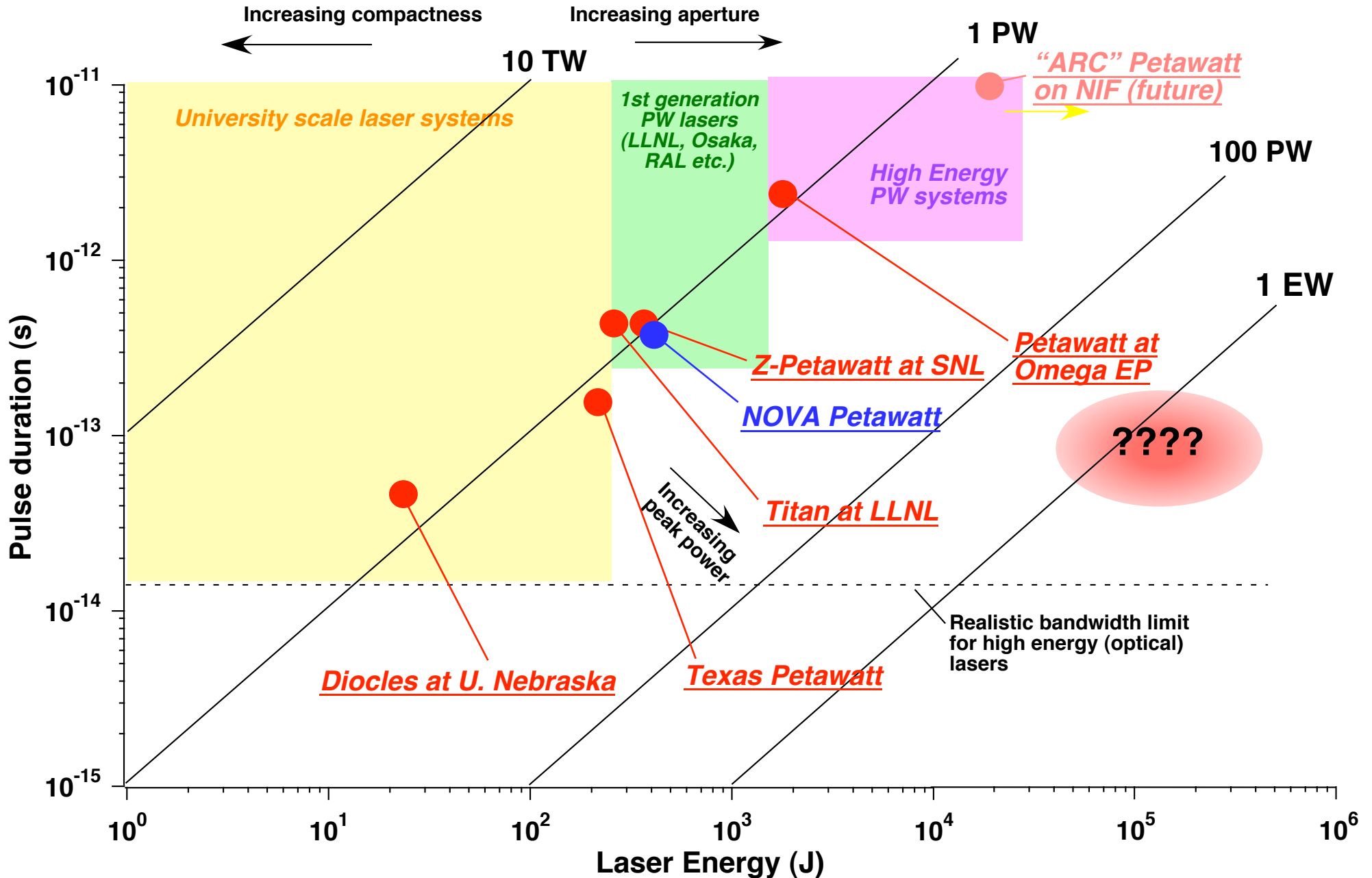
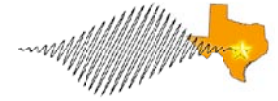
Hot electron and positron spectrum measured from the LLNL PW laser



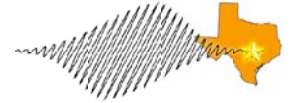
Positron spectrum

Production of a pair plasma results and the explosion of this plasma results in an  $e^- e^+$  fireball  $\rightarrow$  with potential relevance to gamma ray bursts

# How high in laser power can we go using chirped pulse amplification?



# The science enabled by an exawatt laser has yet to be examined in detail



***Focused intensity:  $> 10^{25}$  W/cm<sup>2</sup>***

***Electron quiver energy:  $\sim 10$  GeV***

***Light pressure:  $\sim 10^{15}$  bar***

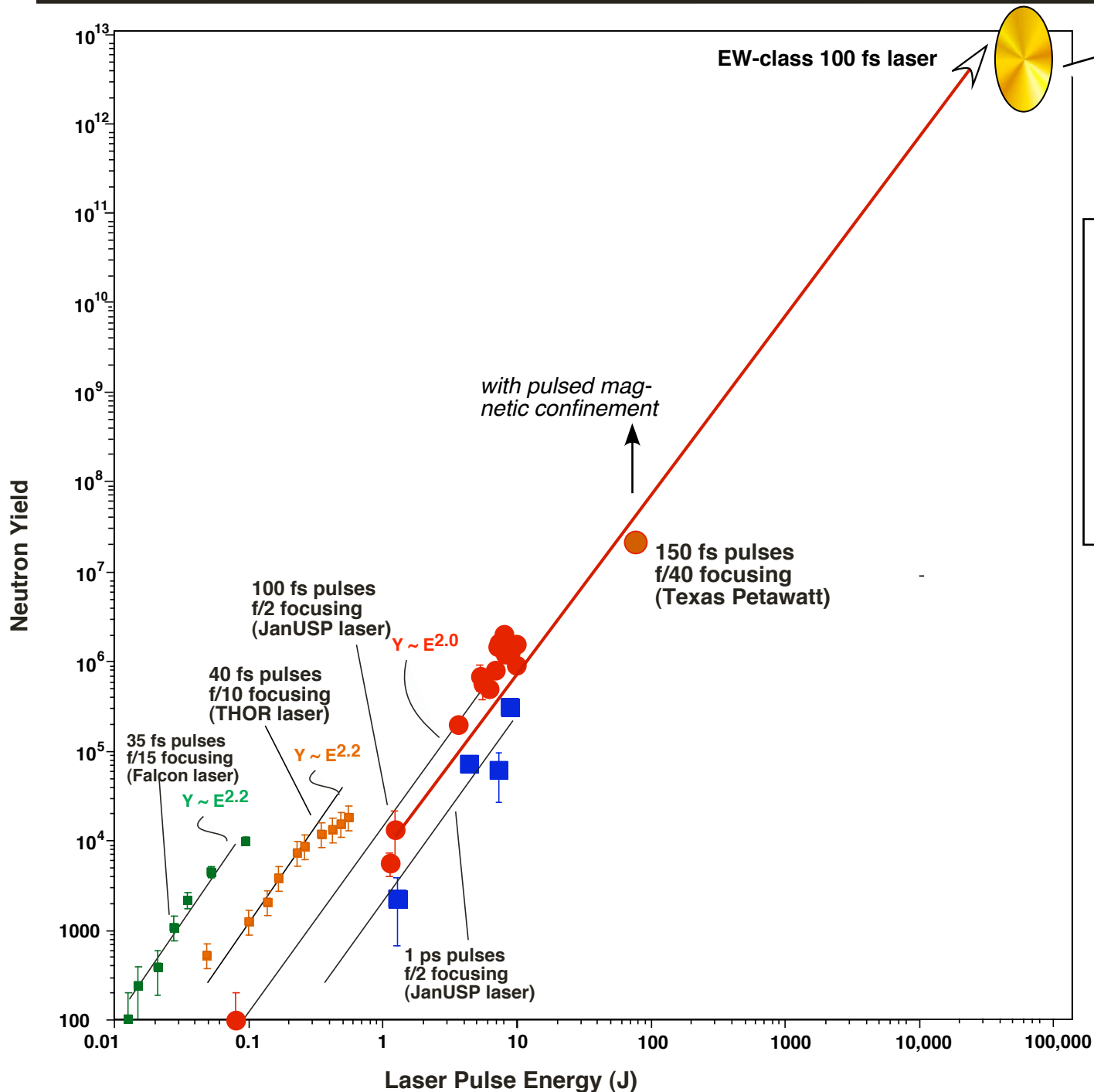
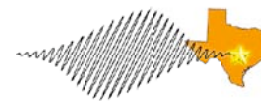
***Proton quiver energy:  $\sim 100$  MeV***



- ***Macroscopic relativistic electron plasmas (pair plasmas)***
- ***Relativistic ion plasmas***
- ***TeV electron acceleration***
- ***Hawking Unruh radiation***
- ***Vacuum birefringence***



# An exawatt-class laser could produce a unique plasma environment for possible nucleosynthesis experiments



## Plasma conditions:

$kT_i$  20 keV

plasma volume 1 cm<sup>3</sup>  
(cylinder d = 5 mm; l = 2 cm)

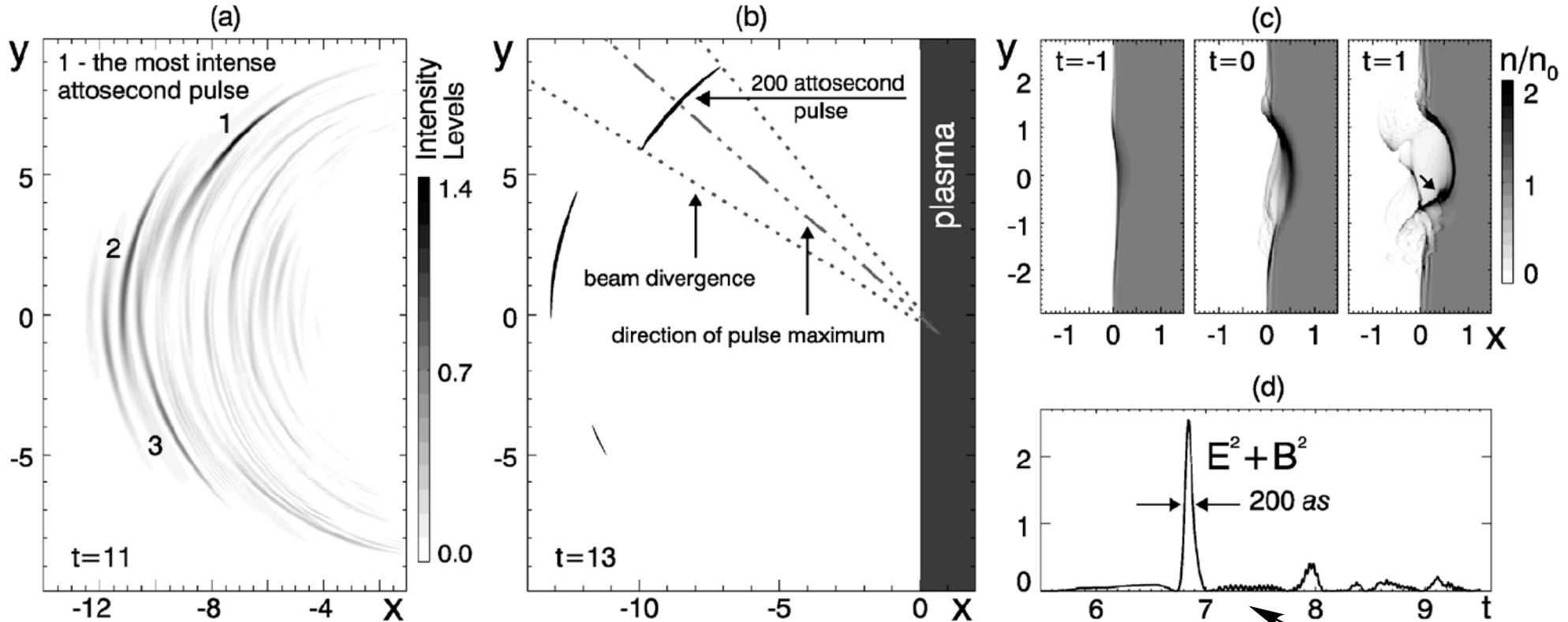
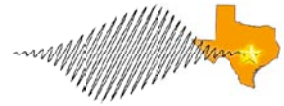
ion density 10<sup>19</sup> - 10<sup>20</sup> cm<sup>-3</sup>

Inertial lifetime of plasma 10 ns  
(possibly longer with addition of pulsed magnetic field)

**Possible environment  
for nucleosynthesis?**

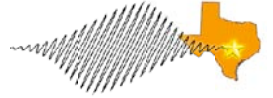
10<sup>7</sup> less density than NIF but...  
10<sup>3</sup> times longer plasma life  
10<sup>2</sup> - 10<sup>3</sup> times greater volume  
~ 10<sup>2</sup> - 10<sup>3</sup> times greater shot rate

# Relativistic motion of macroscopic amounts of matter (plasma) might become possible

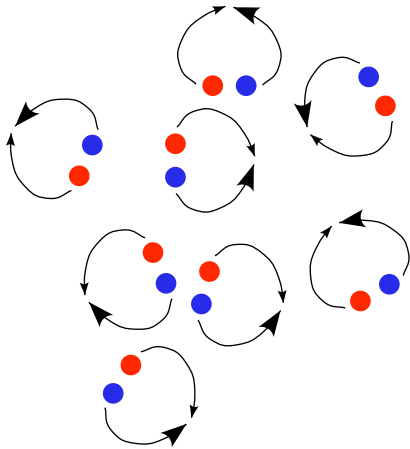


**Attosecond pulses are reflected from the relativistically moving plasma surface**

# A strong field can induce effects in virtual pairs produced in the vacuum



*Cartoon of virtual pairs in the vacuum*



Virtual  $e^-e^+$  pair lifetime  $\Delta t \sim \hbar/mc^2 \approx 10^{-21} \text{ s}$

Field required to accelerate electron to  $mc^2 \approx eEc\Delta t$

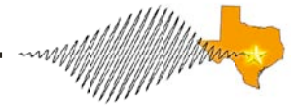
$$E_{crit} = m^2 c^3 / e\hbar = 1.6 \times 10^{16} \text{ V/cm}$$

Polarizing field

**Vacuum physics at extreme fields has never been explored  
Unanswered questions in QED/Standard Model:  
e.g. why don't vacuum fluctuations gravitate?**

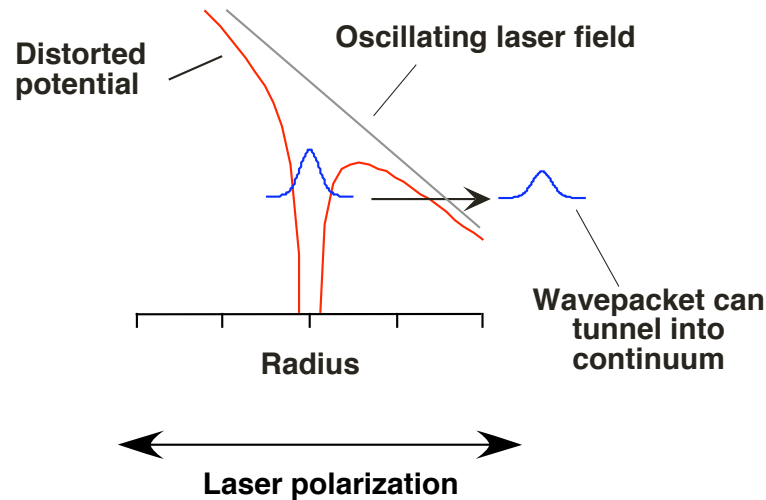


# At sufficient intensities it should be possible to observe the optical nonlinearity of vacuum



Atom/vacuum subject to an oscillating field  $E = E_0 \sin(\omega t - kz)$

## Atom



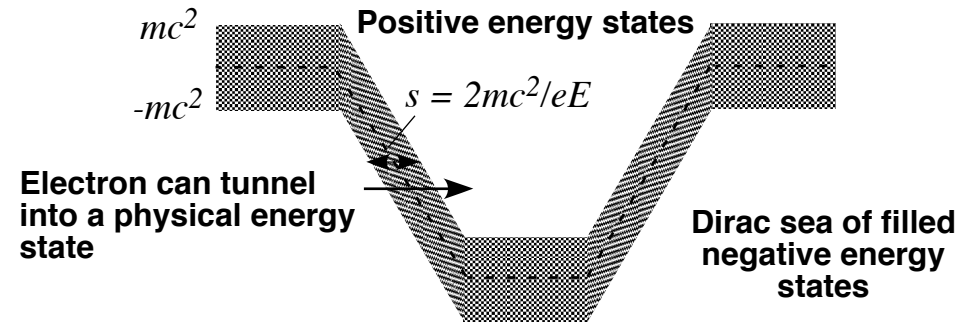
Tunnel ionization rate

$$W \sim \exp\left[-\frac{2}{3}\left(\frac{I_p}{I_H}\right)^{3/2} \frac{E_{at}}{E_0}\right]$$

$$E_{at} = 5 \times 10^9 \text{ V/cm} \quad \text{Intensity} \sim 3.5 \times 10^{16} \text{ W/cm}^2$$

Polarization of atoms gives rise to nonlinearities  
with  $E \sim 0.1\% E_{at}$   
 $I \sim 10^9 \text{ W/cm}^2$

## Vacuum



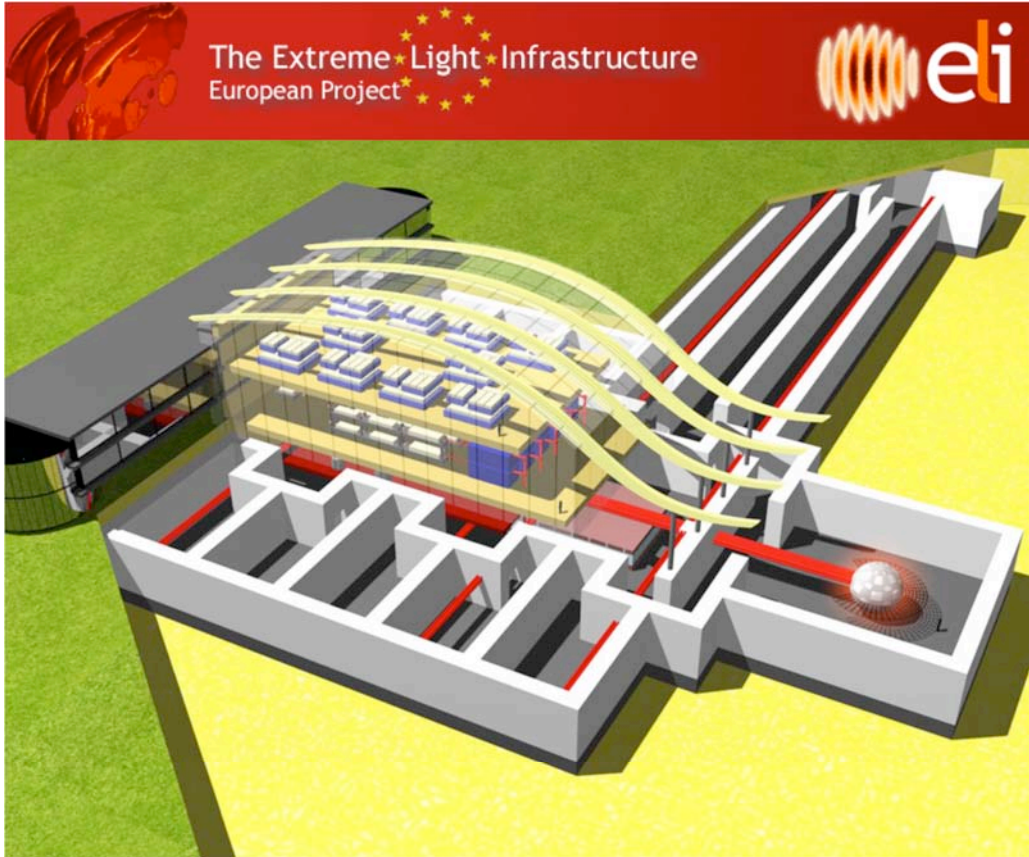
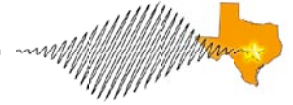
$$W \sim \exp\left[-\frac{s}{\lambda_c} \frac{\hbar}{mc}\right] = \exp\left[-2 \frac{E_s}{E_0}\right]$$

$$E_s = m^2 c^3 / e \hbar = 1.3 \times 10^{16} \text{ V/cm} \quad \text{Intensity} \sim 2.2 \times 10^{29} \text{ W/cm}^2$$

Polarization of vacuum gives rise to nonlinearities  
with  $E \sim 0.1\% E_s$   
 $I \sim 10^{22} \text{ W/cm}^2$

With a 10-100 PW laser it might be possible to observe birefringence of the vacuum

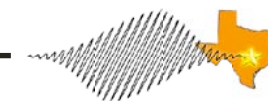
# The Europeans have initiated an EU funded project to build multiple 10 PW-class lasers



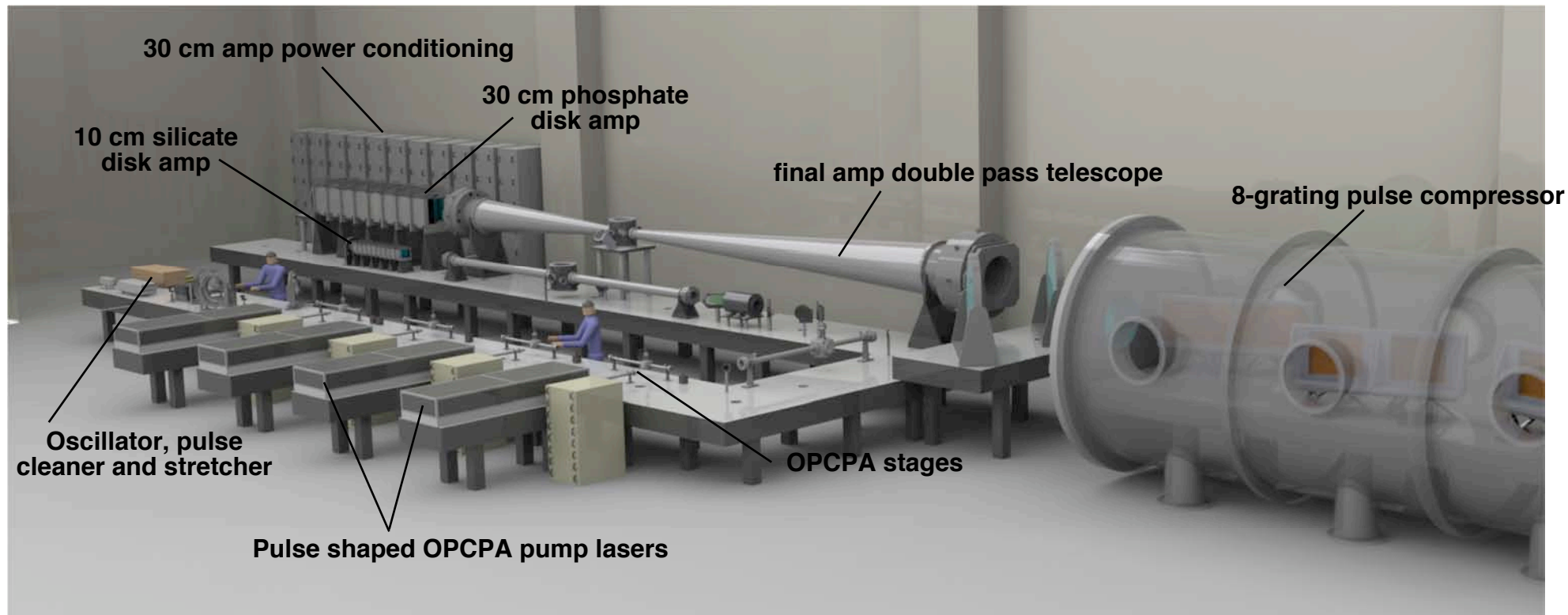
## Three ELI Pillars

- **Bucharest, Romania: ELI - NP**  
Devoted to nuclear physics with intense lasers and gamma beams
- **Prague, Czech Republic: ELI - CZ**  
Devoted to work on electron acceleration
- **Szeged, Hungary; ELI - AS**  
Devoted to attosecond pulse generation

# The hybrid mixed glass architecture would enable construction of a compact 10 PW laser



## Mechanical Engineering conception of the 10 PW Hybrid Mixed glass laser



### Laser output:

Energy: 1500 J,

Pulse duration: <150 fs

repetition rate: 1 shot/min

Laser Wavelength: 1054 nm

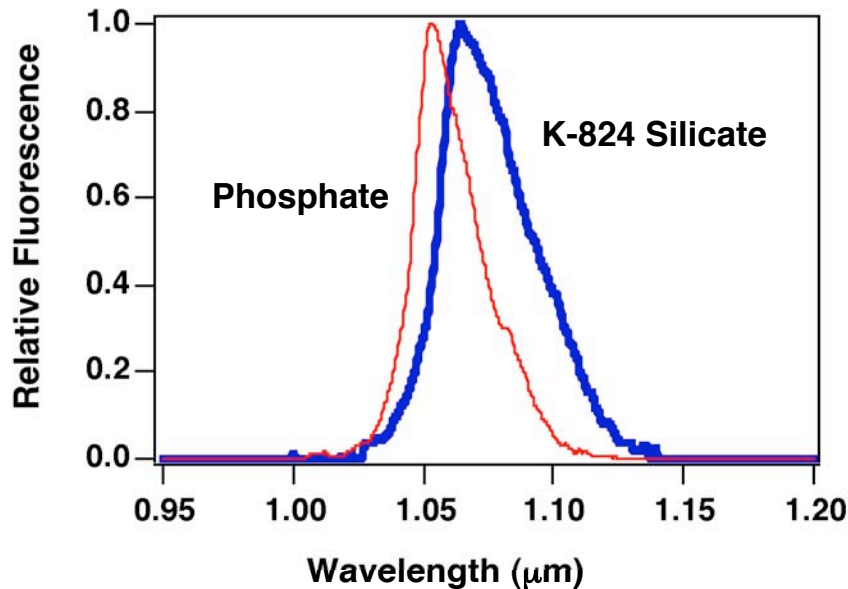
Temporal pulse contrast:  $10^{10}:1$  at  $> 10$  ps



# We have isolated two candidate glasses which have a broad gain spectrum and good gain properties



## K-824 Tantalate/Silicate glass



Peak Wavelength: 1065 nm

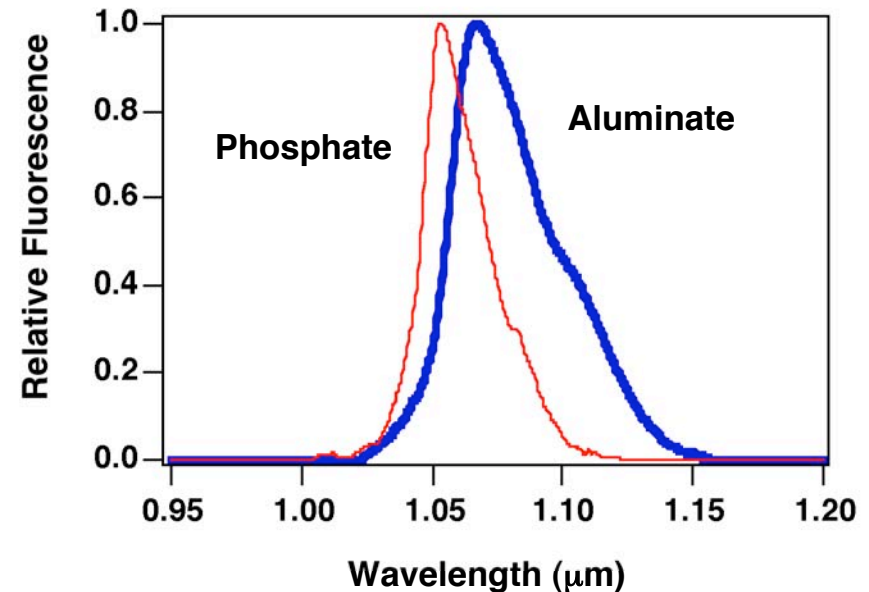
Peak cross section:  $2.4 \times 10^{-20} \text{ cm}^2$

Linewidth (FWHM): 38.2 nm

---

Ta <sub>2</sub> O <sub>5</sub>	57%
SiO <sub>2</sub>	17%
MgO	10%
BaO	15%
Nd <sub>2</sub> O <sub>3</sub>	~1%

## L-65 Aluminate glass



Peak Wavelength: 1067 nm

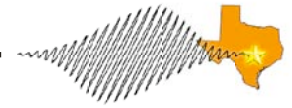
Peak cross section:  $1.8 \times 10^{-20} \text{ cm}^2$

Linewidth (FWHM): 41.2 nm

---

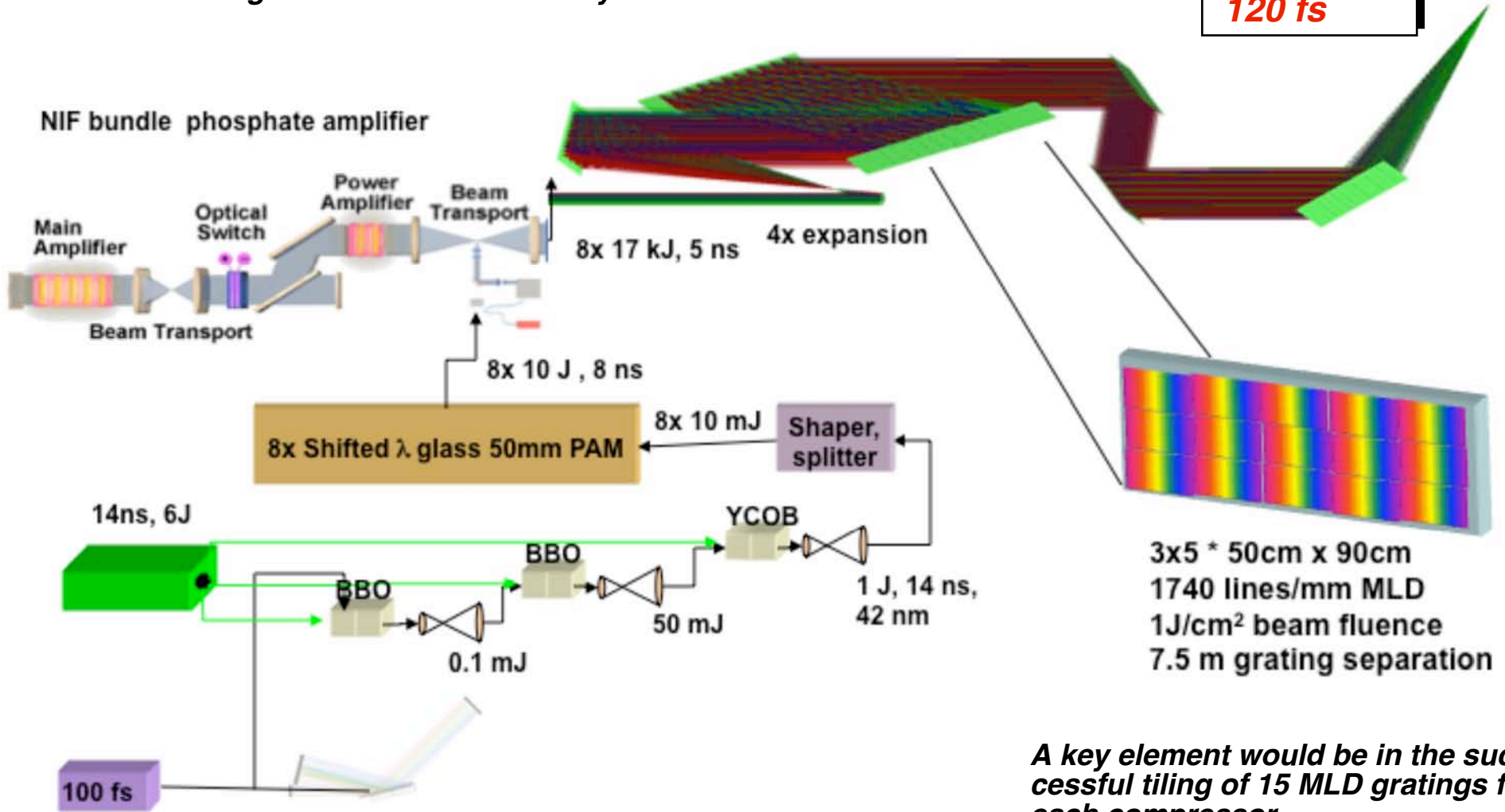
Al <sub>2</sub> O <sub>3</sub>	42%
CaO	38%
BaO	10%
SiO <sub>2</sub>	8%
Nd <sub>2</sub> O <sub>3</sub>	~4%

# The architecture of a mixed glass exawatt laser would be straightforward



*Final gain would be in 8 NIF-style beamlines*

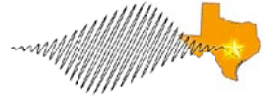
**8x 15 kJ =  
120 kJ in  
120 fs**



*A key element would be in the successful tiling of 15 MLD gratings for each compressor*

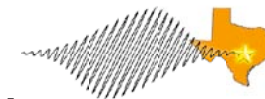
# The high energy amplifier architecture of a mixed-glass Exawatt laser would be based on NIF technology

---

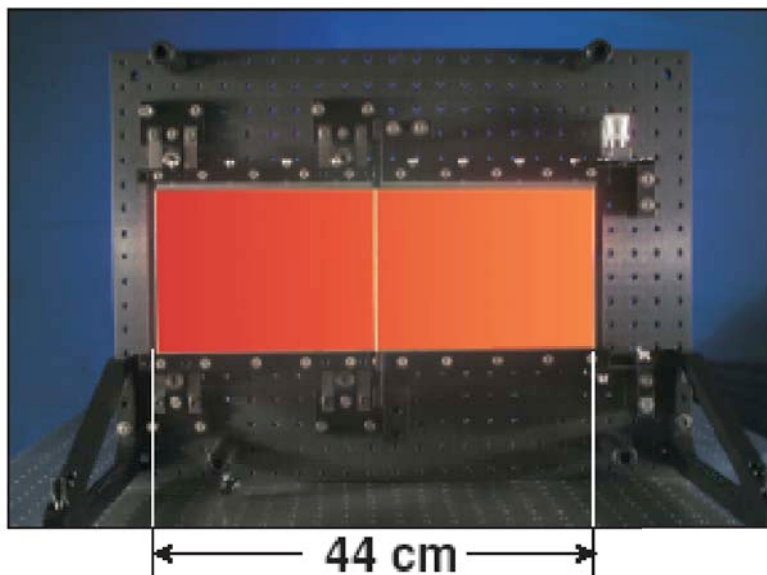




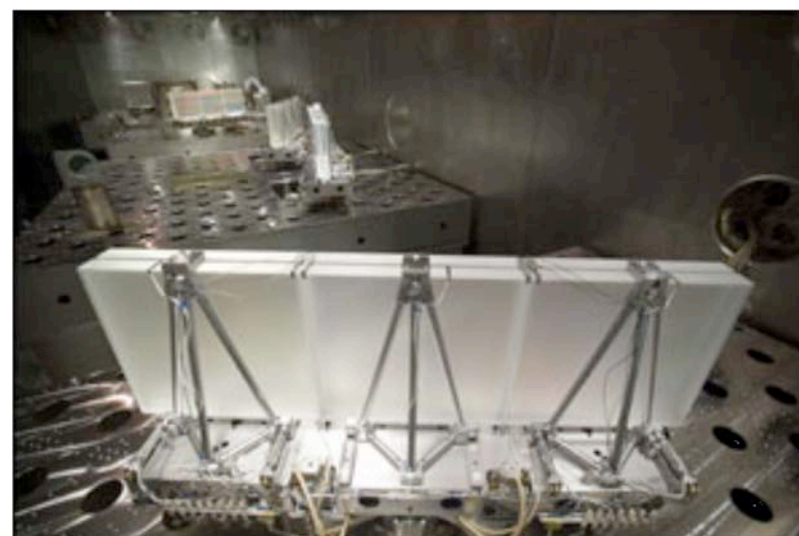
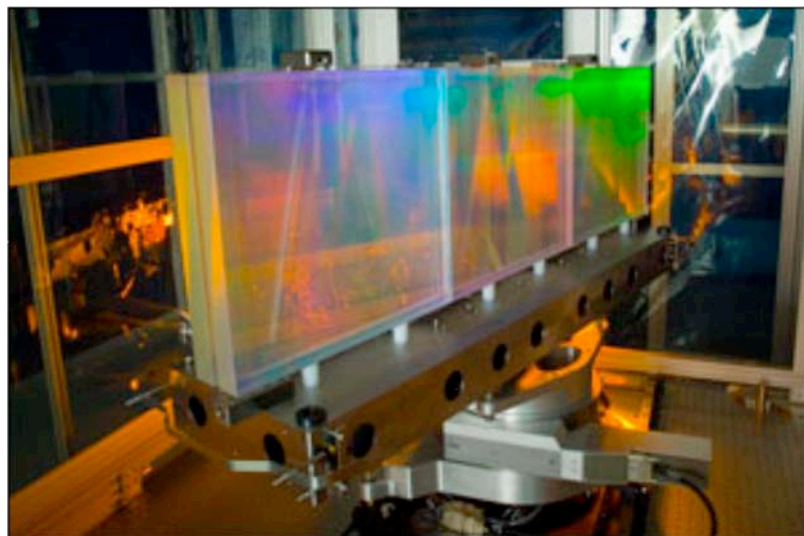
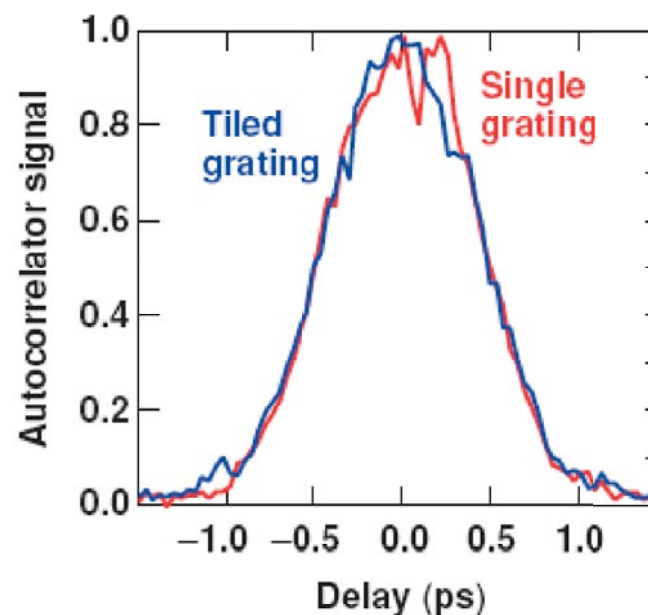
# The idea of tiling multiple gratings for compression of $1 \mu\text{m}$ pulses has been demonstrated at Omega EP



*Two-grating phased array at the U. of Rochester*

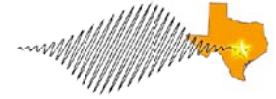


*Pulse compression data using the two grating array (U. of Rochester)*



# **We are presently at a cross roads in examining how to get to move beyond 1 PW powers**

---



- Can 10 PW be built in five years and*
- Is there a technology route to an exawatt in a decade?*
- Push to shorter pulses or higher pulse energy?*
- Utilize Ti:sapphire or some other gain medium?*
- Can OPCPA be employed all the way to an exawatt?*
- How can 10 PW to exawatt pulses be compressed?*
- What will it cost to build a 10 PW laser or an exawatt?*

***Using hybrid OPCPA/Mixed laser glass technology, ~100 fs PW lasers at  $E > 100J$  are possible and it is possible to build a 10 PW laser on this technology now***